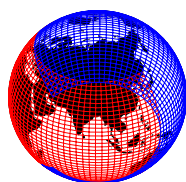
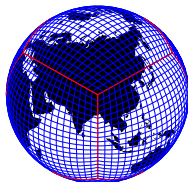


Atmosphere Modeling I: Intro & Dynamics

the CAM (Community Atmosphere Model) FV (Finite Volume) and SE (Spectral element) dynamical cores

Peter Hjort Lauritzen

Atmospheric Modeling & Predictability Section (AMP)
Climate and Global Dynamics Laboratory (CGD)
National Center for Atmospheric Research (NCAR)



NCAR is sponsored by the National Science Foundation (NSF)

1 Atmosphere intro

- Multi-scale nature of atmosphere dynamics
- Resolved and un-resolved scales
- 'Define' dynamical core and parameterizations

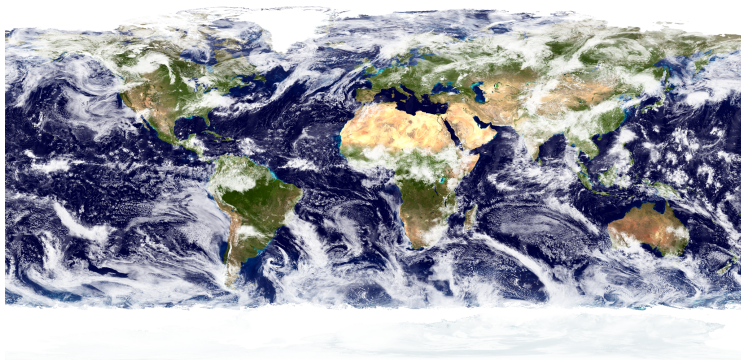
2 CAM-FV dynamical core (current 'work horse' dynamical core for $\approx 1^\circ$ applications)

- Horizontal and vertical grid
- Equations of motion
- The Lin and Rood (1996) advection scheme
- Finite-volume discretization of the equations of motion
- The 'CD' grid approach
- Vertical remapping
- Tracers

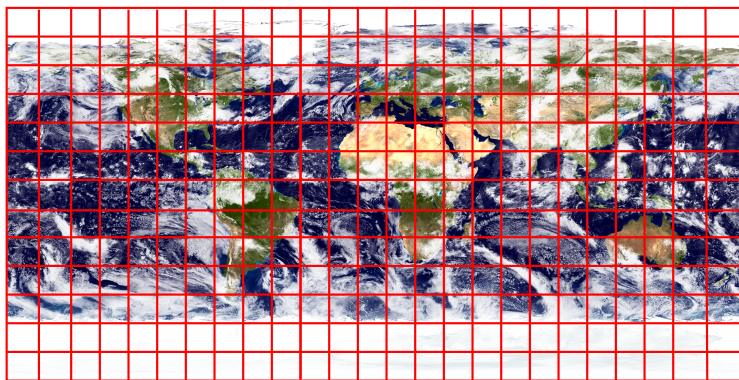
3 CESM simpler models (focus on CAM)

4 Other dynamical core options in CAM

- CAM-EUL (Eulerian): Based on spherical harmonic functions (only simple physics support)
- CAM-SE (Spectral-Elements): Default dynamical core in CAM for high (hydrostatic) horizontal resolution and mesh-refinement applications
- Dynamical cores being integrated into CAM: MPAS and FV3



Source: NASA Earth Observatory



- Red lines: regular latitude-longitude grid
- Grid-cell size defines the smallest scale that can be resolved (\neq **effective resolution!**)
- Many important processes taking place sub-grid-scale that must be parameterized
- Loosely speaking, the parameterizations compute grid-cell average tendencies due to sub-grid-scale processes in terms of the (resolved scale) atmospheric state
- In modeling jargon parameterizations are also referred to as *physics* (what is unphysical about resolved scale dynamics?)

Effective resolution: smallest scale (highest wave-number $k = k_{eff}$) that a model can accurately represent

- k_{eff} can be assessed analytically for linearized equations (Von Neumann analysis)
- In a full model one can assess k_{eff} using total kinetic energy spectra (TKE) of, e.g., horizontal wind \vec{v} (see Figure below)

Effective resolution is typically 4-10 grid-lengths depending on numerical method!

⇒ be careful analyzing phenomena at the grid scale (e.g., extremes)

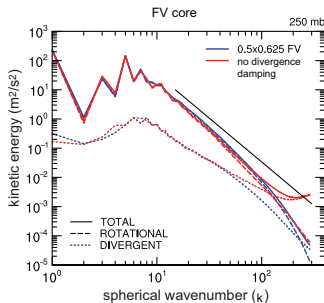
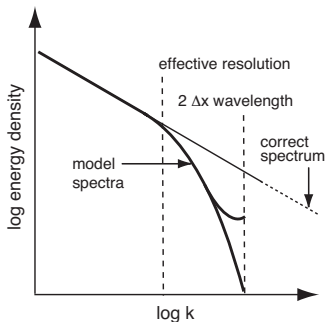
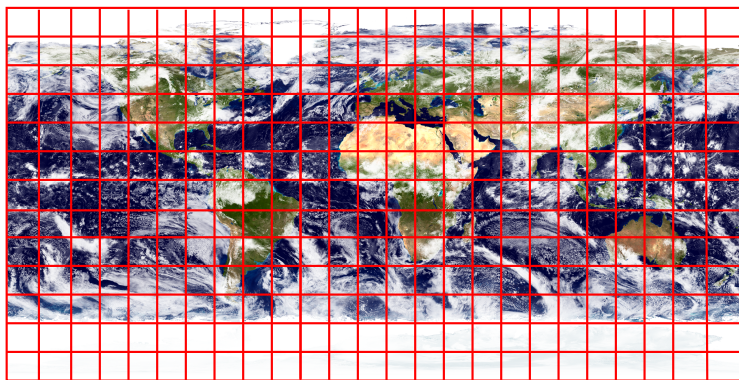


Figure from Skamarock (2011): (left) Schematic depicting the possible behavior of spectral tails derived from model forecasts. (right) TKE (solid lines) as a function of spherical wavenumber for the CCSM finite-volume dynamical core derived from aquaplanet simulations. The total KE is broken into divergent and rotational components (dashed lines) and the solid black lines shows the k^{-3} slope.



- Red lines: regular latitude-longitude grid
- Grid-cell size defines the smallest scale that can be resolved (\neq **effective resolution!**)
- Many important processes taking place sub-grid-scale that must be parameterized
- Loosely speaking, the parameterizations compute grid-cell average tendencies due to sub-grid-scale processes in terms of the (resolved scale) atmospheric state
- In modeling jargon parameterizations are also referred to as *physics* (what is unphysical about resolved scale dynamics?)

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)

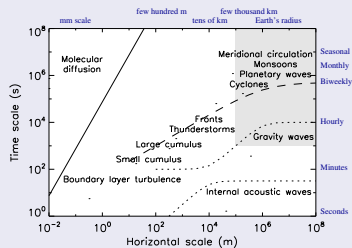


Figure indicates schematically the time scales and horizontal spatial scales of a range of atmospheric phenomena (Figure from Thuburn 2011).

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)

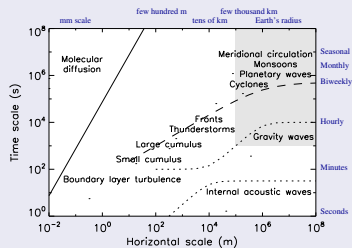


Figure indicates schematically the time scales and horizontal spatial scales of a range of atmospheric phenomena (Figure from Thuburn 2011).

- $\mathcal{O}(10^4 \text{ km})$: large scale circulations (Asian summer monsoon).

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)

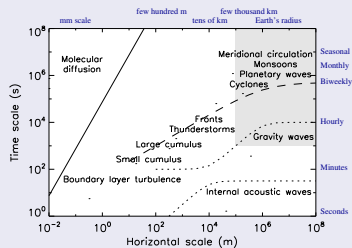


Figure indicates schematically the time scales and horizontal spatial scales of a range of atmospheric phenomena (Figure from Thuburn 2011).

- $\mathcal{O}(10^4 \text{ km})$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4 \text{ km})$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)

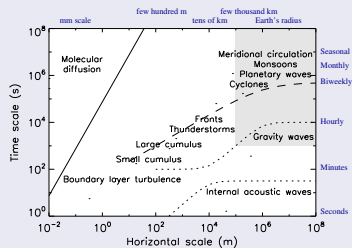


Figure indicates schematically the time scales and horizontal spatial scales of a range of atmospheric phenomena (Figure from Thuburn 2011).

- $\mathcal{O}(10^4 \text{ km})$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4 \text{ km})$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)
- $\mathcal{O}(10^3 \text{ km})$: cyclones and anticyclones

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)

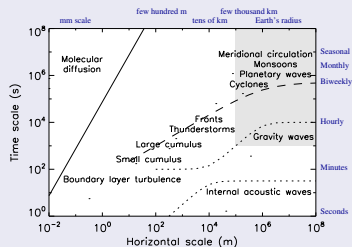


Figure indicates schematically the time scales and horizontal spatial scales of a range of atmospheric phenomena (Figure from Thuburn 2011).

- $\mathcal{O}(10^4 \text{ km})$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4 \text{ km})$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)
- $\mathcal{O}(10^3 \text{ km})$: cyclones and anticyclones
- $\mathcal{O}(10 \text{ km})$: the transition zones between relatively warm and cool air masses can collapse in scale to form fronts with widths of a few tens of km

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)

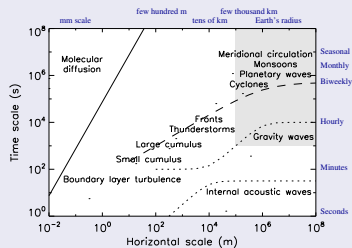


Figure indicates schematically the time scales and horizontal spatial scales of a range of atmospheric phenomena (Figure from Thuburn 2011).

- $\mathcal{O}(10^4 \text{ km})$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4 \text{ km})$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)
- $\mathcal{O}(10^3 \text{ km})$: cyclones and anticyclones
- $\mathcal{O}(10 \text{ km})$: the transition zones between relatively warm and cool air masses can collapse in scale to form fronts with widths of a few tens of km
- $\mathcal{O}(10^3 \text{ km} - 100 \text{ m})$: convection can be organized on a huge range of different scales (tropical intraseasonal oscillations; supercell complexes and squall lines; individual small cumulus clouds formed from turbulent boundary layer eddies)

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)

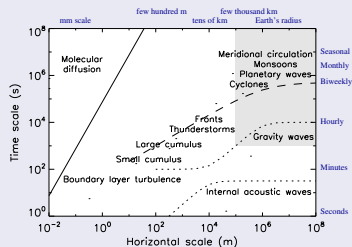
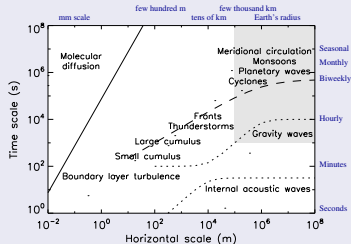


Figure indicates schematically the time scales and horizontal spatial scales of a range of atmospheric phenomena (Figure from Thuburn 2011).

- $\mathcal{O}(10^4 \text{ km})$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4 \text{ km})$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)
- $\mathcal{O}(10^3 \text{ km})$: cyclones and anticyclones
- $\mathcal{O}(10 \text{ km})$: the transition zones between relatively warm and cool air masses can collapse in scale to form fronts with widths of a few tens of km
- $\mathcal{O}(10^3 \text{ km} - 100 \text{ m})$: convection can be organized on a huge range of different scales (tropical intraseasonal oscillations; supercell complexes and squall lines; individual small cumulus clouds formed from turbulent boundary layer eddies)
- $\mathcal{O}(10 \text{ m} - 1 \text{ mm})$: turbulent eddies in boundary layer (lowest few hundred *m*'s of the atmosphere, where the dynamics is dominated by turbulent transports); range in scale from few hundred *m*'s (the boundary layer depth) down to *mm* scale at which molecular diffusion becomes significant.

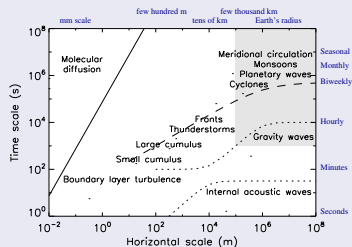
Multi-scale nature of atmosphere dynamics (from Thuburn 2011)



- All of the phenomena along the dashed line are important for weather and climate, and so need to be represented in numerical models.
- **Important phenomena occur at all scales - there is no significant spectral gap!** Moreover, there are strong interactions between the phenomena at different scales, and these interactions need to be represented.
- The lack of any spectral gap makes the modeling of weather/climate very **challenging**
- The emphasis in this lecture is how we model resolved dynamics; however, it should be borne in mind that equally important is how we represent unresolved processes, and the interactions between resolved and unresolved processes.

- $\mathcal{O}(10^4 km)$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4 km)$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)
- $\mathcal{O}(10^3 km)$: cyclones and anticyclones
- $\mathcal{O}(10 km)$: the transition zones between relatively warm and cool air masses can collapse in scale to form fronts with widths of a few tens of km
- $\mathcal{O}(10^3 km - 100 m)$: convection can be organized on a huge range of different scales (tropical intraseasonal oscillations; supercell complexes and squall lines; individual small cumulus clouds formed from turbulent boundary layer eddies)
- $\mathcal{O}(10 m - 1 mm)$: turbulent eddies in boundary layer (lowest few hundred m 's of the atmosphere, where the dynamics is dominated by turbulent transports); range in scale from few hundred m 's (the boundary layer depth) down to mm scale at which molecular diffusion becomes significant.

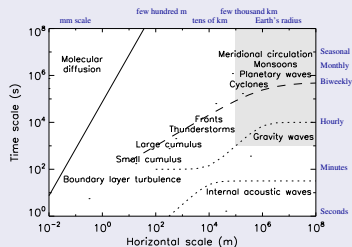
Multi-scale nature of atmosphere dynamics (from Thuburn 2011)



- Two dotted curves correspond to dispersion relations for internal inertio-gravity waves and internal acoustic waves (relatively fast processes)
- these lines lie significantly below the energetically dominant processes on the dashed line
 - \Rightarrow they are energetically weak compared to the dominant processes along the dashed curve
 - \Rightarrow we do relatively little damage if we distort their propagation (will return to this later)
 - the fact that these waves are fast puts constraints on the size of Δt (at least for explicit and semi-implicit time-stepping schemes)!

- $\mathcal{O}(10^4 km)$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4 km)$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)
- $\mathcal{O}(10^3 km)$: cyclones and anticyclones
- $\mathcal{O}(10 km)$: the transition zones between relatively warm and cool air masses can collapse in scale to form fronts with widths of a few tens of km
- $\mathcal{O}(10^3 km - 100 m)$: convection can be organized on a huge range of different scales (tropical intraseasonal oscillations; supercell complexes and squall lines; individual small cumulus clouds formed from turbulent boundary layer eddies)
- $\mathcal{O}(10 m - 1 mm)$: turbulent eddies in boundary layer (lowest few hundred m 's of the atmosphere, where the dynamics is dominated by turbulent transports); range in scale from few hundred m 's (the boundary layer depth) down to mm scale at which molecular diffusion becomes significant.

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)

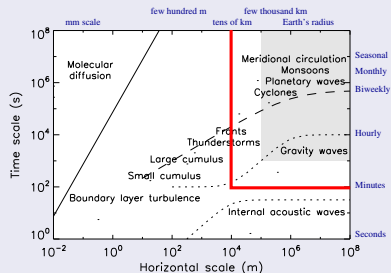


Horizontal resolution:

- the shaded region shows the resolved space/time scales in typical current day climate models (approximately $1^{\circ} - 2^{\circ}$ resolution)
- highest resolution at which uniform resolution CAM is run/developed is on the order of 10 – 25km
- as the resolution is increased some 'large-scale' parameterizations may no longer be necessary (e.g., large scale convection) and we might need to redesign some parameterizations that were developed for horizontal resolutions of hundreds of km's

- $\mathcal{O}(10^4 \text{ km})$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4 \text{ km})$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)
- $\mathcal{O}(10^3 \text{ km})$: cyclones and anticyclones
- $\mathcal{O}(10 \text{ km})$: the transition zones between relatively warm and cool air masses can collapse in scale to form fronts with widths of a few tens of km
- $\mathcal{O}(10^3 \text{ km} - 100 \text{ m})$: convection can be organized on a huge range of different scales (tropical intraseasonal oscillations; supercell complexes and squall lines; individual small cumulus clouds formed from turbulent boundary layer eddies)
- $\mathcal{O}(10 \text{ m} - 1 \text{ mm})$: turbulent eddies in boundary layer (lowest few hundred m's of the atmosphere, where the dynamics is dominated by turbulent transports); range in scale from few hundred m's (the boundary layer depth) down to mm scale at which molecular diffusion becomes significant.

Multi-scale nature of atmosphere dynamics (from Thuburn 2011)



Horizontal resolution:

- the shaded region shows the resolved space/time scales in typical current day climate models (approximately $1^{\circ} - 2^{\circ}$ resolution)
- highest resolution at which uniform resolution CAM is run/developed is on the order of $10 - 25\text{ km}$
- as the resolution is increased some 'large-scale' parameterizations may no longer be necessary (e.g., large scale convection) and we might need to redesign some parameterizations that were developed for horizontal resolutions of hundreds of km's

- $\mathcal{O}(10^4\text{ km})$: large scale circulations (Asian summer monsoon).
- $\mathcal{O}(10^4\text{ km})$: undulations in the jet stream and pressure patterns associated with the largest scale Rossby waves (called *planetary waves*)
- $\mathcal{O}(10^3\text{ km})$: cyclones and anticyclones
- $\mathcal{O}(10\text{ km})$: the transition zones between relatively warm and cool air masses can collapse in scale to form fronts with widths of a few tens of km
- $\mathcal{O}(10^3\text{ km} - 100\text{ m})$: convection can be organized on a huge range of different scales (tropical intraseasonal oscillations; supercell complexes and squall lines; individual small cumulus clouds formed from turbulent boundary layer eddies)
- $\mathcal{O}(10\text{ m} - 1\text{ mm})$: turbulent eddies in boundary layer (lowest few hundred m's of the atmosphere, where the dynamics is dominated by turbulent transports); range in scale from few hundred m's (the boundary layer depth) down to mm scale at which molecular diffusion becomes significant.

Parameterization suite

- Moist processes: deep convection, shallow convection, large-scale condensation
- Radiation and Clouds: cloud microphysics, precipitation processes, radiation
- Turbulent mixing: planetary boundary layer parameterization, vertical diffusion, gravity wave drag



'Resolved' dynamics

'Roughly speaking, the **dynamical core** solves the governing fluid and thermodynamic equations on resolved scales, while the parameterizations represent sub-grid-scale processes and other processes not included in the dynamical core such as radiative transfer.' - Thuburn (2008)

Parameterization suite

- Moist processes: deep convection, shallow convection, large-scale condensation
- Radiation and Clouds: cloud microphysics, precipitation processes, radiation
- Turbulent mixing: planetary boundary layer parameterization, vertical diffusion, gravity wave drag

Strategies for coupling:

- **process-split**: dynamical core & parameterization suite are based on the same state and their tendencies are added to produce the updated state (used in CAM-EUL)
- **time-split**: dynamic core & parameterization suite are calculated sequentially, each based on the state produced by the other (used in CAM-FV; **the order matters!**).
- different coupling approaches discussed in the context of CCM3 in Williamson (2002)
- simulations are very dependent on coupling time-step (e.g. Williamson and Olson, 2003)
- (re-)emerging research topic: physics-dynamics coupling (PDC) conference series (Gross et al., 2016)



'Resolved' dynamics

'Roughly speaking, the **dynamical core** solves the governing fluid and thermodynamic equations on resolved scales, while the parameterizations represent sub-grid-scale processes and other processes not included in the dynamical core such as radiative transfer.' - Thuburn (2008)

Spherical (horizontal) discretization grid

CAM-FV uses regular latitude-longitude grid:

- horizontal position: (λ, θ) , where λ longitude and θ latitude.
- horizontal resolution specified when creating a new case:

```
./create_newcase -res res ...
```

where, e.g., **res**=f09_f09_mg17 which is the $\Delta\lambda \times \Delta\theta = 0.9^\circ \times 1.25^\circ$ horizontal resolution configuration of the FV dynamical core corresponding to **nlon**=288, **nlat**=192.

Changing resolution requires rebuilding (not a namelist variable).

- Note: Convergence of the meridians near the poles.



Vertical coordinate: hybrid sigma ($\sigma = p/p_s$)-pressure (p) coordinate

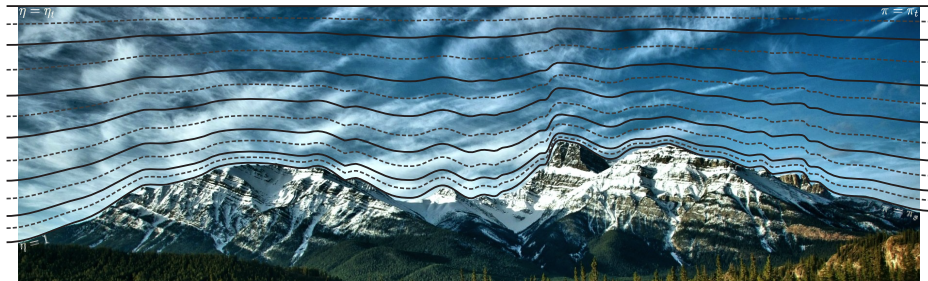


Figure courtesy of David Hall (CU Boulder).

Sigma layers at the bottom (following terrain) with isobaric (pressure) layers aloft.

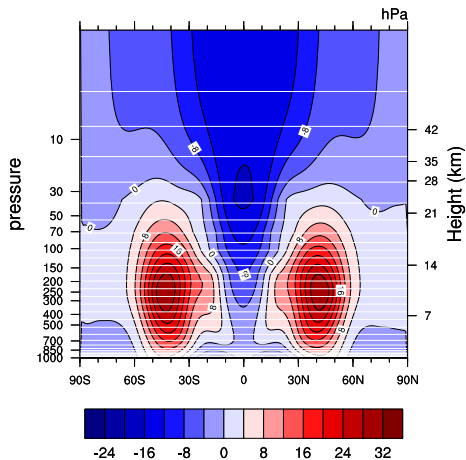
Pressure at model level interfaces

$$p_{k+1/2} = A_{k+1/2} p_0 + B_{k+1/2} p_s,$$

where p_s is surface pressure, p_0 is the model top pressure, and $A_{k+1/2} (\in [0 : 1])$ and $B_{k+1/2} (\in [1 : 0])$ hybrid coefficients (in model code: *hyai* and *hybi*). Similarly for model level mid-points.

Note: vertical index is 1 at model top and *klev* at surface.

Vertical coordinate: hybrid sigma ($\sigma = p/p_s$)-pressure (p) coordinate



Time & zonally averaged zonal wind (Held-Suarez forcing); overlaid CAM5 levels ($klev = 30$).

Why do we use terrain-following coordinates?

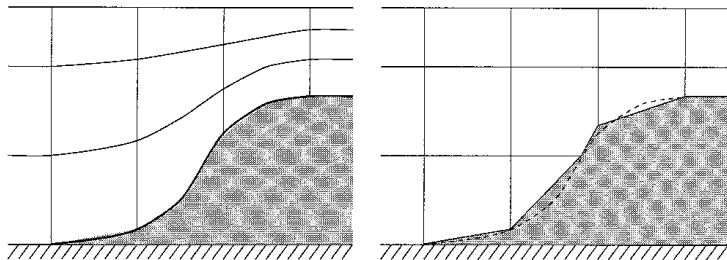


Figure: Representation of a smoothly varying bottom (dashed line) in (left) a terrain-following coordinate model, and (right) a height coordinate model with piecewise constant slopes (cut-cells, shaved-cells)

Figure is from Adcroft et al. (1997).

→ The main reason is that the lower boundary condition is very simple when using terrain-following coordinates!

Aside: hybrid sigma ($\sigma = p/p_s$)-pressure (p) coordinate

While terrain-following coordinates simplify the bottom boundary condition, they may introduce errors:

- Pressure gradient force (PDF) errors: $\frac{1}{\rho} \nabla p_z = \frac{1}{\rho} \nabla_{\eta} p + \frac{1}{\rho} \frac{dp}{dz} \nabla_{\eta} z$, (Kasahara, 1974) where ρ is density, p pressure and z height.
- Errors in modeling flow along constant z -surfaces near the surface

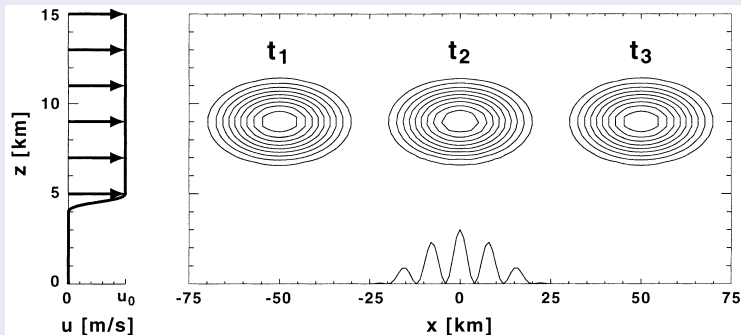
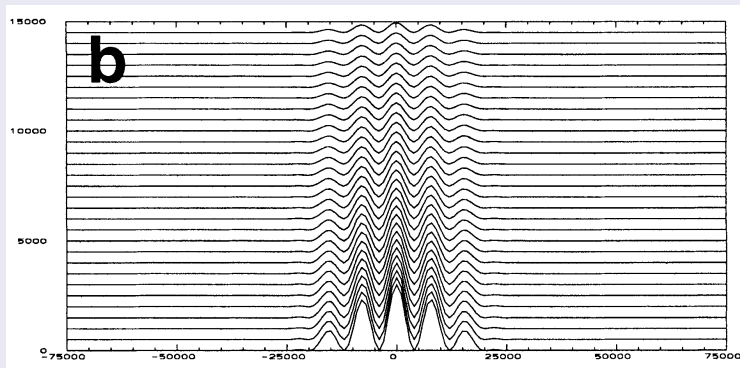


FIG. 4. Vertical cross section of the idealized two-dimensional advection test. The topography is located entirely within a stagnant pool of air, while there is a uniform horizontal velocity aloft. The analytical solution of the advected anomaly is shown at three instances.

Aside: hybrid sigma ($\sigma = p/p_s$)-pressure (p) coordinate

While terrain-following coordinates simplify the bottom boundary condition, they may introduce errors:

- Pressure gradient force (PDF) errors: $\frac{1}{\rho} \nabla p_z = \frac{1}{\rho} \nabla_{\eta} p + \frac{1}{\rho} \frac{dp}{dz} \nabla_{\eta} z$, (Kasahara, 1974) where ρ is density, p pressure and z height.
- Errors in modeling flow along constant z -surfaces near the surface

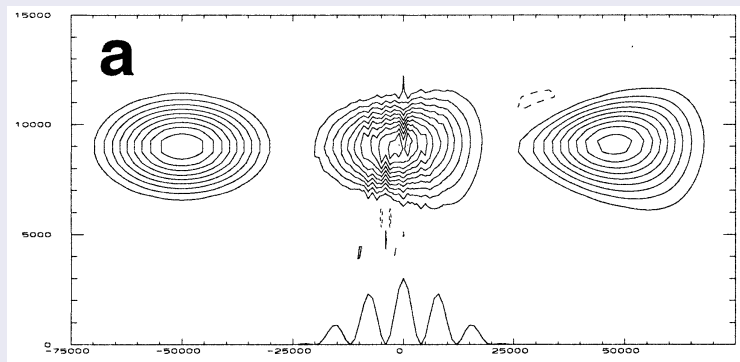


Schär et al. (2002)

Aside: hybrid sigma ($\sigma = p/p_s$)-pressure (p) coordinate

While terrain-following coordinates simplify the bottom boundary condition, they may introduce errors:

- Pressure gradient force (PDF) errors: $\frac{1}{\rho} \nabla p_z = \frac{1}{\rho} \nabla_{\eta} p + \frac{1}{\rho} \frac{dp}{dz} \nabla_{\eta} z$, (Kasahara, 1974) where ρ is density, p pressure and z height.
- Errors in modeling flow along constant z -surfaces near the surface



Schär et al. (2002)

- CAM-FV uses a Lagrangian ('floating') vertical coordinate ξ so that

$$\frac{d\xi}{dt} = 0,$$

i.e. vertical surfaces are material surfaces (no flow across them).

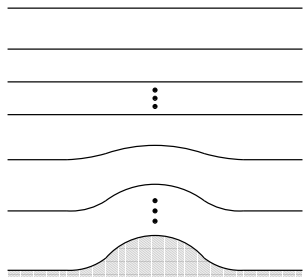


Figure shows 'usual' hybrid $\sigma - p$ vertical coordinate $\eta(p_s, p)$ (where p_s is surface pressure):

- $\eta(p_s, p)$ is a monotonic function of p .
- $\eta(p_s, p_s) = 1$
- $\eta(p_s, 0) = 0$
- $\eta(p_s, p_{top}) = \eta_{top}$.

Boundary conditions are:

- $\frac{d\eta(p_s, p_s)}{dt} = 0$
- $\frac{d\eta(p_s, p_{top})}{dt} = \omega(p_{top}) = 0$

(ω is vertical velocity in pressure coordinates)

- CAM-FV uses a Lagrangian ('floating') vertical coordinate ξ so that

$$\frac{d\xi}{dt} = 0,$$

i.e. vertical surfaces are material surfaces (no flow across them).

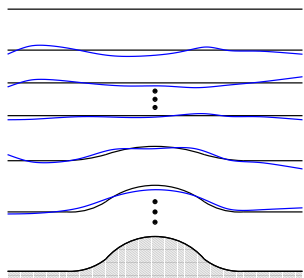


Figure:

- set $\xi = \eta$ at time t_{start} (black lines).
- for $t > t_{start}$ the vertical levels deform as they move with the flow (blue lines).
- to avoid excessive deformation of the vertical levels (non-uniform vertical resolution) the prognostic variables defined in the Lagrangian layers ξ are periodically remapped (= conservative interpolation) back to the Eulerian reference coordinates η (more on this later).

Why use floating Lagrangian vertical coordinates?

Vertical advection terms disappear (3D model becomes 'stacked shallow-water models'; only 2D numerical methods are needed)

- The vertical resolution is implicitly set during `./create_newcase` depending on (physics) configuration. For example, `klev=26` for CAM4, `klev=30` for CAM5 and `klev=32` for CAM6.
- The vertical resolution can be changed with

```
./xmlchange CAM_CONFIG_OPTS=-nlev 30
```

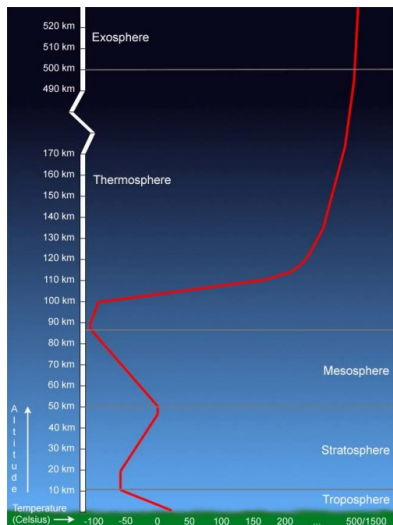
- If horizontal or vertical resolution is changed the user must point to an initial condition file matching that resolution. Non-default initial condition file is set in CAM namelist (`user_n1_cam` located in the `case` directory):

```
ncdata='inputdata/atm/cam/inic/fv/cami-mam3_0000-01-01_0.9x1.25_L30_c100618.nc'
```

- Changing vertical or horizontal resolution requires a 're-compile'.
- **WARNING!** CAM physics parameterizations are sensitive to resolution (especially vertical resolution) - usually a retuning of parameters is necessary to get an 'acceptable' climate.

The vertical extent is from the surface to

- approximately 40 km's / 2Pa for CAM
- approximately 100 km's / 10^{-6} hPa for WACCM (Whole Atmosphere Community Climate Model)
- approximately 500 km's / 10^{-9} hPa for WACCM-x



The following approximations are made to the compressible Euler equations:

- **spherical geoid:** geopotential Φ is only a function of radial distance from the center of the Earth r : $\Phi = \Phi(r)$ (for planet Earth the true gravitational acceleration is much stronger than the centrifugal force).
⇒ Effective gravity acts only in radial direction

The following approximations are made to the compressible Euler equations:

- **spherical geoid:** geopotential Φ is only a function of radial distance from the center of the Earth r : $\Phi = \Phi(r)$ (for planet Earth the true gravitational acceleration is much stronger than the centrifugal force).
⇒ Effective gravity acts only in radial direction
- **quasi-hydrostatic approximation** (also simply referred to as *hydrostatic approximation*): Involves ignoring the acceleration term in the vertical component of the momentum equations so that it reads:

$$\rho g = -\frac{\partial p}{\partial z}, \quad (1)$$

where g gravity, ρ density and p pressure. Good approximation down to horizontal scales greater than approximately $10km$.

The following approximations are made to the compressible Euler equations:

- **spherical geoid**: geopotential Φ is only a function of radial distance from the center of the Earth r : $\Phi = \Phi(r)$ (for planet Earth the true gravitational acceleration is much stronger than the centrifugal force).
⇒ Effective gravity acts only in radial direction
- **quasi-hydrostatic approximation** (also simply referred to as *hydrostatic approximation*): Involves ignoring the acceleration term in the vertical component of the momentum equations so that it reads:

$$\rho g = -\frac{\partial p}{\partial z}, \quad (1)$$

where g gravity, ρ density and p pressure. Good approximation down to horizontal scales greater than approximately 10km .

- **shallow atmosphere**: a collection of approximations. Coriolis terms involving the horizontal components of Ω are neglected (Ω is angular velocity), factors $1/r$ are replaced with $1/a$ where a is the mean radius of the Earth and certain other metric terms are neglected so that the system retains conservation laws for energy and angular momentum.

Adiabatic frictionless equations of motion using Lagrangian vertical coordinates

Assuming a Lagrangian vertical coordinate the hydrostatic equations of motion integrated over a layer can be written as

$$\begin{aligned} \text{mass air:} & \quad \frac{\partial(\delta p)}{\partial t} = -\nabla_h \cdot (\vec{v}_h \delta p), \\ \text{mass tracers:} & \quad \frac{\partial(\delta p q)}{\partial t} = -\nabla_h \cdot (\vec{v}_h q \delta p), \\ \text{horizontal momentum:} & \quad \frac{\partial \vec{v}_h}{\partial t} = -(\zeta + f) \vec{k} \times \vec{v}_h - \nabla_h \kappa - \nabla_p \Phi, \\ \text{thermodynamic:} & \quad \frac{\partial(\delta p \Theta)}{\partial t} = -\nabla_h \cdot (\vec{v}_h \delta p \Theta) \end{aligned}$$

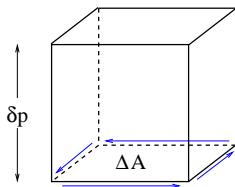
where δp is the layer thickness, \vec{v}_h is horizontal wind, q tracer mixing ratio, ζ vorticity, f Coriolis, κ kinetic energy, Θ potential temperature. The momentum equations are written in vector invariant form.

Adiabatic frictionless equations of motion using Lagrangian vertical coordinates

Assuming a Lagrangian vertical coordinate the hydrostatic equations of motion integrated over a layer can be written as

$$\begin{aligned} \text{mass air:} & \quad \frac{\partial(\delta p)}{\partial t} = -\nabla_h \cdot (\vec{v}_h \delta p), \\ \text{mass tracers:} & \quad \frac{\partial(\delta p q)}{\partial t} = -\nabla_h \cdot (\vec{v}_h q \delta p), \\ \text{horizontal momentum:} & \quad \frac{\partial \vec{v}_h}{\partial t} = -(\zeta + f) \vec{k} \times \vec{v}_h - \nabla_h \kappa - \nabla_p \Phi, \\ \text{thermodynamic:} & \quad \frac{\partial(\delta p \Theta)}{\partial t} = -\nabla_h \cdot (\vec{v}_h \delta p \Theta) \end{aligned}$$

The equations of motion are discretized using an Eulerian finite-volume approach.



Integrate the flux-form continuity equation horizontally over a control volume:

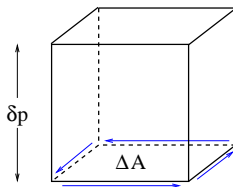
$$\frac{\partial}{\partial t} \iint_A \delta p \, dA = - \iint_A \nabla_h (\vec{v}_h \delta p) \, dA, \quad (2)$$

where A is the horizontal extent of the control volume. Using Gauss's divergence theorem for the right-hand side of (2) we get:

$$\frac{\partial}{\partial t} \iint_A \delta p \, dA = - \oint_{\partial A} \delta p \vec{v} \cdot \vec{n} \, dA, \quad (3)$$

where ∂A is the boundary of A and \vec{n} is outward pointing normal unit vector of ∂A .

Finite-volume discretization of continuity equation



Integrate the flux-form continuity equation horizontally over a control volume:

$$\frac{\partial}{\partial t} \iint_A \delta p \, dA = - \iint_A \nabla_h (\vec{v}_h \delta p) \, dA, \quad (2)$$

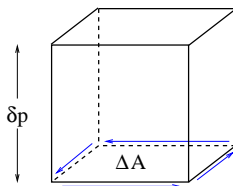
where A is the horizontal extent of the control volume. Using Gauss's divergence theorem for the right-hand side of (2) we get:

$$\frac{\partial}{\partial t} \iint_A \delta p \, dA = - \oint_{\partial A} \delta p \vec{v} \cdot \vec{n} \, dA, \quad (3)$$

Right-hand side of (3) represents the instantaneous flux of mass through the vertical faces of the control volume.

Next: integrate over one time-step Δt_{dyn} and discretize left-hand side.

Finite-volume discretization of continuity equation



Integrate the flux-form continuity equation horizontally over a control volume:

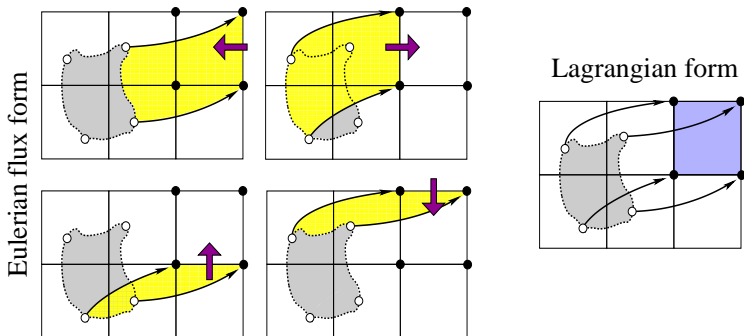
$$\frac{\partial}{\partial t} \iint_A \delta p \, dA = - \iint_A \nabla_h (\vec{v}_h \delta p) \, dA, \quad (2)$$

$$\Delta A \overline{\delta p}^{n+1} - \Delta A \overline{\delta p}^n = -\Delta t_{dyn} \int_{t=n\Delta t}^{t=(n+1)\Delta t} \left[\oint_{\partial A} \delta p \vec{v} \cdot \vec{n} \, dA \right] dt, \quad (3)$$

where n is time-level index and $\overline{(\cdot)}$ is cell-averaged value.

The right-hand side represents the mass transported through all of the four vertical control volume faces into the cell during one time-step. Graphical illustration on next slide!

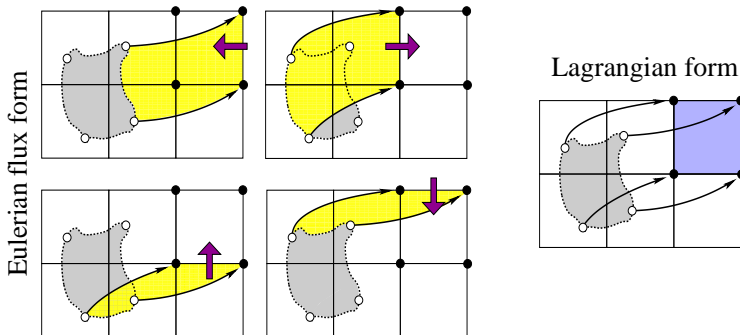
Finite-volume discretization of continuity equation: Tracking mass



The yellow areas are 'swept' through the control volume faces during one time-step. The grey area is the corresponding Lagrangian area (area moving with the flow with no flow through its boundaries that ends up at the Eulerian control volume after one time-step). Black arrows show parcel trajectories.

Note **equivalence** between Eulerian flux-form and Lagrangian form!

(Lauritzen et al., 2011b)



Until now everything has been exact. How do we approximate the fluxes numerically?

- In CAM-FV the Lin and Rood (1996) scheme is used which is a dimensionally split scheme (that is, rather than 'explicitly' estimating the boundaries of the yellow areas and integrate over them, fluxes are estimated by successive applications of one-dimensional operators in each coordinate direction).

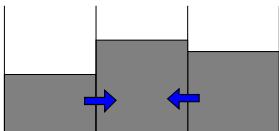
$$\overline{\delta p}^{n+1} = \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right],$$

where

$F^{\lambda,\theta}$ = flux divergence in λ or θ coordinate direction

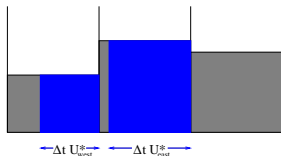
$f^{\lambda,\theta}$ = advective update in λ or θ coordinate direction

$$\overline{\delta p}^{n+1} = \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right],$$



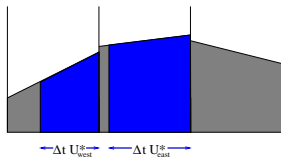
- Figure: Graphical illustration of flux-divergence operator F^λ . Shaded areas show cell average values for the cell we wish to make a forecast for and the two adjacent cells.

$$\overline{\delta p}^{n+1} = \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta (\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda (\overline{\delta p}^n) \right) \right],$$



- $u_{East/West}^*$ are the time-averaged winds on each face (more on how these are obtained later).
- F^λ is proportional to the difference between mass 'swept' through East and West cell face.
- $f^\lambda = F^\lambda + \overline{\overline{\delta p}} \Delta t_{dyn} D$, where D is divergence.
- On Figure we assume constant sub-grid-cell reconstructions for the fluxes.

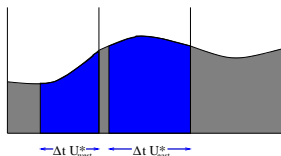
$$\overline{\delta p}^{n+1} = \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right],$$



Higher-order approximation to the fluxes:

- Piecewise linear sub-grid-scale reconstruction (van Leer, 1977): Fit a linear function using neighboring grid-cell average values with mass-conservation as a constraint (i.e. area under linear function = cell average).

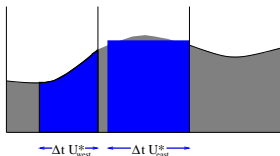
$$\overline{\delta p}^{n+1} = \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right],$$



Higher-order approximation to the fluxes:

- Piecewise linear sub-grid-scale reconstruction (van Leer, 1977): Fit a linear function using neighboring grid-cell average values with mass-conservation as a constraint (i.e. area under linear function = cell average).
- Piecewise parabolic sub-grid-scale reconstruction (Colella and Woodward, 1984): Fit parabola using neighboring grid-cell average values with mass-conservation as a constraint. Note: Reconstruction is C^0 across cell edges.

$$\overline{\delta p}^{n+1} = \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right],$$



Higher-order approximation to the fluxes:

- Piecewise linear sub-grid-scale reconstruction (van Leer, 1977): fit a linear function using neighboring grid-cell average values with mass-conservation as a constraint (i.e. area under linear function = cell average).
- Piecewise parabolic sub-grid-scale reconstruction (Colella and Woodward, 1984): fit parabola using neighboring grid-cell average values with mass-conservation as a constraint. Note: reconstruction is continuous at cell edges.
- Reconstruction function may 'overshoot' or 'undershoot' which may lead to unphysical and/or oscillatory solutions. Use limiters to render reconstruction function shape-preserving.

$$\overline{\delta p}^{n+1} = \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right],$$

Advantages:

- Inherently mass conservative (note: conservation does not necessarily imply accuracy!).
- Formulated in terms of one-dimensional operators.
- Preserves constant mass field for a non-divergent flow field (if the finite-difference approximation to divergence is zero).
- Preserves linear correlations between trace species (if shape-preservation filters are not applied)
- Has shape-preserving options.
- CAM-FV uses the PPM (Piecewise Parabolic Method; Colella and Woodward, 1984) with shape-preserving filters described in Lin and Rood (1996)

- In top layers operators are reduced to first order:

if ($k \leq k_{lev}/8$) IORD=JORD=1

E.g., for $k_{lev}=30$ the operators are altered in the top 3 layers.

- The advective $f^{\lambda,\theta}$ (*inner*) operators are 'hard-coded' to 1st order. For a linear analysis of the consequences of using *inner* and *outer* operators of different orders see Lauritzen (2007).

Hydrostatic equations of motion integrated over a Lagrangian layer

$$\begin{aligned}\frac{\partial(\delta p)}{\partial t} &= -\nabla_h \cdot (\vec{v}_h \delta p), \\ \frac{\partial(\delta p q)}{\partial t} &= -\nabla_h \cdot (\vec{v}_h \delta p), \\ \frac{\partial \vec{v}_h}{\partial t} &= -(\zeta + f) \vec{k} \times \vec{v}_h - \nabla_h \kappa - \nabla_p \Phi, \\ \frac{\partial(\delta p \Theta)}{\partial t} &= -\nabla_h \cdot (\vec{v}_h \delta p \Theta)\end{aligned}$$

The equations of motion are discretized using an Eulerian finite-volume approach.

Hydrostatic equations of motion integrated over a Lagrangian layer

$$\begin{aligned}\bar{\delta p}^{n+1} &= \bar{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\bar{\delta p}^n + f^\theta(\bar{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\bar{\delta p}^n + f^\lambda(\bar{\delta p}^n) \right) \right], \\ \frac{\partial(\delta p q)}{\partial t} &= -\nabla_h \cdot (\vec{v}_h \delta p), \\ \frac{\partial \vec{v}_h}{\partial t} &= -(\zeta + f) \vec{k} \times \vec{v}_h - \nabla_h \kappa - \nabla_p \Phi, \\ \frac{\partial(\delta p \Theta)}{\partial t} &= -\nabla_h \cdot (\vec{v}_h \delta p \Theta)\end{aligned}$$

Hydrostatic equations of motion integrated over a Lagrangian layer

$$\begin{aligned}\overline{\delta p}^{n+1} &= \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right], \\ \overline{\delta p q}^{n+1} &= \text{super-cycled (discussed later),} \\ \frac{\partial \vec{v}_h}{\partial t} &= -(\zeta + f) \vec{k} \times \vec{v}_h - \nabla_h \kappa - \nabla_p \Phi, \\ \frac{\partial(\delta p \Theta)}{\partial t} &= -\nabla_h \cdot (\vec{v}_h \delta p \Theta)\end{aligned}$$

Hydrostatic equations of motion integrated over a Lagrangian layer

$$\begin{aligned}\overline{\delta p}^{n+1} &= \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right], \\ \overline{\delta p q}^{n+1} &= \text{super-cycled (discussed later)}, \\ \vec{v}_h^{n+1} &= \vec{v}_h^n - \vec{\Gamma}^1 \left[(\zeta + f) \vec{k} \times \vec{v}_h \right] - \nabla_h \left(\vec{\Gamma}^2 \kappa \right) - \Delta t_{dyn} \hat{P}, \\ \frac{\partial(\delta p \Theta)}{\partial t} &= -\nabla_h \cdot (\vec{v}_h \delta p \Theta)\end{aligned}$$

- $\vec{\Gamma}^1$ is operator using combinations of $F^{\lambda, \theta}$ and $f^{\lambda, \theta}$ as components to approximate the time-volume-average of the vertical component of absolute vorticity. Similarly for $\vec{\Gamma}^2$ but for kinetic energy. ∇_h is simply approximated by finite differences. For details see Lin (2004).
- \hat{P} is a finite-volume discretization of the pressure gradient force (see Lin 1997 for details).

Hydrostatic equations of motion integrated over a Lagrangian layer

$$\begin{aligned}\overline{\delta p}^{n+1} &= \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right], \\ \overline{\delta p q}^{n+1} &= \text{super-cycled (discussed later)}, \\ \vec{v}_h^{n+1} &= \vec{v}_h^n - \vec{\Gamma}^1 \left[(\zeta + f) \vec{k} \times \vec{v}_h \right] - \nabla_h \left(\vec{\Gamma}^2 \kappa \right) - \Delta t_{dyn} \hat{P}, \\ \overline{\Theta \delta p}^{n+1} &= \overline{\Theta \delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\Theta \delta p}^n + f^\theta(\overline{\Theta \delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\Theta \delta p}^n + f^\lambda(\overline{\Theta \delta p}^n) \right) \right],\end{aligned}$$

Hydrostatic equations of motion integrated over a Lagrangian layer

$$\begin{aligned}\overline{\delta p}^{n+1} &= \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right], \\ \overline{\delta p q}^{n+1} &= \text{super-cycled (discussed later)}, \\ \vec{v}_h^{n+1} &= \vec{v}_h^n - \vec{\Gamma}^1 \left[(\zeta + f) \vec{k} \times \vec{v}_h \right] - \nabla_h \left(\vec{\Gamma}^2 \kappa \right) - \Delta t_{dyn} \hat{P}, \\ \overline{\Theta \delta p}^{n+1} &= \overline{\Theta \delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\Theta \delta p}^n + f^\theta(\overline{\Theta \delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\Theta \delta p}^n + f^\lambda(\overline{\Theta \delta p}^n) \right) \right],\end{aligned}$$

- No explicit diffusion operators in equations (so far!).
- Implicit diffusion through shape-preservation constraints in F and f operators.
- CAM-FV has 'control' over vorticity at the grid scale through implicit diffusion in the operators F and f but it does not have explicit control over divergence near the grid scale.

Hydrostatic equations of motion integrated over a Lagrangian layer

$$\begin{aligned}\overline{\delta p}^{n+1} &= \overline{\delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\theta(\overline{\delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\delta p}^n + f^\lambda(\overline{\delta p}^n) \right) \right], \\ \overline{\delta pq}^{n+1} &= \text{super-cycled (discussed later)}, \\ \vec{v}_h^{n+1} &= \vec{v}_h^n - \vec{\Gamma}^1 \left[(\zeta + f) \vec{k} \times \vec{v}_h \right] - \nabla_h \left(\vec{\Gamma}^2 \kappa \right) - \Delta t_{dyn} \hat{P} + \Delta t_{dyn} \nabla_h \left(\nu \nabla_h^\ell D \right), \ell = 0, 2 \\ \overline{\Theta \delta p}^{n+1} &= \overline{\Theta \delta p}^n + F^\lambda \left[\frac{1}{2} \left(\overline{\Theta \delta p}^n + f^\theta(\overline{\Theta \delta p}^n) \right) \right] + F^\theta \left[\frac{1}{2} \left(\overline{\Theta \delta p}^n + f^\lambda(\overline{\Theta \delta p}^n) \right) \right],\end{aligned}$$

- No explicit diffusion operators in equations.
- Implicit diffusion through shape-preservation constraints in F and f operators.
- The above discretization leads to ‘control’ over vorticity at the grid scale through implicit diffusion but no explicit control over divergence.
- **Add divergence damping (2^{nd} -order or 4^{th} -order) term to momentum equations.**
Optionally a ‘Laplacian-like’ damping of wind components is used in upper 3 levels to slow down excessive polar night jet that appears at high horizontal resolutions.
namelist variable: `div24de12flag`

More details: Lauritzen et al. (2011a); for a stability analysis of divergence damping in CAM see Whitehead et al. (2011)

Total kinetic energy spectra

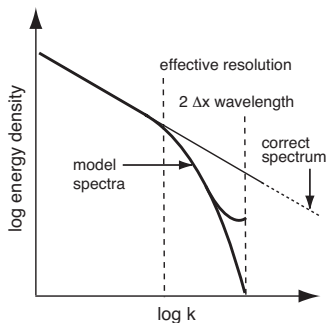
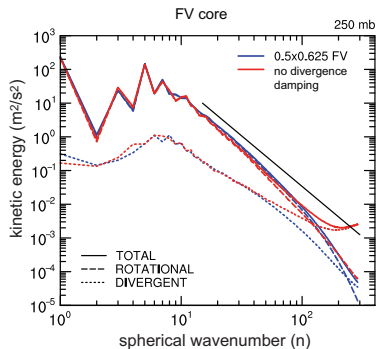


Figure: (left) Solid black line shows k^{-3} slope (courtesy of D.L. Williamson). (right) Schematic of 'effective resolution' (Figure from Skamarock (2011)).

- (left) Without divergence damping there is a spurious accumulation of total kinetic energy associated with divergent modes near the grid scale.
- (right) Note: total kinetic energy spectra can also be used to assess 'effective resolution' (see, e.g., discussion in Skamarock, 2011)

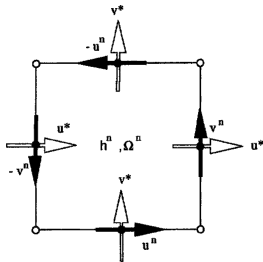


Figure from Lin and Rood (1997).

Definition of Arakawa C and D horizontal staggering (Arakawa and Lamb, 1977):

- C: velocity components at the center of cell faces and orthogonal to cell faces and mass variables at the cell center. Natural choice for mass-flux computations when using Lin and Rood (1996) scheme.
- D: velocity components parallel to cell faces and mass variables at the cell center. Natural choice for computing the circulation of vorticity ($\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$).

Time-stepping: the 'CD'- grid approach

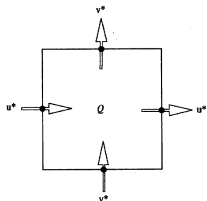


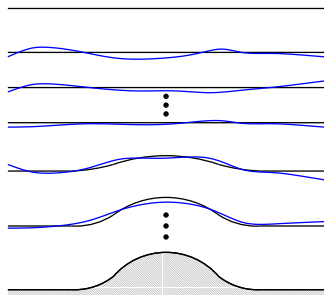
Figure from Lin and Rood (1997).

- For the flux- and advection operators (F and f , respectively) in the Lin and Rood (1996) scheme the time-centered advective winds (u^* , v^*) for the cell faces are needed:
- An option: extrapolate winds (as in semi-Lagrangian models) \Rightarrow can result in noise near steep topography (Lin and Rood, 1997).

- Instead, the equations of motion are integrated forward in time for $\frac{1}{2}\Delta t_{dyn}$ using a C grid horizontal staggering.
- These C -grid winds (u^* , v^*) are then used for the 'full' time-step update (everything else from the C -grid forecast is 'thrown away').
- The 'full' time-step update is performed on a D -grid.
- For a linear stability analysis of the 'CD'-grid approach see Skamarock (2008).

Vertical remapping

- CAM-FV uses a Lagrangian ('floating') vertical coordinate ξ .
- ξ is retained *ksplit* dynamics time-steps Δt_{dyn} .
- Hereafter the prognostic variables are remapped to the Eulerian vertical grid η .
- For horizontal resolution of 1° CAM *ksplit* = 4 and $\Delta t_{dyn} = 225s$
⇒ vertical remapping time-step is 900s
- Δt_{dyn} is chosen based on stability
(limited by gravity wave speed in CAM; advective winds in WACCM)
- Meridians are converging towards the poles: to stabilize the model (and reduce noise) FFT filters are applied along latitudes North/South of approximately $36^\circ N/S$.



- Continuity equation for air is coupled with momentum and thermodynamic equations:

- Continuity equation for air is coupled with momentum and thermodynamic equations:
 - thermodynamic variables and other prognostic variables feed back on the velocity field

- Continuity equation for air is coupled with momentum and thermodynamic equations:
 - thermodynamic variables and other prognostic variables feed back on the velocity field
 - which, in turn, feeds back on the solution to the continuity equation.

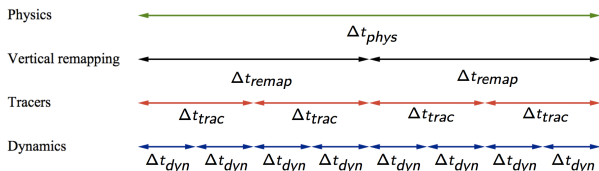
- Continuity equation for air is coupled with momentum and thermodynamic equations:
 - thermodynamic variables and other prognostic variables feed back on the velocity field
 - which, in turn, feeds back on the solution to the continuity equation.
 - Hence the continuity equation for air can not be solved in isolation and one must obey the maximum allowable time-step restrictions imposed by the fastest waves in the system.

Super-cycling (also referred to as sub-cycling) of tracers

- Continuity equation for air is coupled with momentum and thermodynamic equations:
 - thermodynamic variables and other prognostic variables feed back on the velocity field
 - which, in turn, feeds back on the solution to the continuity equation.
 - Hence the continuity equation for air can not be solved in isolation and one must obey the maximum allowable time-step restrictions imposed by the fastest waves in the system.
- The tracer transport equation can be solved in isolation given prescribed winds and air densities, and is therefore not susceptible to the time-step restrictions imposed by the fastest waves in the system.

Super-cycling (also referred to as sub-cycling) of tracers

- Continuity equation for air is coupled with momentum and thermodynamic equations:
 - thermodynamic variables and other prognostic variables feed back on the velocity field
 - which, in turn, feeds back on the solution to the continuity equation.
 - Hence the continuity equation for air can not be solved in isolation and one must obey the maximum allowable time-step restrictions imposed by the fastest waves in the system.
- The tracer transport equation can be solved in isolation given prescribed winds and air densities, and is therefore not susceptible to the time-step restrictions imposed by the fastest waves in the system.
- For efficiency: Use longer time-step for continuity equation for tracers than for air.



Δt_{dyn} = dynamics time-step; Δt_{trac} = tracer time-step; Δt_{remap} = remap time-step; Δt_{phys} = physics time-step (typically 1800s)

Leads to a major 'speed-up' of dynamics.

Time-steps and namelist variables to control time-steps

The finite-volume fluid flow solver is coded in terms of nested loops:

```
do iv = 1, nv ! vertical re-mapping sub-cycle loop
  do n = 1, n2 ! tracer sub-cycle loop
    do it = 1, nsplit ! dynamics sub-cycle loop

      enddo
    enddo
  enddo
```

where nv , $n2$, and $nsplit$ are defined in terms of 'fv_' namelist variables

```
nv = fv_nspltvrm
n2 = (fv_nsplttrac+nv-1)/fv_nspltvrm
nsplit = (fv_nsplsplit+n2*f_v_nspltvrm-1) / (n2*f_v_nspltvrm)
```

and the time-steps are given by

$$\begin{aligned}\Delta t_{remap} &= \Delta t_{phys} / fv_nspltvrm = 900s \text{ (in CAM } 1^\circ) \\ \Delta t_{trac} &= \Delta t_{phys} / fv_nspltvrm*n2 = 900s \text{ (in CAM } 1^\circ) \\ \Delta t_{dyn} &= \Delta t_{phys} / fv_nspltvrm*n2*f_v_nsplsplit = 225s \text{ (in CAM } 1^\circ)\end{aligned}$$

Free-stream preserving 'super-cycling' of tracers with respect to air ρ

Simply solving the tracer continuity equation for $\overline{q\delta\rho}^{n+1}$ using Δt_{trac} will lead to inconsistencies. Why?

Continuity equation for air $\delta\rho$

$$\frac{\partial\delta\rho}{\partial t} + \nabla \cdot (\delta\rho \vec{v}_h) = 0, \quad (4)$$

and a tracer with mixing ratio q

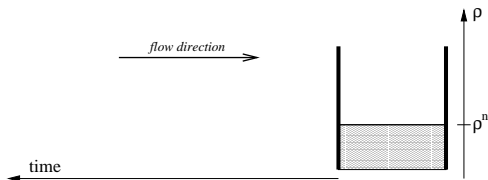
$$\frac{\partial(\delta\rho q)}{\partial t} + \nabla \cdot (\delta\rho q \vec{v}_h) = 0, \quad (5)$$

For $q = 1$ equation (5) reduces to (4). If this is satisfied in the numerical discretizations, the scheme is 'free-stream' preserving.

Solving (5) with $q = 1$ using Δt_{trac} will NOT produce the same solution as solving (4) nsplit times using Δt_{dyn} !

Graphical illustration of 'free stream' preserving transport of tracers

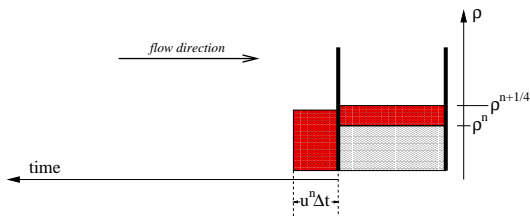
Assume no flux through East cell wall.



- Solve continuity equation for air δp together with momentum and thermodynamics equations.

Graphical illustration of 'free stream' preserving transport of tracers

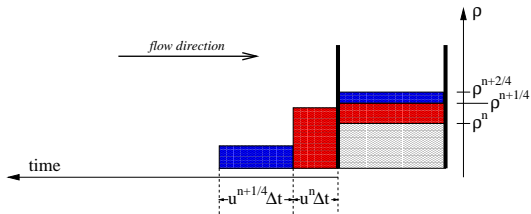
Assume no flux through East cell wall.



- Solve continuity equation for air δp together with momentum and thermodynamics equations.

Graphical illustration of 'free stream' preserving transport of tracers

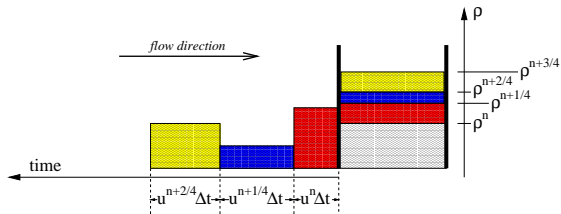
Assume no flux through East cell wall.



- Solve continuity equation for air δp together with momentum and thermodynamics equations.
- Repeat *nsplit* times

Graphical illustration of 'free stream' preserving transport of tracers

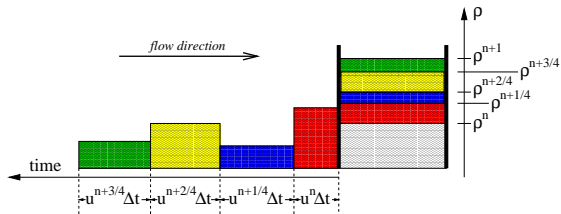
Assume no flux through East cell wall.



- Solve continuity equation for air δp together with momentum and thermodynamics equations.
- Repeat *nsplit* times

Graphical illustration of 'free stream' preserving transport of tracers

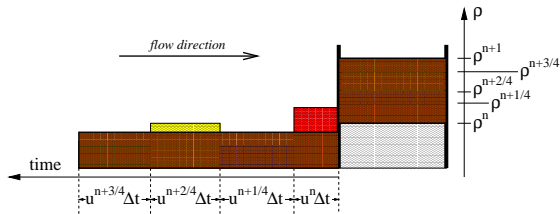
Assume no flux through East cell wall.



- Solve continuity equation for air δp together with momentum and thermodynamics equations.
- Repeat *nsplit* times

Graphical illustration of 'free stream' preserving transport of tracers

Assume no flux through East cell wall.



- Solve continuity equation for air δp together with momentum and thermodynamics equations.
- Repeat $nsplit$ times
- Brown area = average flow of mass through cell face.
- Compute time-averaged value of q across brown area using Lin and Rood (1996) scheme: $\overline{\langle q \rangle}$.
- Forecast for tracer is: $\overline{\langle q \rangle} \times \sum_{i=1}^{nsplit} \delta p^{n+i}/nsplit$
- Yields 'free stream' preserving solution!

- CAM-FV has a very efficient and quite consistent treatment of the tracers.
- This is very important: Number of trace species in climate models are increasing and accounts for most of the computational 'work' in the dynamical core.

- CAM-FV has a very efficient and quite consistent treatment of the tracers.
- This is very important: Number of trace species in climate models are increasing and accounts for most of the computational 'work' in the dynamical core.
- Rasch et al. (2006) did a comprehensive study of the characteristics of atmospheric transport using three dynamical cores in CAM (CAM-FV, CAM-EUL, ~~CAM-STD~~):

What is CAM-EUL? (Collins et al., 2004):

- Based on the spectral transform method and semi-implicit time-stepping
- EUL = Eulerian discretization in grid-point space.
- Tracer transport with non-conservative semi-Lagrangian scheme ('fixers' restore formal mass-conservation)

The results from this study favor use of the CAM-FV core for tracer transport. Compared to CAM-EUL, CAM-FV

- is inherently conservative,
- less diffusive (e.g. maintains strong gradients better),
- maintains thermodynamic relationships among variables more accurately,
- preserved linear-correlations (Lauritzen and Thuburn, 2012) well.

- CAM-FV has a very efficient and quite consistent treatment of the tracers.
- This is very important: Number of trace species in climate models are increasing and accounts for most of the computational 'work' in the dynamical core.
- Rasch et al. (2006) did a comprehensive study of the characteristics of atmospheric transport using three dynamical cores in CAM (CAM-FV, CAM-EUL, ~~CAM-STD~~):

What is CAM-EUL? (Collins et al., 2004):

- Based on the spectral transform method and semi-implicit time-stepping
- EUL = Eulerian discretization in grid-point space.
- Tracer transport with non-conservative semi-Lagrangian scheme ('fixers' restore formal mass-conservation)

The results from this study favor use of the CAM-FV core for tracer transport. Compared to CAM-EUL, CAM-FV

- is inherently conservative,
- less diffusive (e.g. maintains strong gradients better),
- maintains thermodynamic relationships among variables more accurately,
- preserved linear-correlations (Lauritzen and Thuburn, 2012) well.

However, with respect to some other climate statistics CAM-FV needs higher horizontal resolution to produce results equivalent to those produced using the spectral transform dynamical core in CAM (CAM-EUL). **Effective resolution is coarser in CAM-EUL!** See Williamson (2008) for details.

<http://www.cesm.ucar.edu/models/simpler-models/>
<http://www.cesm.ucar.edu/models/simpler-models-indev/> (in development)

Simpler Models

This webpage documents simpler model configurations that are released and supported by the CESM project. As part of CESM2.0, several dynamical core and aquaplanet configurations will be made available. The documentation on these web pages provides information on how to use these configurations and applies to CESM2.0 or later releases. In order to make use of these configurations, users must download CESM2.0 or subsequent releases and guidance on doing that can be found [here](#).

For questions about the aquaplanet configuration, please contact Brian Medeiros (brianpm@ucar.edu) and for questions about the dry dynamical core configuration, please contact Isla Simpson (islas@ucar.edu). If you would like to contribute to the development of other configurations, please contact Lorenzo Polvani (lp@columbia.edu) or Amy Clement (aclement@rsmas.miami.edu).

Currently available simpler models

Atmosphere (CAM)

- [Dry Dynamical Core](#)

CESM Project

CESM is a fully-coupled, community, global climate model that provides state-of-the-art computer simulations of the Earth's past, present, and future climate states.

CESM is sponsored by the National Science Foundation (NSF) and the U.S. Department of Energy (DOE). Administration of the CESM is maintained by the Climate and Global Dynamics Laboratory (CGD) at the National Center for Atmospheric Research (NCAR).

Simpler Models

[Aquaplanet](#)

[Dry Dynamical Core](#)



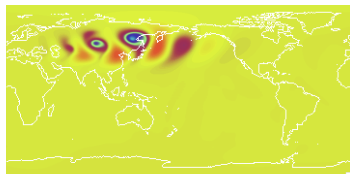
Moist baroclinic wave



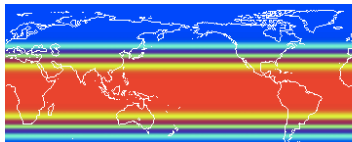
Ullrich et al. (2014): $Q \neq 0$; supports deep atmosphere approximation

P.H.Lauritzen & S.Goldhaber

PS, day 10



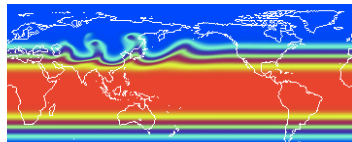
Q, level 976hPa, Day 0



Baroclinic wave used for
DCMIP 2016



Q, level 976hPa, Day 10



```
./create_newcase -compset FADIAB -res ne30_ne30  
./xmlchange -append CAM_CONFIG_OPTS="-analytic_ic"  
echo "analytic_ic_type = 'baroclinic_wave'">> user_nl_cam
```

Moist baroclinic wave with Kessler Micro Physics



Ullrich et al. (2014) baroclinic with 3 tracers (cloud ice, rain water, water vapor)+Kessler (1969) physics

P.H.Lauritzen, C.Zarzycki & S.Goldhaber

A. KESSLER PHYSICS

The cloud microphysics update according to the following equation set:

$$\frac{\Delta \theta}{\Delta t} = -\frac{L}{c_p \bar{\pi}} \left(\frac{\Delta q_{cs}}{\Delta t} + E_r \right) \quad (78)$$

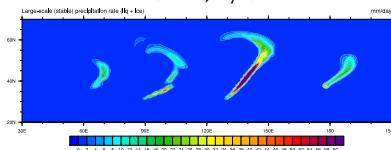
$$\frac{\Delta q_r}{\Delta t} = \frac{\Delta q_{cs}}{\Delta t} + E_r \quad (79)$$

$$\frac{\Delta q_c}{\Delta t} = -\frac{\Delta q_{cs}}{\Delta t} - A_r - C_r \quad (80)$$

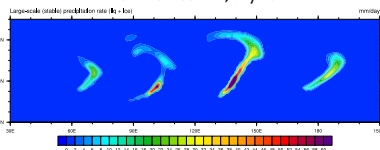
$$\frac{\Delta q_v}{\Delta t} = -E_r + A_r + C_r - V_r \frac{\partial q_r}{\partial z} \quad (81)$$

where L is the latent heat of condensation, A_r is the autoconversion rate of cloud water to rain water, C_r is the collection rate of rain water, E_r is the rain water evaporation rate, and V_r is the rain water terminal velocity.

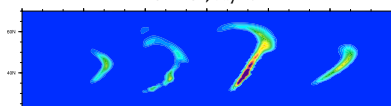
CAM-FV, day 10



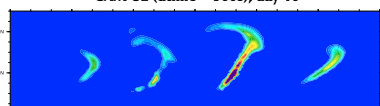
CAM-SE-CSLAM, day 10



CAM-SE, day 10



CAM-SE (dttime = 900s), day 10



`./create_newcase -compset FKESSLER -res ne30_ne30`

Simpler CAM model configurations

- **-compset FHS94**: Held-Suarez test case (Held and Suarez, 1994):
 - Simple Newtonian relaxation of the temperature field to a zonally symmetric state
 - Rayleigh damping of low-level winds representing boundary-layer friction

$$\frac{\partial v}{\partial t} = \dots - k_v(\sigma)v$$

$$\frac{\partial T}{\partial t} = \dots - k_T(\phi, \sigma)[T - T_{\text{eq}}(\phi, p)]$$

$$T_{\text{eq}} = \max \left\{ 200\text{K}, \left[315\text{K} - (\Delta T)_y \sin^2 \phi - (\Delta \theta)_z \log \left(\frac{p}{p_0} \right) \cos^2 \phi \right] \left(\frac{p}{p_0} \right)^{\kappa} \right\}$$

$$k_T = k_a + (k_s - k_a) \max \left(0, \frac{\sigma - \sigma_b}{1 - \sigma_b} \right) \cos^4 \phi$$

$$k_v = k_f \max \left(0, \frac{\sigma - \sigma_b}{1 - \sigma_b} \right)$$

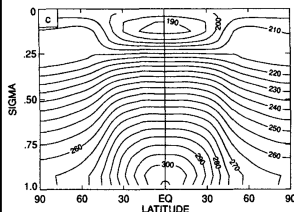
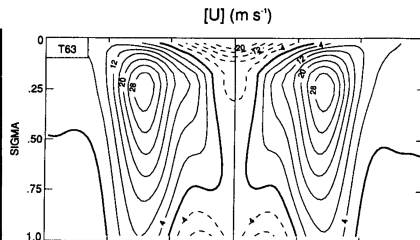
$$\sigma_a = 0.7 \quad k_f = 1 \text{ day}^{-1}$$

$$k_s = 1/40 \text{ day}^{-1} \quad k_a = 1/4 \text{ day}^{-1}$$

$$(\Delta T)_y = 60\text{K} \quad (\Delta \theta)_z = 10\text{K}$$

$$p_0 = 1000 \text{ mb} \quad \kappa = \frac{R}{c_p} = \frac{2}{7} \quad c_p = 1004 \text{ J kg}^{-1} \text{ K}^{-1}$$

$$\Omega = 7.292 \times 10^{-5} \text{ s}^{-1} \quad g = 9.8 \text{ m s}^{-2} \quad a_e = 6.371 \times 10^6 \text{ m}$$



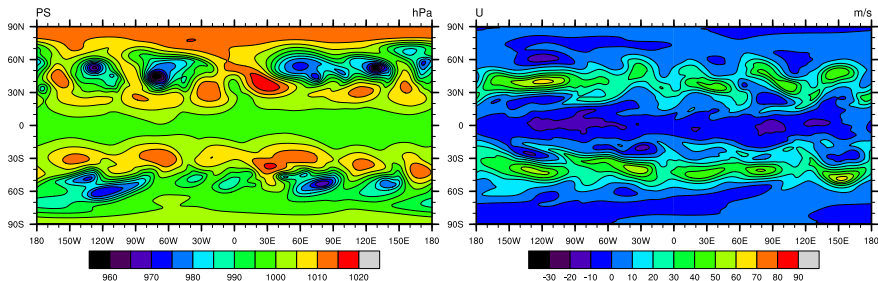
Upper:
Zonal-time
averaged U

Left: Zonal-time
averaged T

Note: this test case can be used to assess how well the dynamical core conserves axial angular momentum (Lebonnois et al., 2012; Lauritzen et al., 2014)

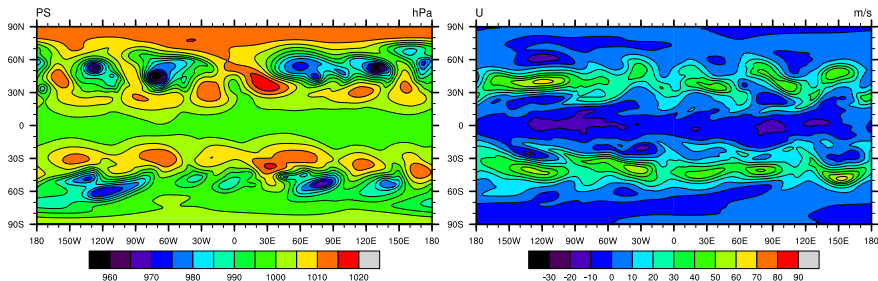
- **-compset FHS94**: Held-Suarez test case (Held and Suarez, 1994):
 - Simple Newtonian relaxation of the temperature field to a zonally symmetric state
 - Rayleigh damping of low-level winds representing boundary-layer friction

Held-Suarez simulation, hour: 12



- **-compset FHS94**: Held-Suarez test case (Held and Suarez, 1994):
 - Simple Newtonian relaxation of the temperature field to a zonally symmetric state
 - Rayleigh damping of low-level winds representing boundary-layer friction

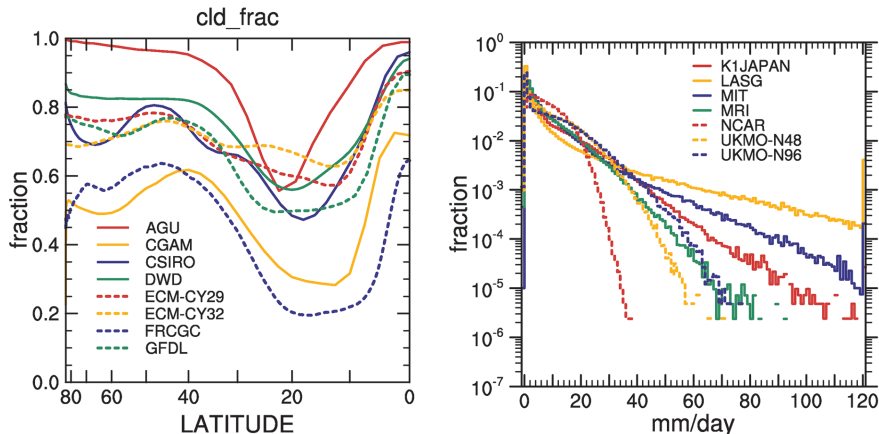
Held-Suarez simulation, hour: 12



- More idealized dynamical core tests: DCMIP (Dynamical Core Model Intercomparison Project; lead by C. Jablonowski); <https://earthsystemcog.org/projects/dcmip-2016/>

Simpler CAM model configurations

- QPC4,5,6: Ocean only planet (referred to as Aquaplanet Earth - APE) with zonally symmetric SST-forcing using 'full' physics package (Neale and Hoskins, 2000). See example of application in Williamson (2008); Blackburn et al. (2013). APE atlas (Williamson et al., 2012).



(left) Zonal-time average cloud fraction.(right) Fraction of time precipitation is in 1 mm/day bins. Figures from Blackburn et al. (2013)

The reformulation of global climate/weather models for massively parallel computer architectures



The reformulation of global climate/weather models for massively parallel computer architectures

Traditionally the equations of motion have been discretized on the traditional regular latitude-longitude grid using either

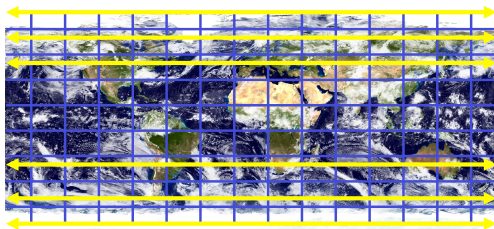
- 1 spherical harmonics based methods (dominated for over 40 years)
- 2 finite-difference/finite-volume methods (e.g., CAM-FV)

Both methods require non-local communication:

- 1 Legendre transform
- 2 'polar^a filters' (due to convergence of the meridians near the poles)

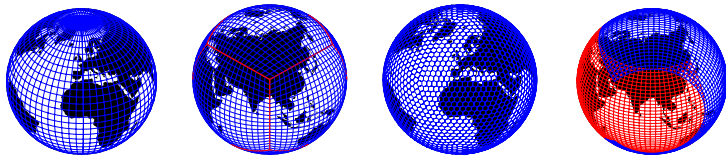
respectively, and are therefore **not** "trivially" amenable for massively parallel compute systems.

^aconfusing terminology: filters are also applied away from polar regions: $\theta \in [\pm 36^\circ, \pm 90]$

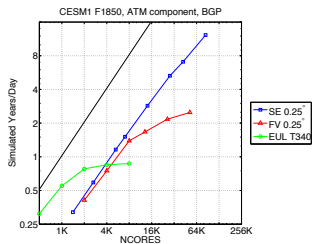


Rectangular computational space

The reformulation of global climate/weather models for massively parallel computer architectures



- Quasi-uniform grid + local numerical method \Rightarrow no non-local communication necessary



Performance in through-put for different dynamical cores in NCAR's global atmospheric climate model:

horizontal resolution: approximately 25km \times 25km grid boxes

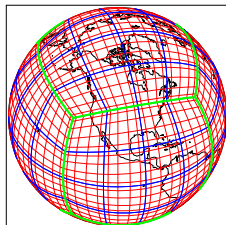
- EUL = spectral transform (lat-lon grid)
- FV = finite-volume (reg. lat-lon grid)
- SE = spectral element (cubed-sphere grid)

Computer = Intrepid (IBM Blue Gene/P Solution) at Argonne National Laboratory

Note that for small compute systems CAM-EUL has SUPERIOR throughput!!

● CAM-SE (Lauritzen et al., 2018): Spectral Elements

- Dynamical core based on HOMME (High-Order Method Modeling Environment, Thomas and Loft 2005).
- Mass-conservative to machine precision and good total energy conservation properties
- Conserves axial angular momentum very well (Lauritzen et al., 2014)
- Discretized on cubed-sphere (uniform resolution or conforming mesh-refinement; Zarzycki et al., 2014) and highly scalable
- 'Work-horse' for high resolution climate applications ($1/4^\circ$)
- New NCAR CAM-SE version using dry-mass vertical coordinates and with comprehensive treatment of condensates and energy released with CESM2
- Optional transport with finite-volume scheme (Lauritzen et al., 2017) and finite-volume physics grid (Herrington et al., 2018)

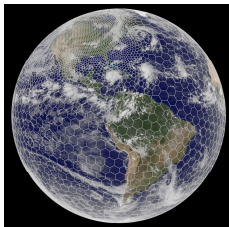


● MPAS (Skamarock et al., 2012): Finite-volume unstructured

- MPAS = Model for Prediction Across Scales
- Centroidal Voronoi tessellation of the sphere
- Fully compressible non-hydrostatic discretization similar to Weather Research Weather (WRF) model (Skamarock and Klemp, 2008)
- Planned integration into CAM

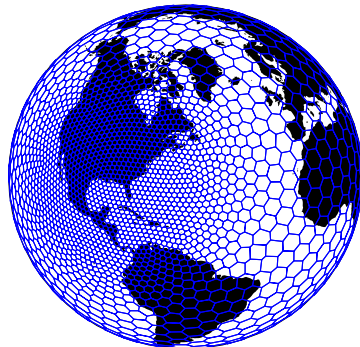
● FV3: Finite-volume

- 'cubed-sphere' version of CAM-FV with non-hydrostatic extension
<https://www.gfdl.noaa.gov/fv3/fv3-documentation-and-references/>
- Ongoing integration into CAM (currently evaluating Aqua-planet simulations)

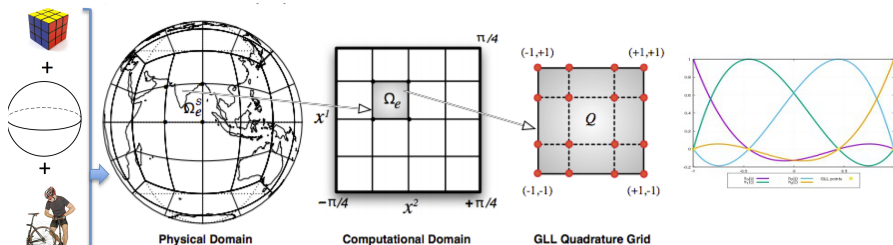


Figures courtesy of R.D. Nair (upper) and W.C. Skamarock (lower).

Both CAM-SE and MPAS support mesh-refinement:



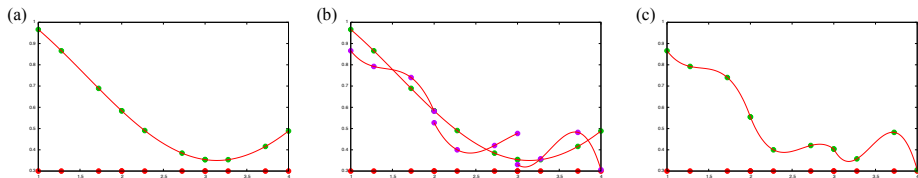
CAM-SE uses a continuous Galerkin finite element method (Taylor et al., 1997) referred to as **Spectral Elements (SE)**:



Figures from Nair et al. (2011)

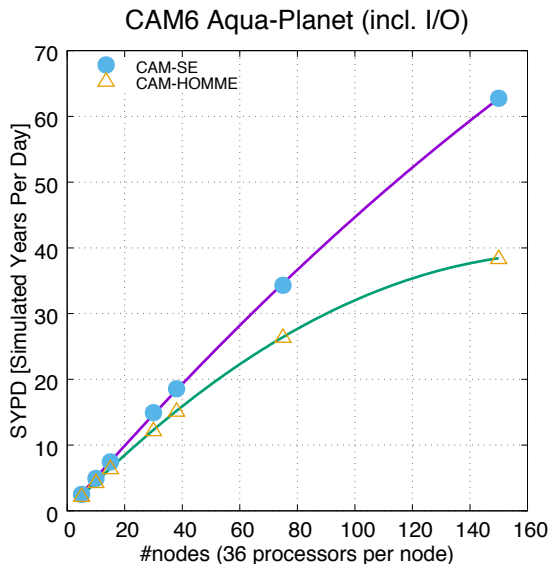
- Physical domain: Tile the sphere with quadrilaterals using the gnomonic cubed-sphere projection
- Computational domain: Mapped local Cartesian domain
- Each element operates with a Gauss-Lobatto-Legendre (GLL) quadrature grid
Gaussian quadrature using the GLL grid will integrate a polynomial of degree $2N - 1$ exactly, where N is degree of polynomial
- Elementwise the solution is projected onto a tensor product of 1D Legendre basis functions by multiplying the equations of motion by test functions; *weak Galerkin formation*
→ all derivatives inside each element can be computed analytically!

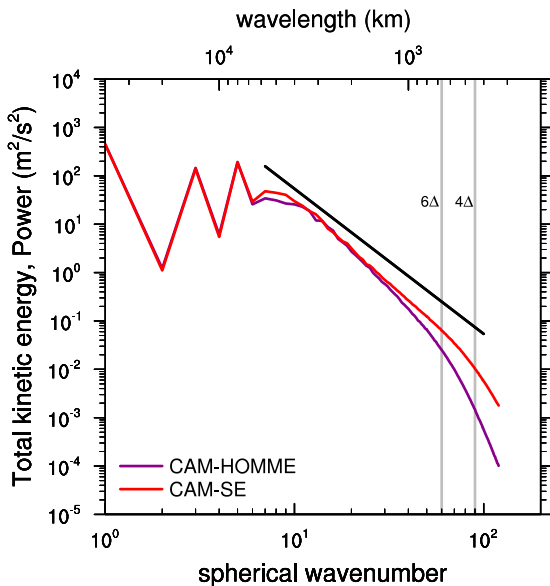
CAM-SE uses a continuous Galerkin finite element method (Taylor et al., 1997) referred to as **Spectral Elements (SE)**:



How do solutions in each element 'communicate' with each other?

- The solution is projected onto the space of globally continuous (C^0) piecewise polynomials
- \rightarrow point values are forced to be C^0 continuous along element boundaries by averaging.
- Note: this is the only operation in which information 'propagates' between elements
- MPI data-communication: only information on the boundary of elements!

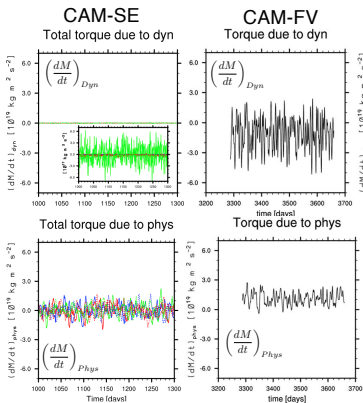




CAM-SE: (Lauritzen et al., 2018) Axial angular momentum conservation (Lauritzen et al., 2014)

M =axial angular momentum integrated over the sphere. For a flat Earth $\frac{dM}{dt} = 0$

In the absence of mountain torque: $0 \sim \left(\frac{dM}{dt}\right)_{dyn} \ll \left(\frac{dM}{dt}\right)_{phys}$



Is conservation of axial angular momentum important?

It is for super-rotating planets (Lebonnois et al., 2012). It is also argued to be important for Earth (Thuburn, 2008); possibly causing bias in CAM-FV (see T.Toniazzo's AMWG presentation from 2015: <http://www.cesm.ucar.edu/events/wg-meetings/2017/presentations/amwg/toniazzo.pdf>)

The total energy equation can be written on the form (Kasahara, 1974)

$$\frac{\partial}{\partial t} \left[\rho \left(\frac{\partial z}{\partial \eta^{(d)}} \right) (K + c_v T + gz) \right] + \nabla_{\eta^{(d)}} \cdot \left[\rho \vec{v} \left(\frac{\partial z}{\partial \eta^{(d)}} \right) (K + gz + c_p T) \right] = - \frac{\partial}{\partial \eta^{(d)}} \left(p \frac{\partial z}{\partial t} \right), \quad (6)$$

where c_p is heat capacity at constant pressure, z height, and $K = \frac{1}{2} \vec{v} \cdot \vec{v}$.

In z -based vertical coordinate (for a moment assume $\eta^{(d)} \equiv z$), then integrating energy equation in the vertical and using that z is constant at the model top (z_{top}) and surface (z_s) we get

$$\frac{\partial}{\partial t} \int_{z=z_s}^{z=z_{top}} (K + c_v T + gz) \rho dz + \nabla_z \cdot \int_{z=z_s}^{z=z_{top}} \vec{v} (K + gz + c_p T) \rho dz = 0. \quad (7)$$

Note: clear separation of kinetic (K), potential (gz) and internal ($c_v T$) energy. Integrating in the horizontal over the entire sphere the flux term drops out, and it is clear that the total energy is conserved for the frictionless and adiabatic system of equations.

In a hybrid-sigma vertical coordinate and assuming that pressure model top is constant we get

$$\frac{1}{g} \frac{\partial}{\partial t} \int_{\eta=0}^{\eta=1} \left(\frac{\partial \mathcal{P}^{(d)}}{\partial \eta^{(d)}} \right) \sum_{\ell} \left[m^{(\ell)} \left(K + c_p^{(\ell)} T + \Phi_s \right) \right] d\eta^{(d)} = 0 \quad (6)$$

where \sum_{ℓ} is sum over dry air, water vapor, cloud liquid, cloud ice, rain and snow ($\ell = 'd', 'wv', 'cl', 'ci', 'rn', 'sw'$).

CAM physics assumes that the 'perfect' dynamical core conserves an energy where $c_p^{(wv)} = c_p^{(d)}$, $c_p^{(\ell)} = 0$ for $\ell = 'cl', 'ci', 'rn', 'sw'$.

One can relatively simply make the continuous equations in CAM-SE conserve the 'CAM physics total energy'. If doing so

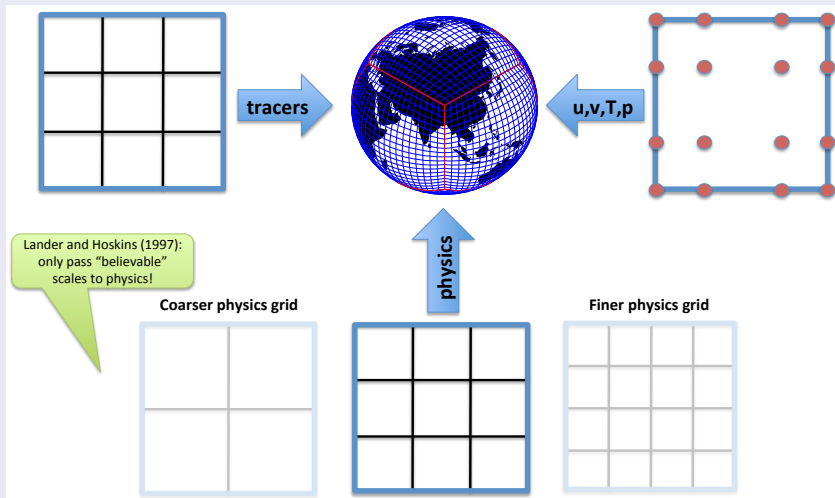
- CAM-SE dynamical core loss of total energy is $0.16W/m^2$ (Aqua-planet simulation).
- For CAM-FV the number is $1.07W/m^2$ (Aqua-planet simulation).

In CESM2 CAM-SE we use the more comprehensive equation that includes condensates in the momentum and thermodynamic equations by default (namelist `se_qsize_condensate_loading=5`).

Why do we care about total energy conservation?

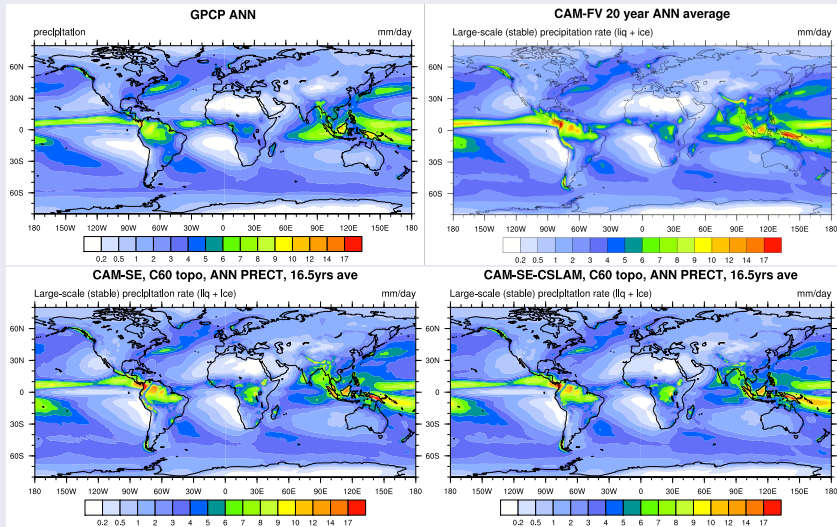
If the total energy budget is not closed in a coupled climate simulation the system will drift ...

CAM-SE has the option to run physics on a finite-volume grid that is coarser, same or finer resolution compared to the dynamics grid. This configuration uses inherently conservative CSLAM (Conservative Semi-Lagrangian Multi-tracer) transport scheme (Lauritzen et al., 2017).

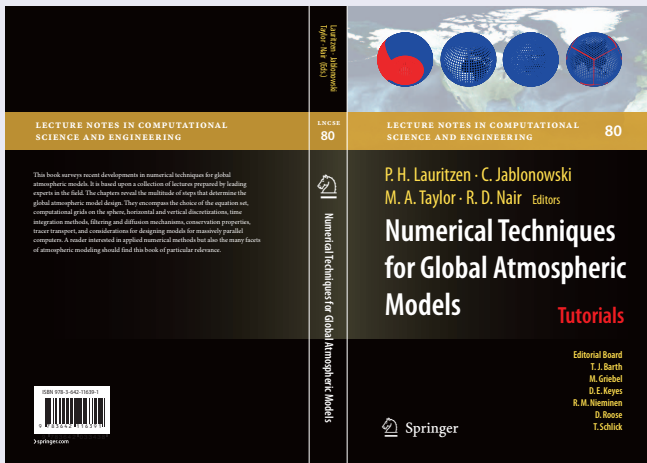


CAM-SE-CSLAM (Herrington et al., 2018) - will be released with CESM2.1

CAM-SE-CSLAM reduces spurious precipitation bias over high orography compared to CAM-FV and CAM-SE (and many other CMIP models)



Interested in numerical methods for global models?



- Book based on the lectures given at the 2008 NCAR ASP (Advance Study Program) Summer Colloquium.
- 16 Chapters; authors include J.Thuburn, J.Tribbia, D.Durran, T.Ringler, W.Skamarock, R.Rood, J.Dennis, Editors, ... Foreword by D. Randall
- More details at: <http://www.cgd.ucar.edu/cms/pel/colloquium.html> and <http://www.cgd.ucar.edu/cms/pel/lncse.html>

Questions?



- Adcroft, A., Hill, C., and Marshall, J. (1997). Representation of topography by shaved cells in a height coordinate ocean model. *Mon. Wea. Rev.*, 125(9):2293–2315.
- Arakawa, A. and Lamb, V. R. (1977). Computational design and the basic dynamical processes of the UCLA general circulation model. *Methods in Computational Physics*, 17:172–265.
- Blackburn, M., WILLIAMSON, D. L., NAKAJIMA, K., OHFUCHI, W., TAKAHASHI, Y. O., HAYASHI, Y.-Y., NAKAMURA, H., ISHIWATARI, M., MCGREGOR, J. L., BORTH, H., WIRTH, V., FRANK, H., BECHTOLD, P., WEDI, N. P., TOMITA, H., SATOH, M., ZHAO, M., HELD, I. M., SUAREZ, M. J., LEE, M.-I., WATANABE, M., KIMOTO, M., LIU, Y., WANG, Z., MOLOD, A., RAJENDRAN, K., KITO, A., and STRATTON, R. (2013). The aqua-planet experiment (ape): Control sst simulation. *Journal of the Meteorological Society of Japan. Ser. II*, 91A:17–56.
- Colella, P. and Woodward, P. R. (1984). The piecewise parabolic method (PPM) for gas-dynamical simulations. *J. Comput. Phys.*, 54:174–201.
- Collins, W. D., Rasch, P. J., Boville, B. A., Hack, J. J., McCaa, J. R., Williamson, D. L., Kiehl, J. T., and Briegleb, B. (2004). Description of the NCAR Community Atmosphere Model (CAM 3.0). NCAR Tech. Note, NCAR/TN-464+STR.
- Dennis, J. M., Edwards, J., Evans, K. J., Guba, O., Lauritzen, P. H., Mirin, A. A., St-Cyr, A., Taylor, M. A., and Worley, P. H. (2012). CAM-SE: A scalable spectral element dynamical core for the Community Atmosphere Model. *Int. J. High Perform. C.*, 26(1):74–89.
- Gross, M., Wan, H., Rasch, P. J., Caldwell, P. M., Williamson, D. L., Klocke, D., Jablonowski, C., Thatcher, D. R., Wood, N., Cullen, M., Beare, B., Willett, M., Lemarié, F., Blayo, E., Malardel, S., Termonia, P., Gassmann, A., Lauritzen, P. H., Johansen, H., Zarzycki, C. M., Sakaguchi, K., and Leung, R. (2016). Recent progress and review of issues related to Physics Dynamics Coupling in geophysical models. *ArXiv e-prints*.
- Held, I. M. and Suarez, M. J. (1994). A proposal for the intercomparison of the dynamical cores of atmospheric general circulation models. *Bull. Amer. Meteor. Soc.*, 75:1825–1830.
- Herrington, A. R., Lauritzen, P. H., Taylor, M. A., Goldhaber, S., Eaton, B. E., Reed, K. A., and Ullrich, P. A. (2018). Physics-dynamics coupling with element-based high-order Galerkin methods: quasi equal-area physics grid. *Mon. Wea. Rev.* revising.
- Kasahara, A. (1974). Various vertical coordinate systems used for numerical weather prediction. *Mon. Wea. Rev.*, 102(7):509–522.
- Lauritzen, P. and Thuburn, J. (2012). Evaluating advection/transport schemes using interrelated tracers, scatter plots and numerical mixing diagnostics. *Quart. J. Roy. Met. Soc.*, 138(665):906–918.
- Lauritzen, P. H. (2007). A stability analysis of finite-volume advection schemes permitting long time steps. *Mon. Wea. Rev.*, 135:2658–2673.
- Lauritzen, P. H., Bacmeister, J. T., Dubos, T., Lebonnois, S., and Taylor, M. A. (2014). Held-Suarez simulations with the Community Atmosphere Model Spectral Element (CAM-SE) dynamical core: A global axial angular momentum analysis using Eulerian and floating Lagrangian vertical coordinates. *J. Adv. Model. Earth Syst.*, 6.
- Lauritzen, P. H., Mirin, A., Truesdale, J., Raeder, K., Anderson, J., Bacmeister, J., and Neale, R. B. (2011a). Implementation of new diffusion/filtering operators in the CAM-FV dynamical core. *Int. J. High Perform. Comput. Appl.*

- Lauritzen, P. H., Nair, R., Herrington, A., Callaghan, P., Goldhaber, S., Dennis, J., Bacmeister, J. T., Eaton, B., Zarzycki, C., Taylor, M. A., Gettelman, A., Neale, R., Dobbins, B., Reed, K., and Dubos, T. (2018). NCAR CESM2.0 release of CAM-SE: A reformulation of the spectral-element dynamical core in dry-mass vertical coordinates with comprehensive treatment of condensates and energy. *J. Adv. Model. Earth Syst.*
- Lauritzen, P. H., Taylor, M. A., Overfelt, J., Ullrich, P. A., Nair, R. D., Goldhaber, S., and Kelly, R. (2017). CAM-SE-CSLAM: Consistent coupling of a conservative semi-lagrangian finite-volume method with spectral element dynamics. *Mon. Wea. Rev.*, 145(3):833–855.
- Lauritzen, P. H., Ullrich, P. A., and Nair, R. D. (2011b). Atmospheric transport schemes: desirable properties and a semi-Lagrangian view on finite-volume discretizations, in: P.H. Lauritzen, R.D. Nair, C. Jablonowski, M. Taylor (Eds.), Numerical techniques for global atmospheric models. *Lecture Notes in Computational Science and Engineering, Springer, 2011*, 80.
- Lebonnois, S., Covey, C., Grossman, A., Parish, H., Schubert, G., Walterscheid, R., Lauritzen, P. H., and Jablonowski, C. (2012). Angular momentum budget in general circulation models of superrotating atmospheres: A critical diagnostic. *J. Geo. Res.: Planets*, 117(E12):n/a–n/a.
- Lin, S. J. (1997). Ti: A finite-volume integration method for computing pressure gradient force in general vertical coordinates. *Quart. J. Roy. Meteor. Soc.*, 123:1749–1762.
- Lin, S.-J. (2004). A 'vertically Lagrangian' finite-volume dynamical core for global models. *Mon. Wea. Rev.*, 132:2293–2307.
- Lin, S. J. and Rood, R. B. (1996). Multidimensional flux-form semi-Lagrangian transport schemes. *Mon. Wea. Rev.*, 124:2046–2070.
- Lin, S.-J. and Rood, R. B. (1997). An explicit flux-form semi-Lagrangian shallow-water model on the sphere. *Q.J.R.Meteorol.Soc.*, 123:2477–2498.
- Neale, R. B. and Hoskins, B. J. (2000). A standard test for agcms including their physical parametrizations: I: the proposal. *Atmos. Sci. Lett.*, 1(2):101–107.
- Rasch, P. J., Coleman, D. B., Mahowald, N., Williamson, D. L., Lin, S. J., Boville, B. A., and Hess, P. (2006). Characteristics of atmospheric transport using three numerical formulations for atmospheric dynamics in a single GCM framework. *J. Climate*, 19:2243–2266.
- Schär, C., Leuenberger, D., Fuhrer, O., Lüthi, D., and Girard, C. (2002). A new terrain-following vertical coordinate formulation for atmospheric prediction models. *Mon. Wea. Rev.*, 130(10):2459–2480.
- Skamarock, W. (2011). Kinetic energy spectra and model filters, in: P.H. Lauritzen, R.D. Nair, C. Jablonowski, M. Taylor (Eds.), Numerical techniques for global atmospheric models. *Lecture Notes in Computational Science and Engineering, Springer*, 80.
- Skamarock, W. C. (2008). A linear analysis of the NCAR CCSM finite-volume dynamical core. *Mon. Wea. Rev.*, 136:2112–2119.
- Skamarock, W. C. and Klemp, J. B. (2008). A time-split nonhydrostatic atmospheric model for weather research and forecasting applications. *J. Comput. Phys.*, 227:3465–3485.
- Skamarock, W. C., Klemp, J. B., Duda, M. G., Fowler, L. D., Park, S.-H., and Ringler, T. D. (2012). A multiscale nonhydrostatic atmospheric model using centroidal Voronoi tessellations and C-grid staggering. *Mon. Wea. Rev.*, 140:3090–3105.
- Taylor, M. A., Tribbia, J., and Iskandarani, M. (1997). The spectral element method for the shallow water equations on the sphere. *J. Comput. Phys.*, 130:92–108.
- Thomas, S. J. and Loft, R. D. (2005). The NCAR spectral element climate dynamical core: Semi-implicit Eulerian formulation. *J. Sci. Comput.*, 25:307–322.

- Thuburn, J. (2008). Some conservation issues for the dynamical cores of NWP and climate models. *J. Comput. Phys.*, 227:3715 – 3730.
- Thuburn, J. (2011). Some basic dynamics relevant to the design of atmospheric model dynamical cores, in: P.H. Lauritzen, R.D. Nair, C. Jablonowski, M. Taylor (Eds.), Numerical techniques for global atmospheric models. *Lecture Notes in Computational Science and Engineering*, Springer, 80.
- van Leer, B. (1977). Towards the ultimate conservative difference scheme. IV: A new approach to numerical convection. *J. Comput. Phys.*, 23:276–299.
- Whitehead, J., Jablonowski, C., Rood, R. B., and Lauritzen, P. H. (2011). A stability analysis of divergence damping on a latitude-longitude grid. *Mon. Wea. Rev.*, 139:2976–2993.
- Williamson, D.L., Blackburn, M., Hoskins, B., Nakajima, K., Ohfuchi, W., Takahashi, Y., Hayashi, Y.-Y., Nakamura, H., Ishiwatari, M., McGregor, J., Borth, H., Wirth, V., Frank, H., Bechtold, P., Wedi, N., Tomita, H., Satoh, M., Zhao, M., Held, I., Suarez, M., Lee, M.-I., Watanabe, M., Kimoto, M., Liu, Y., Wang, Z., Molod, A., Rajendran, K., Kitoh, A., , and Stratton, R. (2012). The ape atlas. *NCAR Technical Note*, NCAR/TN-484+STR.
- Williamson, D. L. (2002). Time-split versus process-split coupling of parameterizations and dynamical core. *Mon. Wea. Rev.*, 130:2024–2041.
- Williamson, D. L. (2008). Equivalent finite volume and Eulerian spectral transform horizontal resolutions established from aqua-planet simulations. *Tellus*, 60:839–847.
- Williamson, D. L. and Olson, J. G. (2003). Dependence of aqua-planet simulations on time step. *Q. J. R. Meteorol. Soc.*, 129(591):2049–2064.
- Zarzycki, C. M., Jablonowski, C., and Taylor, M. A. (2014). Using variable-resolution meshes to model tropical cyclones in the community atmosphere model. *Mon. Wea. Rev.*, 142(3):1221–1239.