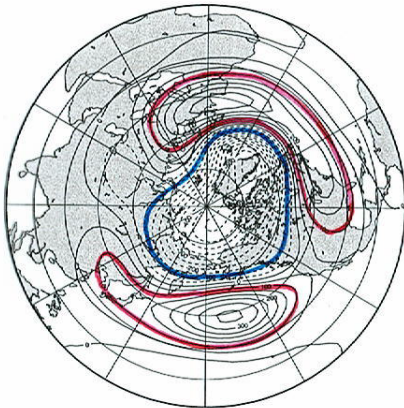


# 3D Spectral Energetics Analysis and Rossby Wave Saturation Theory

Barotropic Component of Geopotential Height  
EOF-1 AO (5.7%)

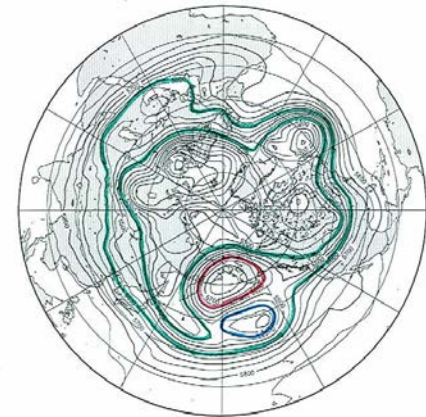


Hiroshi L. Tanaka

*University of Tsukuba  
Japan*



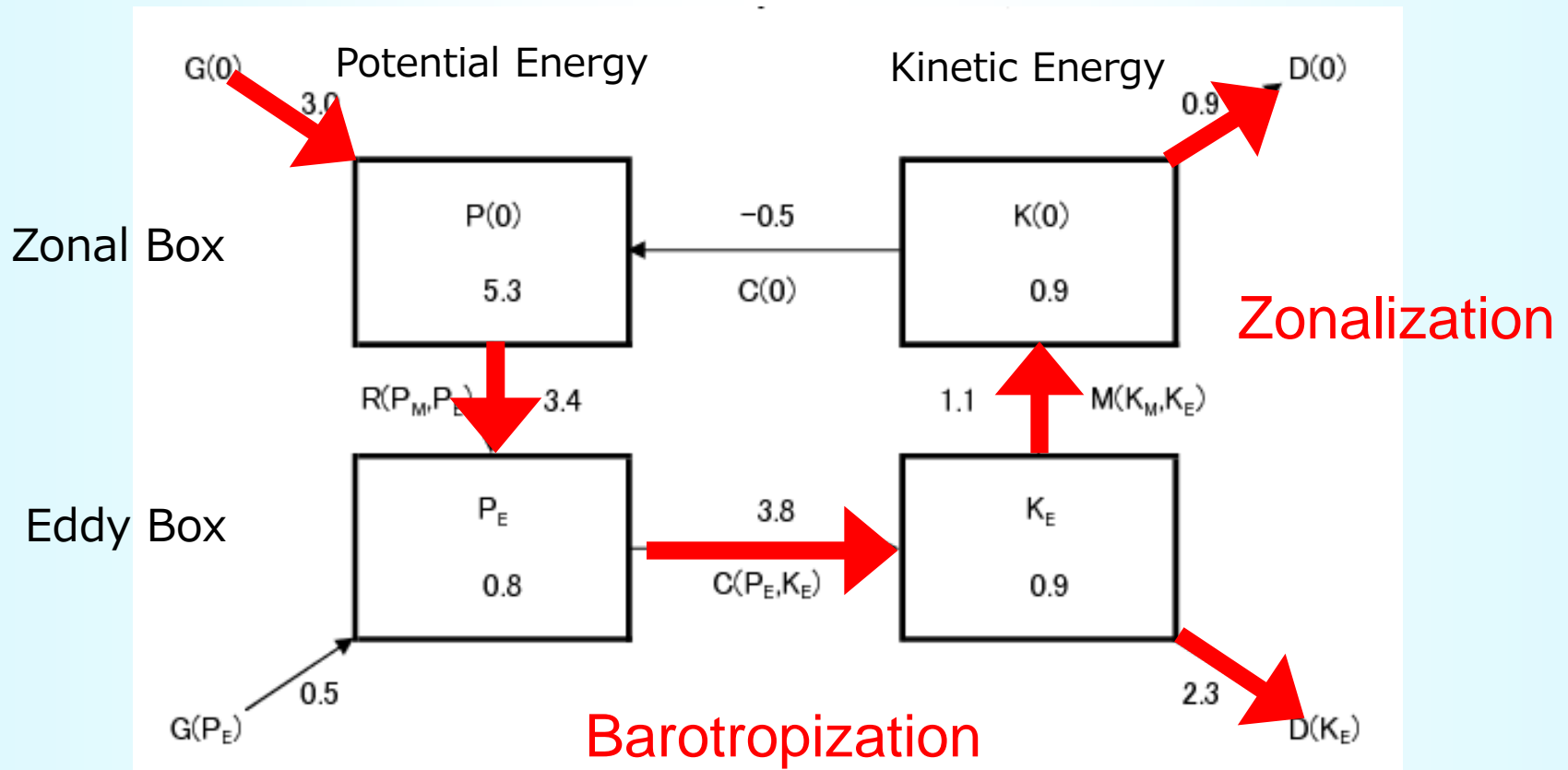
500 hPa Height  
JMA GPV 97031412+00



(Presentation at MODES 2015)

# Lorenz Energy Box Diagram

Energy accumulates at zonal wave 0



Energy accumulates at vertical wave 0

(Kung and Tanaka 1983, JAS)

# Lorenz cycle, Saltzman cycle

(Saltzman 1957; 1970)  $p = \sum_{n=-\infty}^{\infty} p_n \exp(inx)$

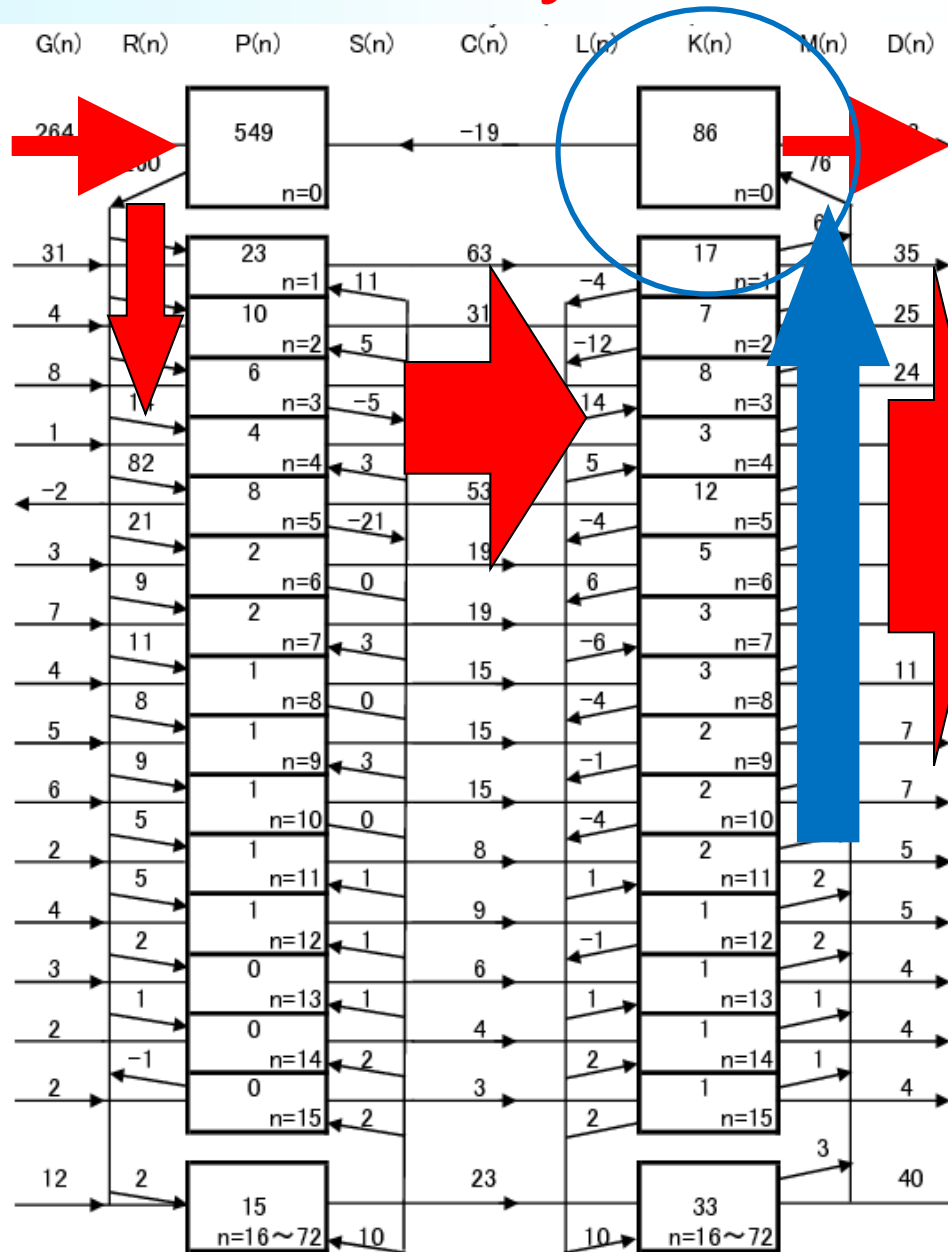
$$\boxed{K(0)} \quad \frac{\partial K_Z}{\partial t} = \sum_{n=1}^N M(n) + C(0) - D(0),$$

$$\boxed{K(n)} \quad \frac{\partial K(n)}{\partial t} = -M(n) + L(n) + C(n) - D(n), \quad n = 1, 2, 3, \dots$$

$$\boxed{P(0)} \quad \frac{\partial P_Z}{\partial t} = -\sum_{n=1}^N R(n) - C(0) + G(0),$$

$$\boxed{P(n)} \quad \frac{\partial P(n)}{\partial t} = R(n) + S(n) - C(n) + G(n), \quad n = 1, 2, 3, \dots$$

# Saltzman cycle



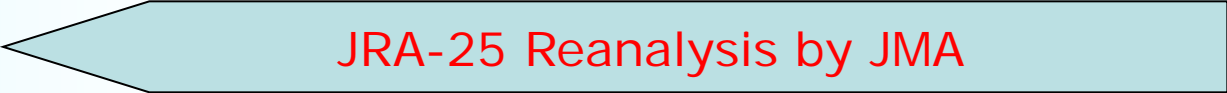


- G: Generation of P(n)
- P: Available potential energy
- R: zonal-wave interaction of P(n)
- S: wave-wave interaction of P(n)
- C: Baroclinic conversion from P(n) to K(n)
- K: Kinetic energy
- M: zonal-wave interaction of K(n)
- L: wave-wave interaction of K(n)
- D: Dissipation of K(n)

(Saltzman 1957 & 1970)

(Kung and Tanaka 1983 & 1984)

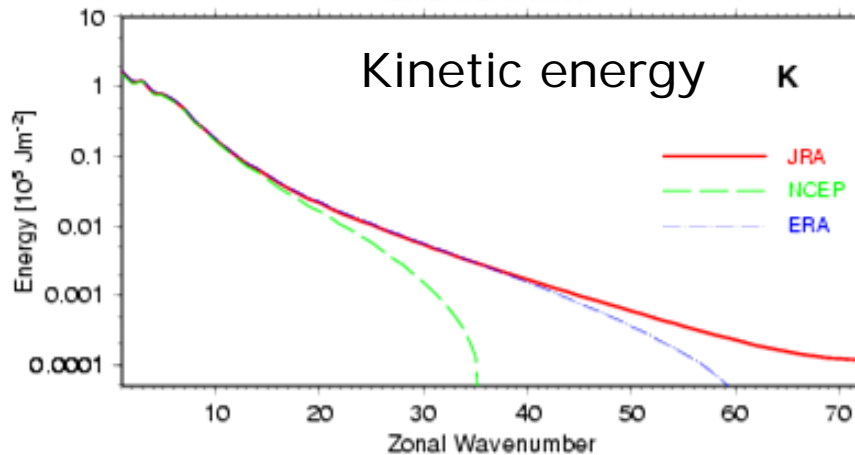
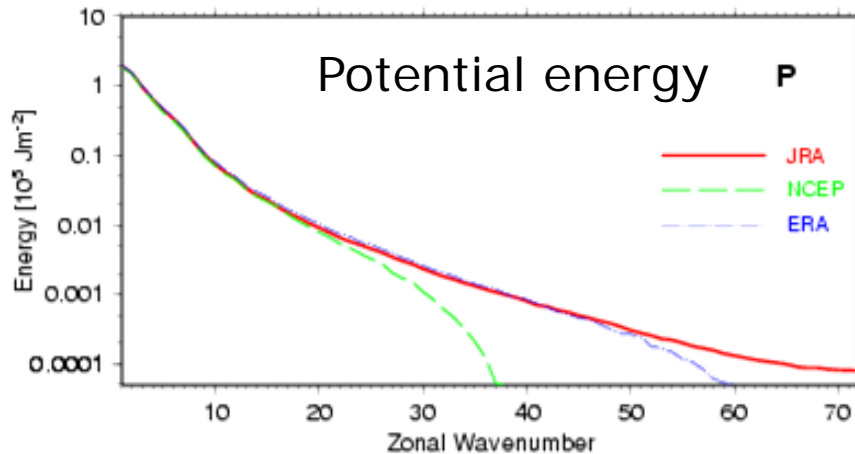


# Data

- JRA-25  JRA-25 Reanalysis by JMA
    - $2.5^\circ \times 2.5^\circ$  , 23 levels (1000 - 0.4 hPa)
  - NCEP/NCAR reanalysis  NOAA CDC
    - $2.5^\circ \times 2.5^\circ$  , 17 levels (1000 - 10 hPa)
  - ERA-40  ECMWF
    - $2.5^\circ \times 2.5^\circ$  , 23 levels (1000 - 1 hPa)
- 1990/91 DJF (3 Month)
    - u, v, T, q

# Available Potential Energy $P(n)$ Kinetic energy $K(n)$

- NCEP up to  $n=35$
- ERA up to  $n=60$

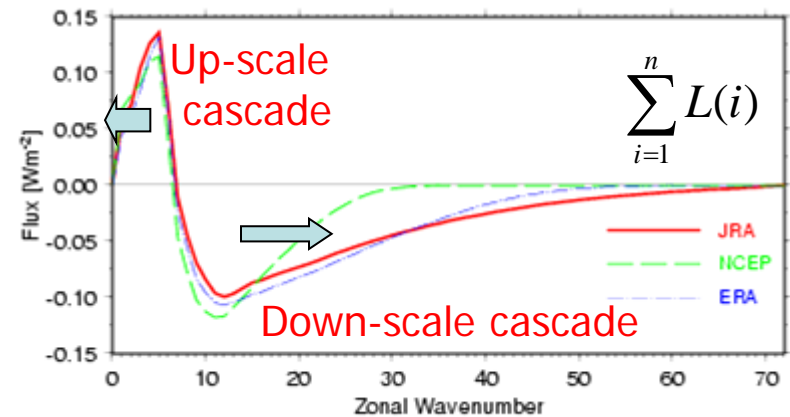
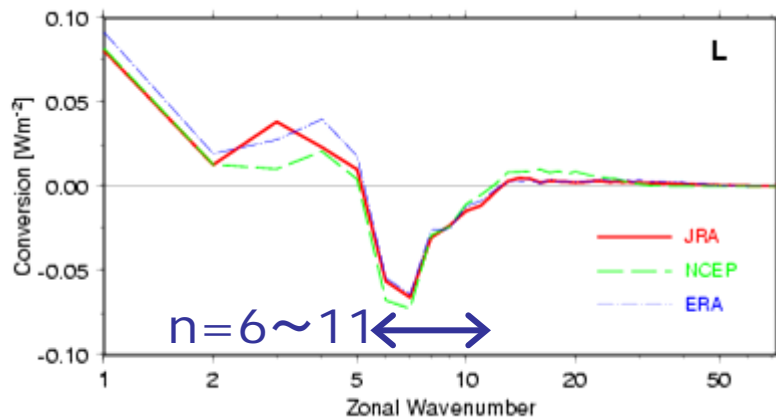
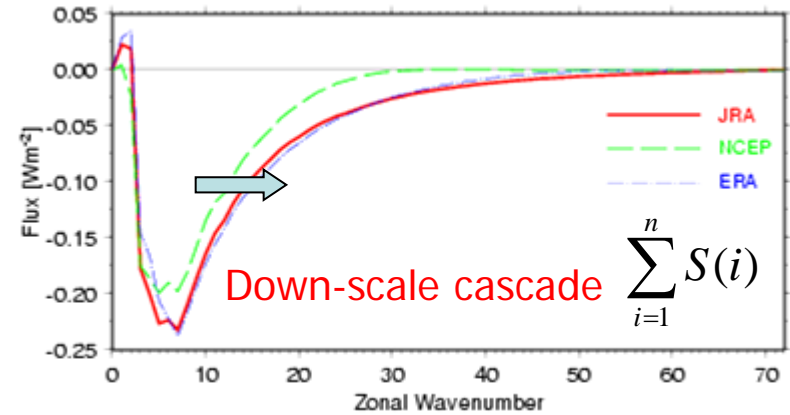
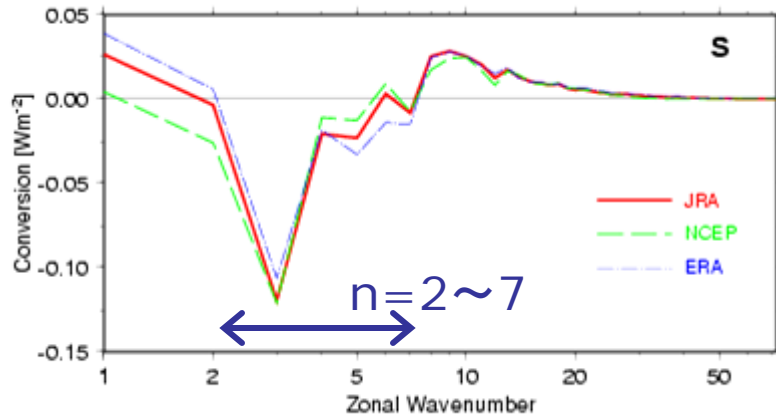


total energy [ $10^4 \text{ Jm}^{-2}$ ]

	NCEP	JRA	ERA
$P_E$	64	67	70
$K_E$	77	82	84

# Wave-wave interactions (S, L)

## Energy Flux



# 2D Spectral model

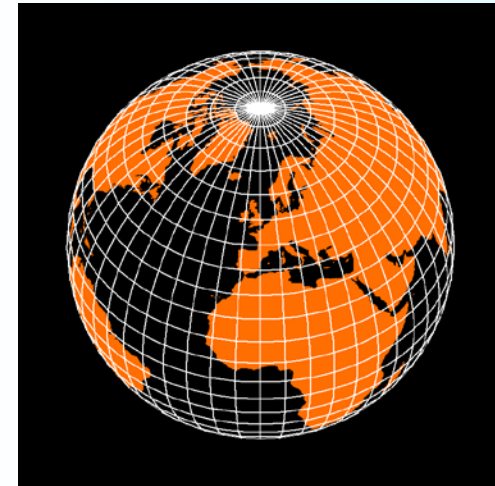
- 1D: Expansion in **Fourier harmonics**
- 2D: Expansion in **spherical harmonics**

$$p = \sum_{n=-\infty}^{\infty} p_n \exp(inx)$$

$$\frac{\partial p}{\partial x} = \sum_{n=-\infty}^{\infty} in p_n \exp(inx)$$

$$Y_l^n(\lambda, \theta) = P_l^n(\theta) \exp(in\lambda)$$

$$p(\lambda, \theta) = \sum_{n=-N}^N \sum_{l=|n|}^L p_{nl} Y_l^n(\lambda, \theta)$$



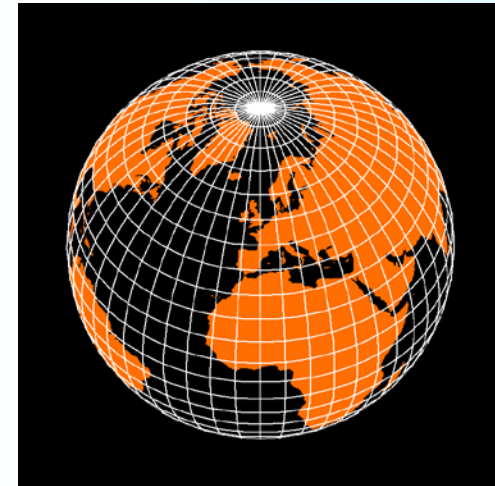
# 3D Spectral model

- Vertical normal mode
- Horizontal normal mode: **Hough harmonics**

Expansion in 3D Normal Mode Functions

$$\Pi_{nlm}(\lambda, \theta, \sigma) = \Theta_{nlm}(\theta) G_m(\sigma) \exp(in\lambda)$$

$$U(\lambda, \theta, \sigma) = \sum_{n=-N}^N \sum_{l=0}^L \sum_{m=0}^M w_{nlm} X_m \Pi_{nlm}(\lambda, \theta, \sigma)$$



$$\frac{dw_i}{d\tau} = -i\sigma_i w_i - i \sum_{jk} r_{ijk} w_j w_k + f_i, \quad i = 1, 2, 3, \dots$$

$$E_i = \frac{1}{2} p_s h_m |w_i|^2, \quad w_{nlm} \rightarrow w_i$$

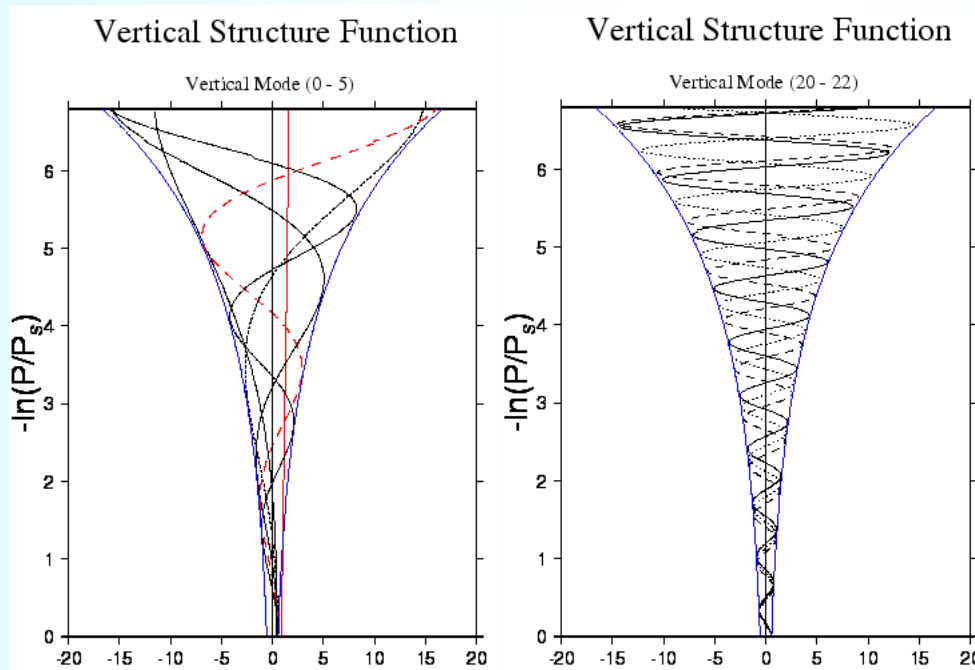
# Vertical energy spectrum

Vertical modes  $\frac{\partial}{\partial \sigma} \left( \sigma^2 \frac{\partial G_m}{\partial \sigma} \right) + \lambda_m G_m = 0, \quad \lambda_m = \frac{R\gamma}{gh_m}$

$G_0(\sigma) = C_1 \sigma^{r_1} + C_2 \sigma^{r_2}$

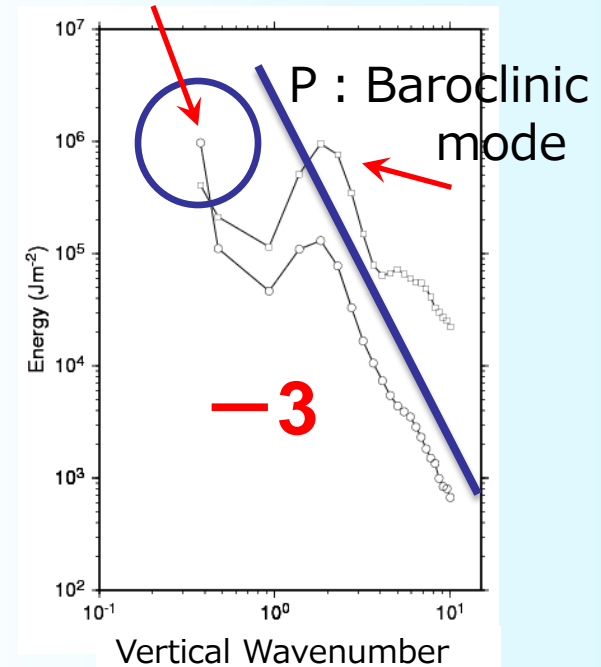
Barotropic and baroclinic modes

$G_m(\sigma) = \sigma^{-\frac{1}{2}} \left( C_1 \sin(\mu \ln \sigma) + C_2 \cos(\mu \ln \sigma) \right) \quad \mu = \sqrt{\lambda_m - \frac{1}{4}}$



Terasaki and Tanaka (2007)

K : Barotropic

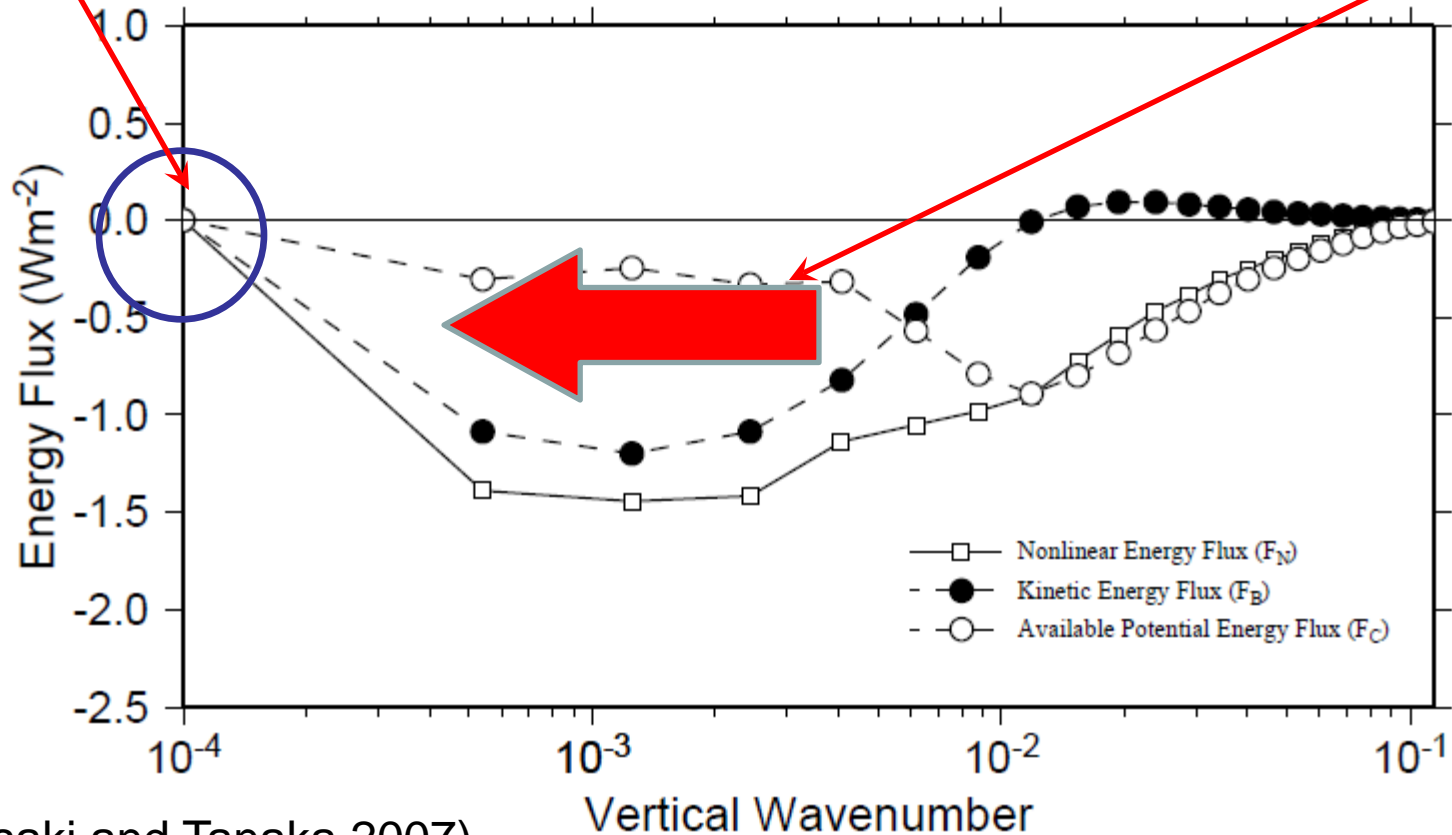


# Barotropization by baroclinic instability

Barotropic  
mode

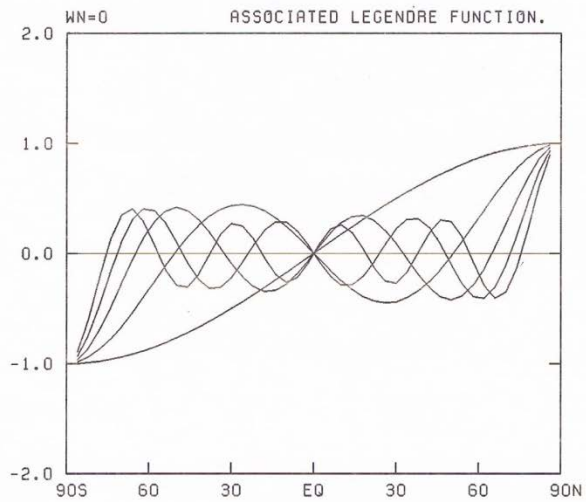
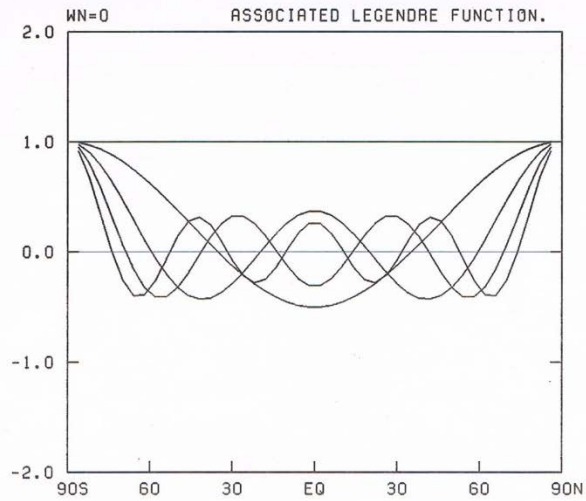
Energy Flux (DJF climate)

Energy flux of  
Kinetic energy



(Terasaki and Tanaka 2007)

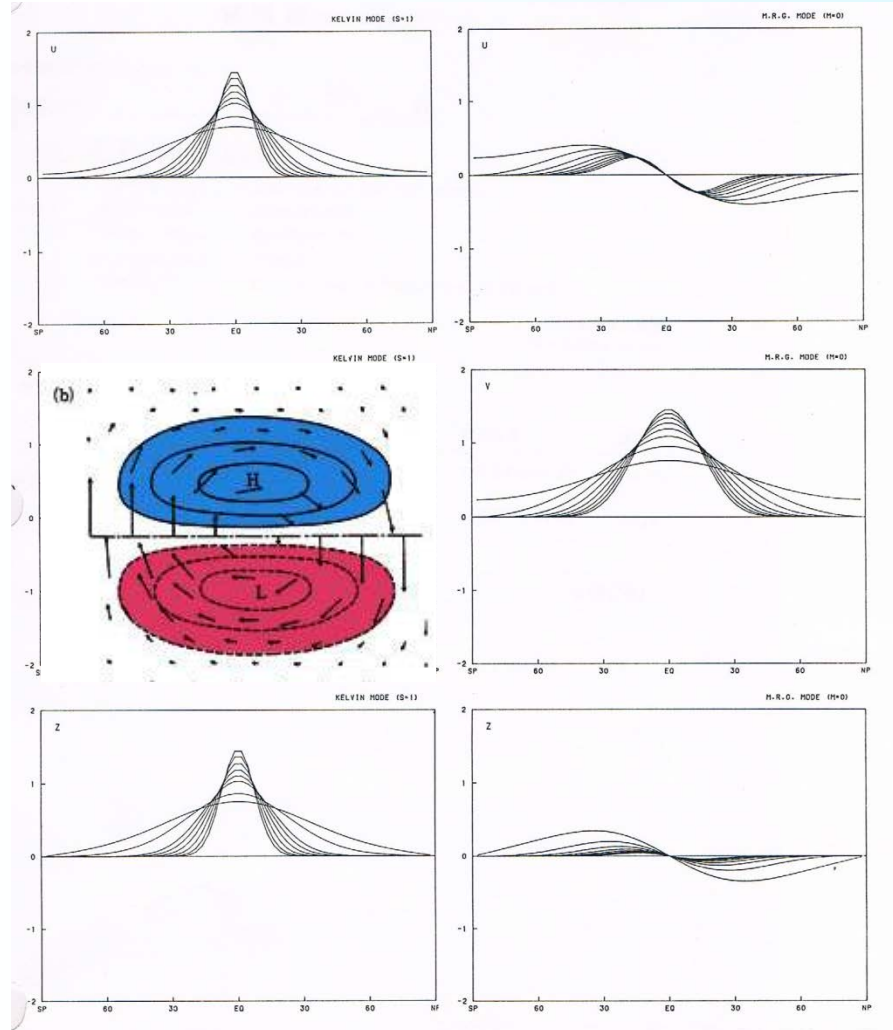
# Spherical harmonics (n=0)



# Hough harmonics

Kelvin mode

Mixed Rossby-gravity mode





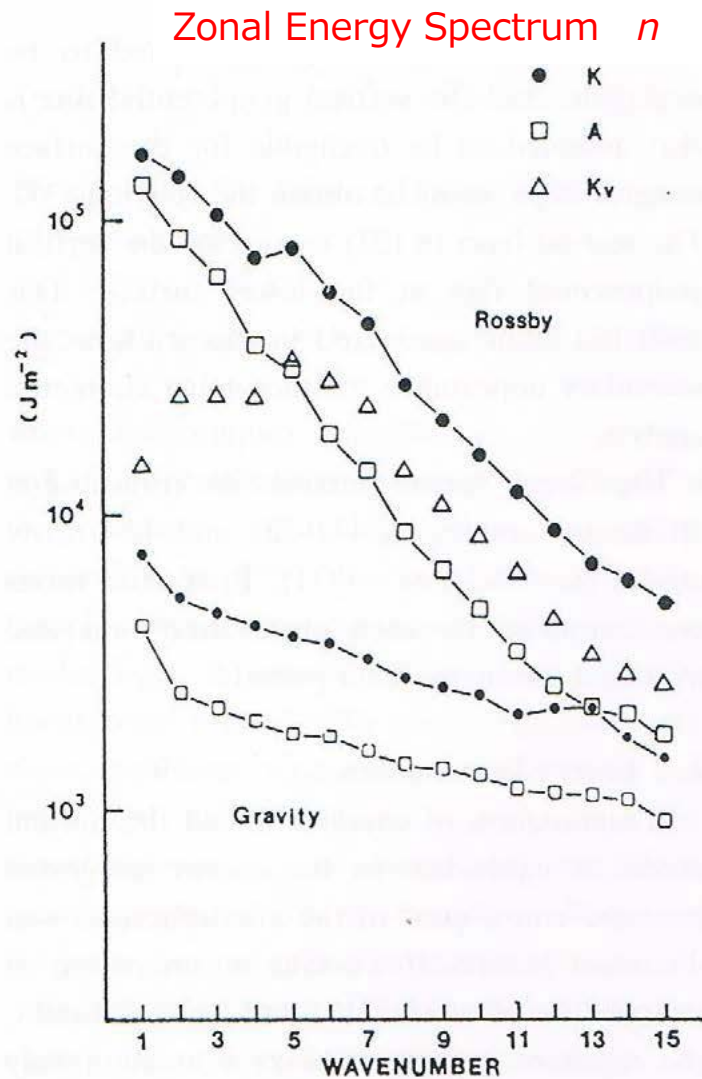


Fig. 2. Energy distributions in the wavenumber domain.  $K$  : kinetic energy,  $A$  : available potential energy,  $K_v$  :  $v$ -component of  $K$

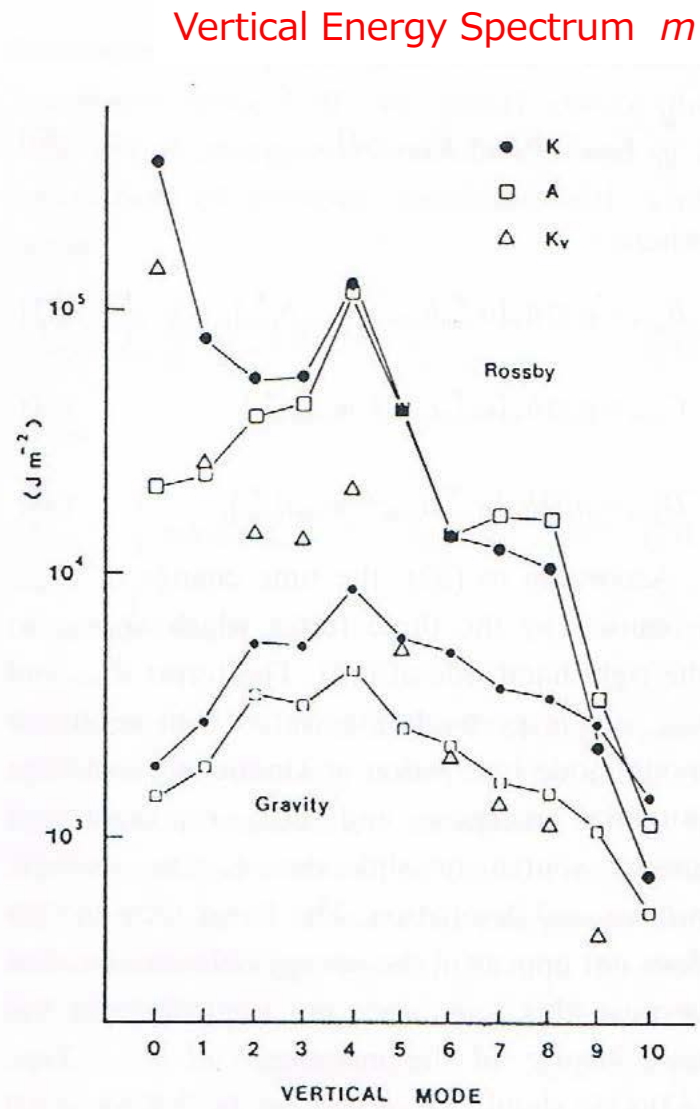
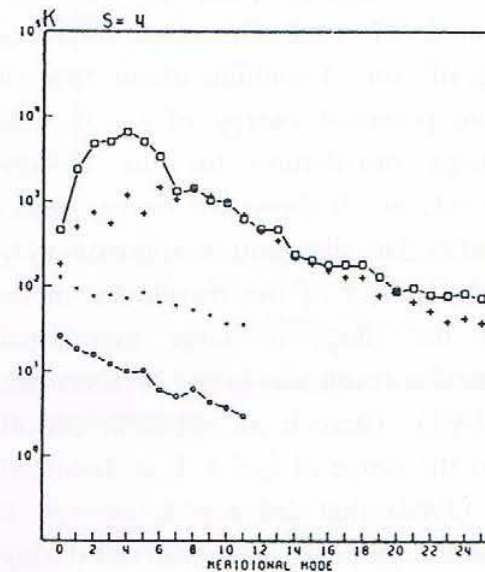
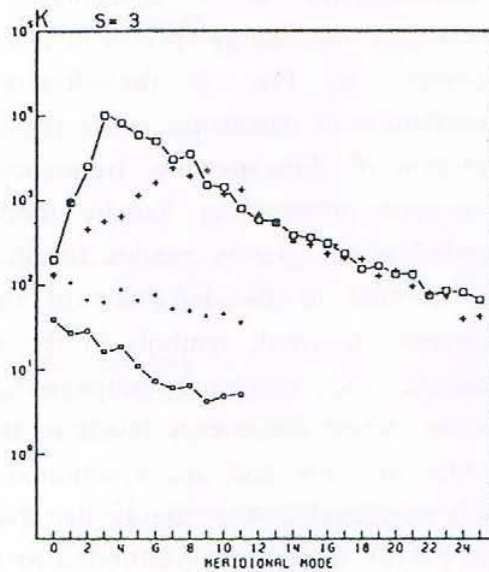
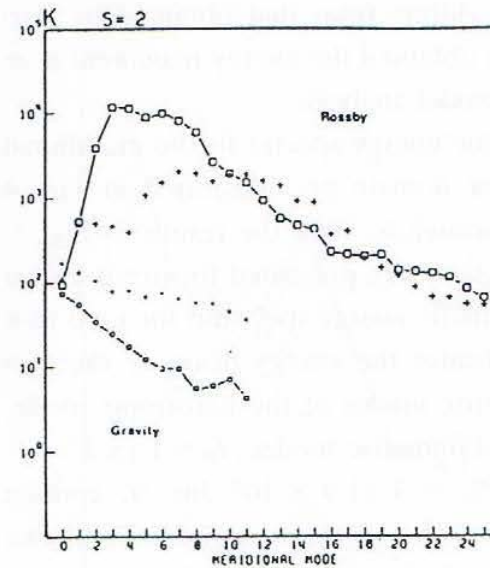
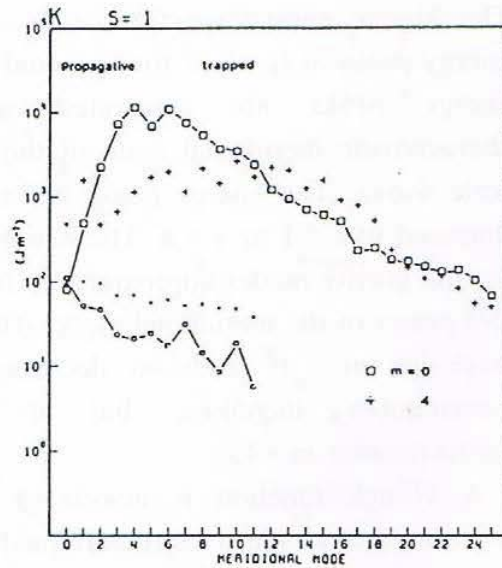


Fig. 3. Eddy energy distributions in the vertical mode domain.

# Meridional Energy Spectrum /



# Energy spectrum in the 3D wavenumber space

$$c = \frac{\beta}{n^2 + l^2 + m^2} = \frac{\beta}{k^2}$$

$n, l, m$  : zonal, meridional and vertical waves

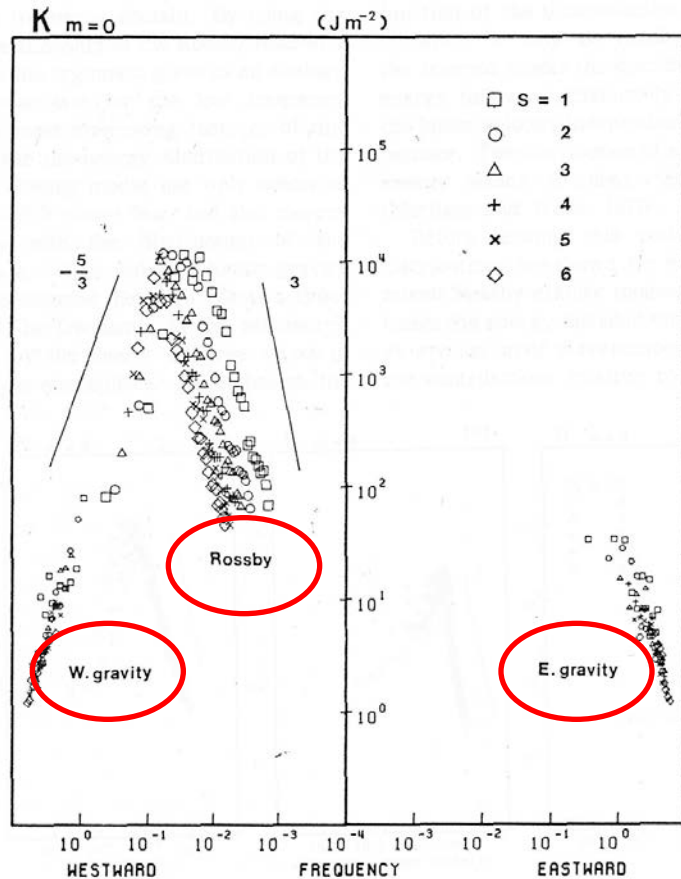
$k$  : total wave  $c = \sigma / n$

$$\frac{dw_i}{d\tau} = -i\sigma_i w_i - i \sum_{jk} r_{ijk} w_j w_k + f_i$$

Use  $c$  for the scale in place of 3D wavenumber

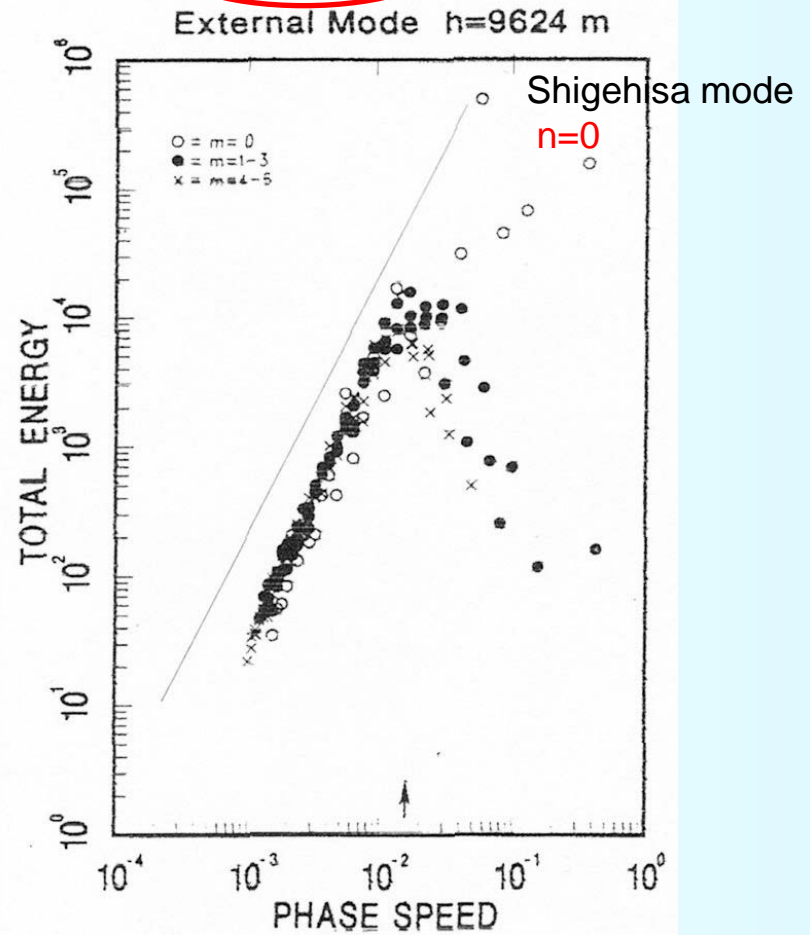
# Observed energy spectrum in **c-domain**

Frequency domain



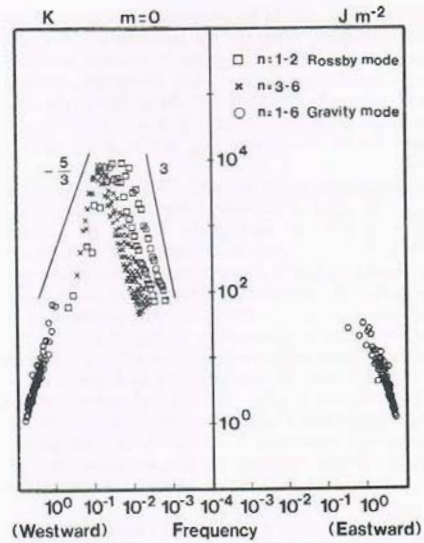
Tanaka (1985)

Phase speed domain

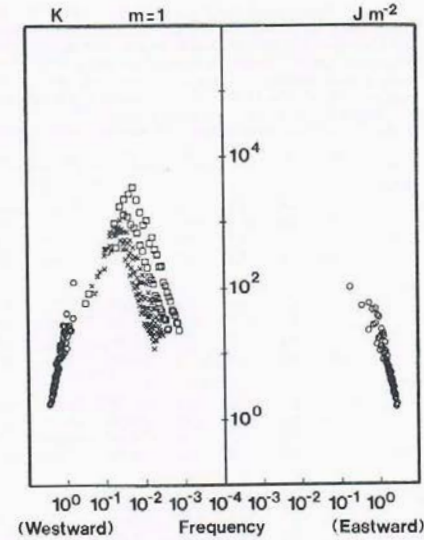


Tanaka and Kasahara (1992)

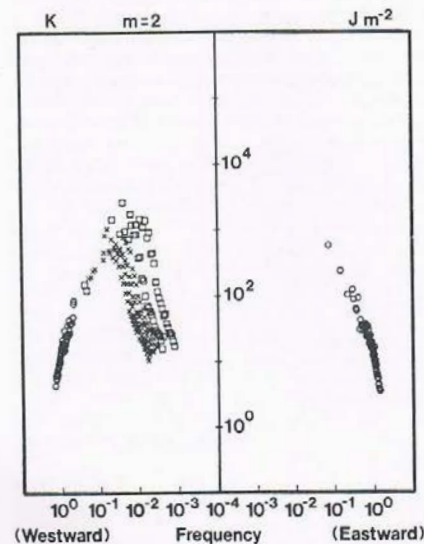
Vertical mode  
 $m=0$



Vertical mode  
 $m=1$



Vertical mode  
 $m=2$



Vertical mode  
 $m=3$

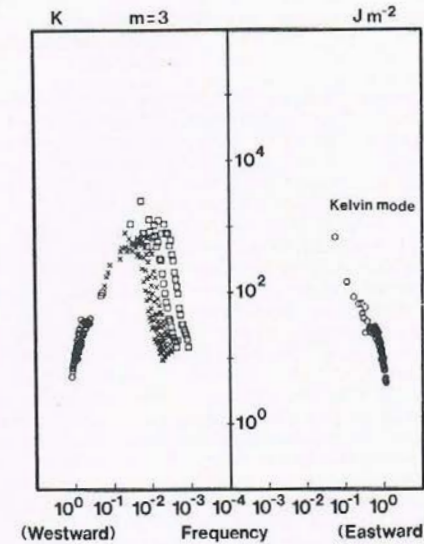
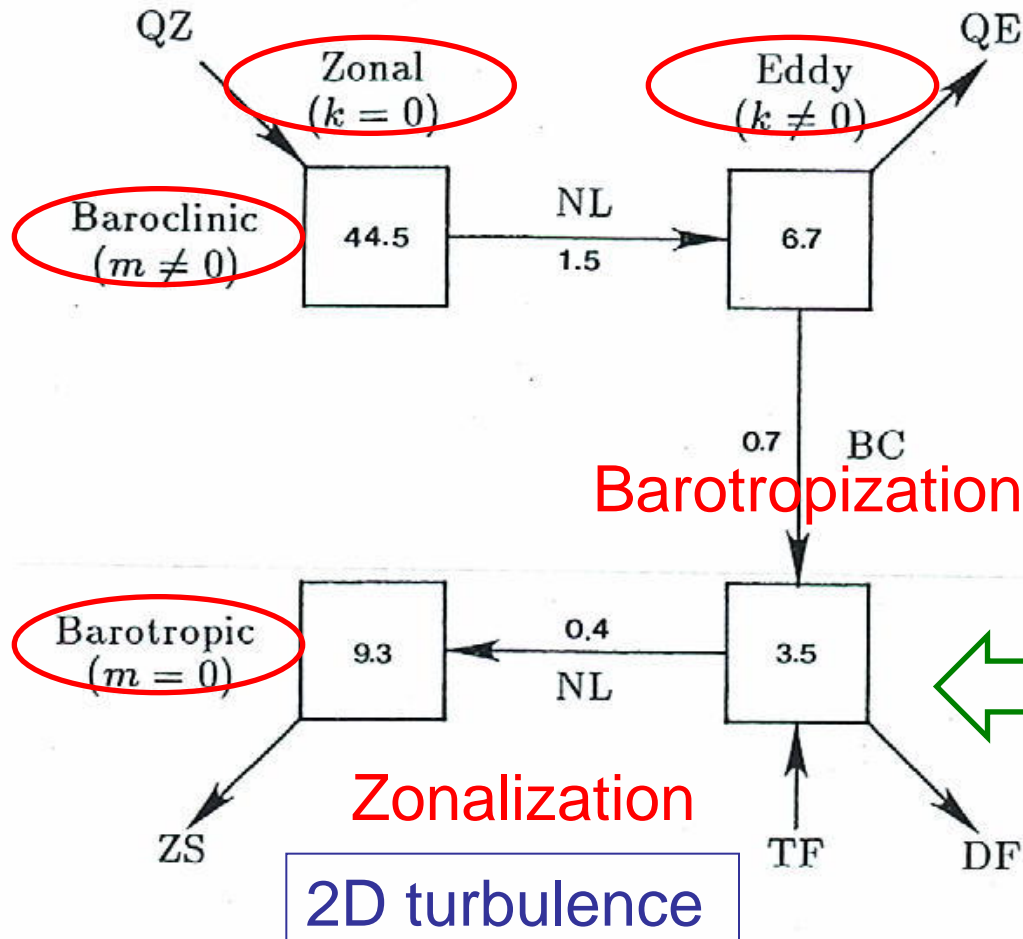


FIG. 3. Kinetic energy spectra in the dimensionless frequency domain (normalized by  $2\Omega$ ) for vertical indices  $m = 0-3$ . The energy of the Rossby modes and gravity modes are plotted for wavenumber  $n = 1-6$ . Note that the frequency on the abscissa is the eigenfrequency of Laplace's tidal equation rather than the analyzed wave frequency in the space-time spectra.

# Energy flow box diagram in barotropic-baroclinic decomposition



## Primitive Equation Model

$$M \frac{\partial U}{\partial t} + L U = N + F, \quad (1)$$

where

$$U = (u, v, \phi')^T, \quad (2)$$

$$M = \text{diag}(1, 1, -\frac{\partial}{\partial p} \frac{p^2}{R\gamma} \frac{\partial}{\partial p}), \quad (3)$$

$$L = \begin{pmatrix} 0 & -2\Omega \sin\theta & \frac{1}{a \cos\theta} \frac{\partial}{\partial \lambda} \\ 2\Omega \sin\theta & 0 & \frac{1}{a} \frac{\partial}{\partial \theta} \\ \frac{1}{a \cos\theta} \frac{\partial}{\partial \lambda} & \frac{1}{a \cos\theta} \frac{\partial(\cdot) \cos\theta}{\partial \theta} & 0 \end{pmatrix}, \quad (4)$$

$$N = \begin{pmatrix} -V \cdot \nabla u - \omega \frac{\partial u}{\partial p} + \frac{\tan\theta}{a} uv \\ -V \cdot \nabla v - \omega \frac{\partial v}{\partial p} - \frac{\tan\theta}{a} vu \\ \frac{\partial}{\partial p} \left( \frac{p^2}{R\gamma} V \cdot \nabla \frac{\partial \phi}{\partial p} + \omega p \left( \frac{p}{R\gamma} \frac{\partial \phi}{\partial p} \right) \right) \end{pmatrix}, \quad (5)$$

$$F = (F_u, F_v, \frac{\partial}{\partial p} \left( \frac{pQ}{C_p \gamma} \right))^T. \quad (6)$$

Numerical  
simulations  
of blocking  
and AO

Primitive  
Equation  
Model



## Barotropic Component of the Atmosphere

### ○ Vertical Transform

$$(u, v, \phi')_0^T = \frac{1}{p_s} \int_0^{p_s} \underline{(u, v, \phi')^T} G_0 dp \quad (1)$$

### ○ Barotropic Model

$$\frac{\partial u}{\partial t} = -\vec{v} \cdot \nabla u + fv - \frac{\partial \phi}{\partial x} + F_x \quad (2)$$

$$\frac{\partial v}{\partial t} = -\vec{v} \cdot \nabla v - fu - \frac{\partial \phi}{\partial y} + F_y \quad (3)$$

$$\frac{\partial \phi}{\partial t} = -\vec{v} \cdot \nabla \phi - \bar{\phi} \nabla \cdot \vec{v} + F_z \quad (4)$$

### ○ 3-D Spectral Transform

$$\underline{U(\lambda, \theta, p, t)} = \sum_{nlm} \underline{w_{nlm}(t)} X_m \Pi_{nlm}(\lambda, \theta, p), \quad (5)$$

$$\underline{w_{nlm}(t)} = \langle \underline{U(\lambda, \theta, p, t)}, X_m^{-1} \Pi_{nlm} \rangle \quad (6)$$

where  $U(\lambda, \theta, p, t) = (u, v, \phi')^T$ ,  $w_{nlm}(t)$  is the spectral expansion coefficient,  $X_m = \text{diag}(c_m, c_m, c_m^2)$ , and  $\underline{\Pi_{nlm}}$  is the 3-D NMF.

# Numerical simulations of blocking and AO

$$\frac{dw_i}{d\tau} = -i\sigma_i w_i - i \sum_{jk} r_{ijk} w_j w_k + f_i$$

$$E_i = \frac{1}{2} p_s h_m |w_i|^2$$

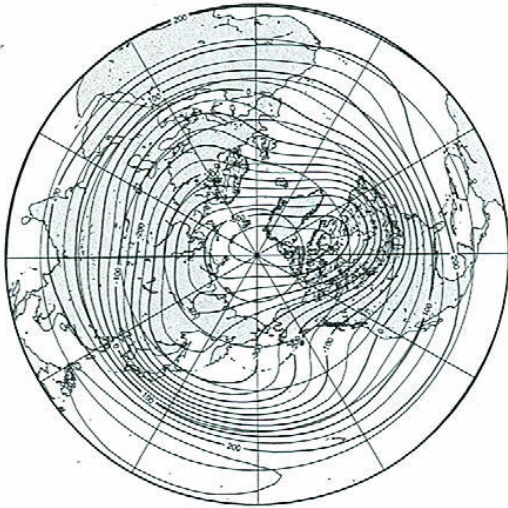
Barotropic S-Model  
(Tanaka 2003, JAS)



# NCEP/NCAR

Barotropic Height

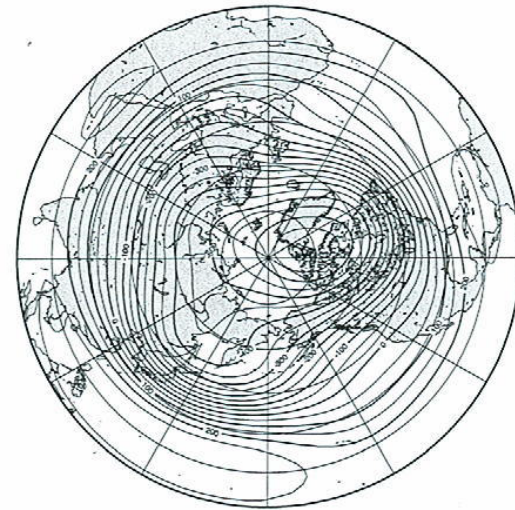
DJF mean for 1950-2000



# Barotropic S-Model

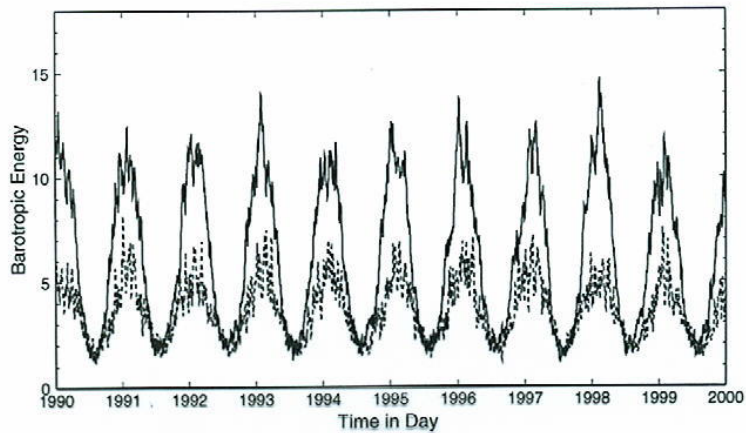
Geopotential Height

Perpetual January (50 Years)



# Barotropic energy

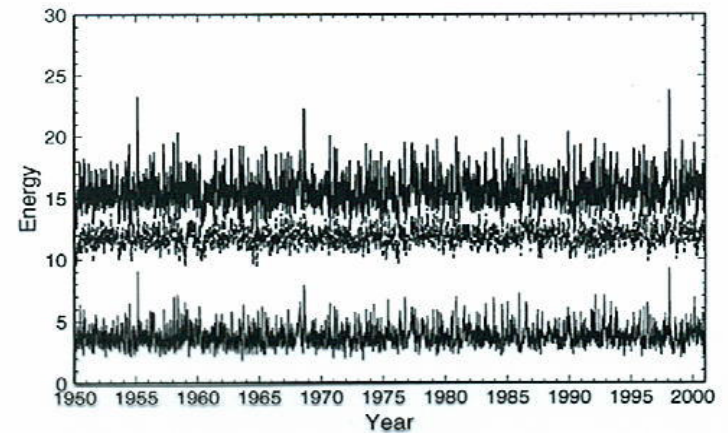
NCEP/NCAR



# Perpetual January

Barotropic Energy (S-Model)

Perpetual January



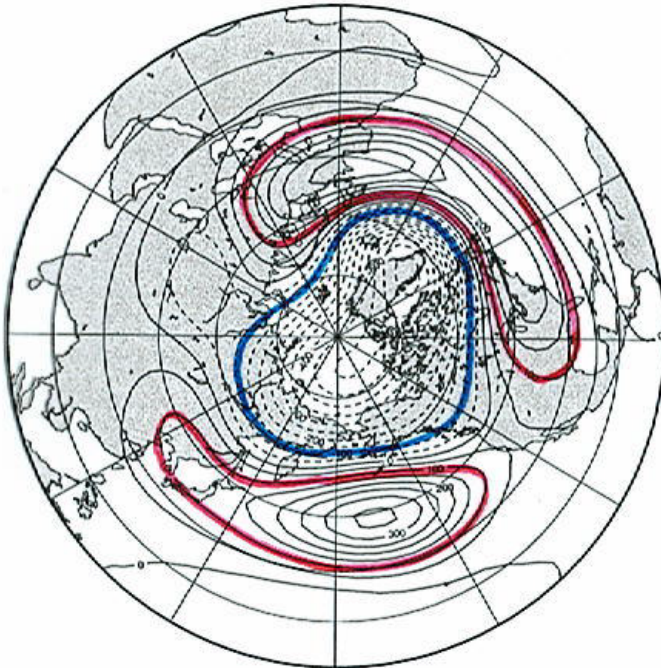
# Arctic Oscillation

## Barotropic height (EOF-1)

NCEP/NCAR

Barotropic Component of Geopotential Height

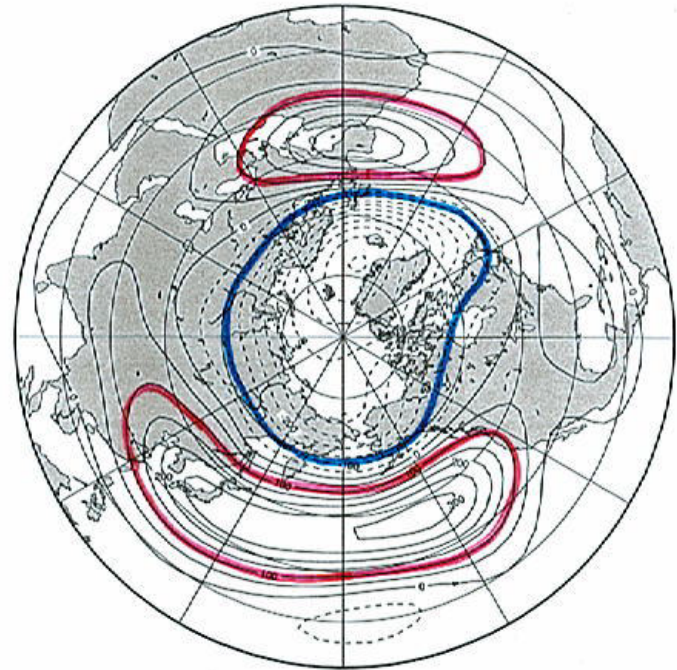
EOF-1 AO (5.7%)



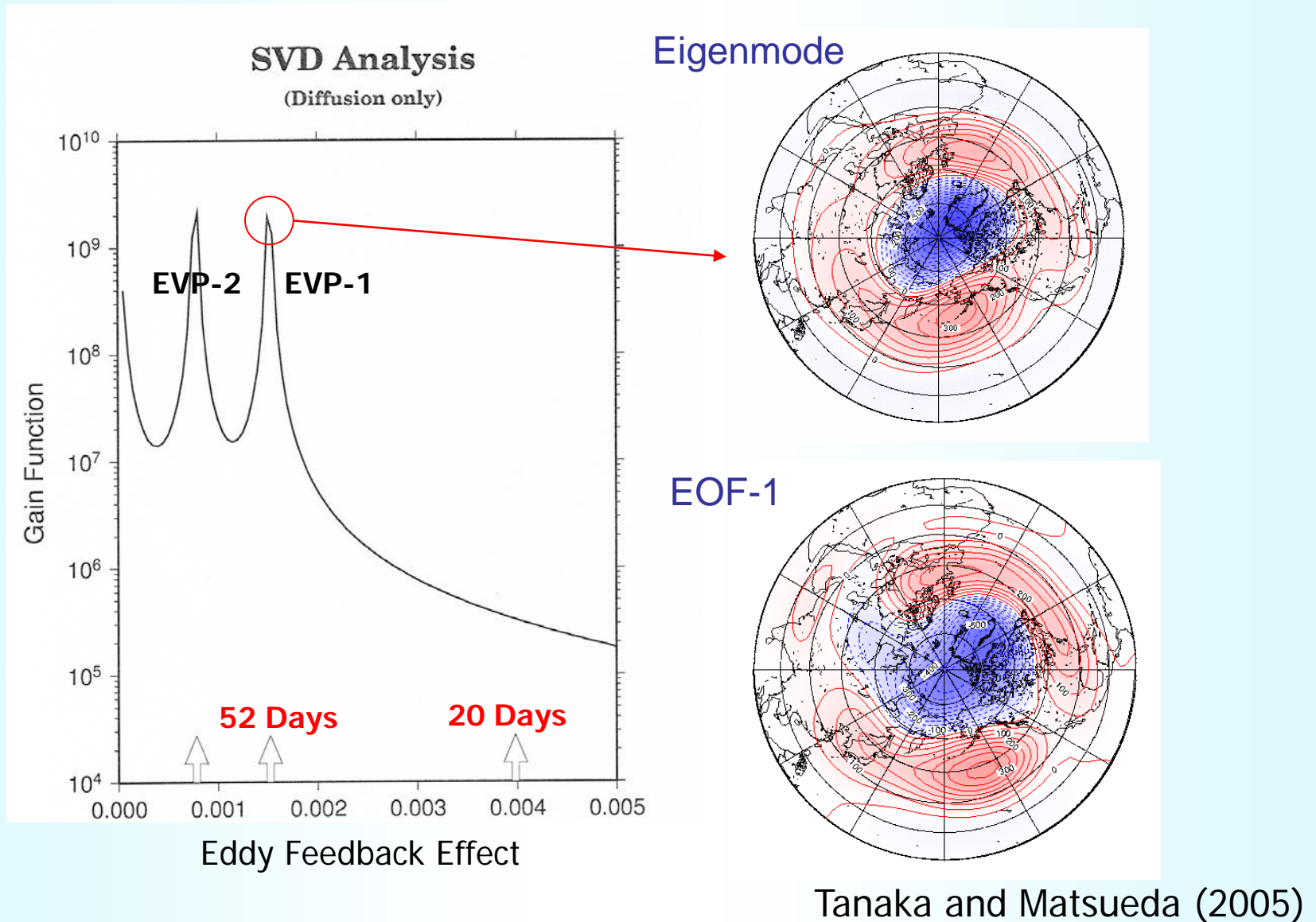
Barotropic S-Model

Barotropic Height

EOF-1 (16%)



# Singular Eigenmode Theory of AO

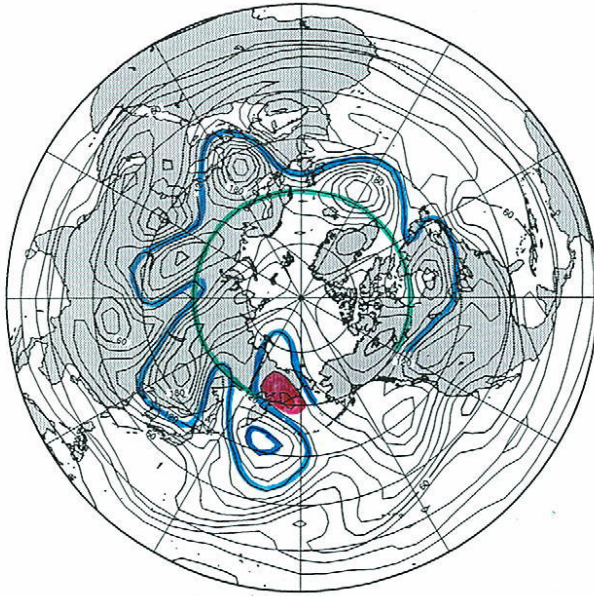




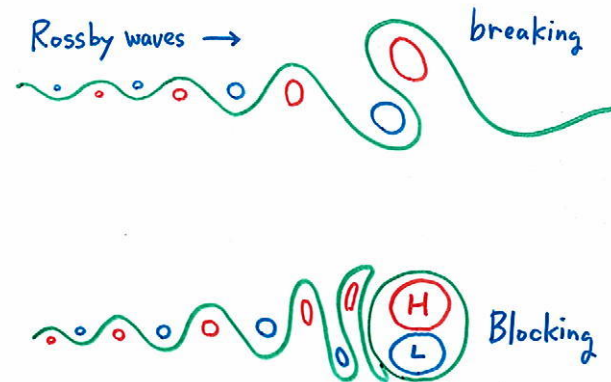


Potential Vorticity

Day 79



# Blocking formation by Rossby wave breaking



(Tanaka and Watarai 1999)

Blocking

Breaking Rossby Waves

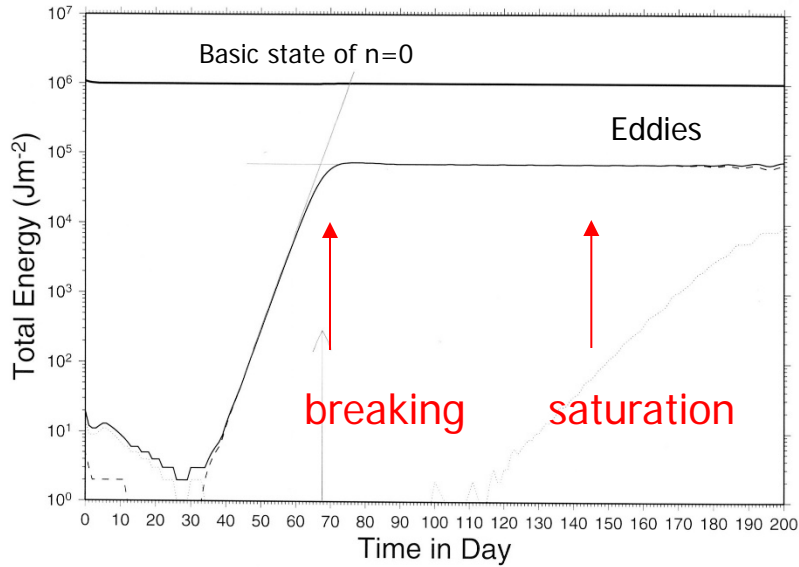


Tanaka (1998)



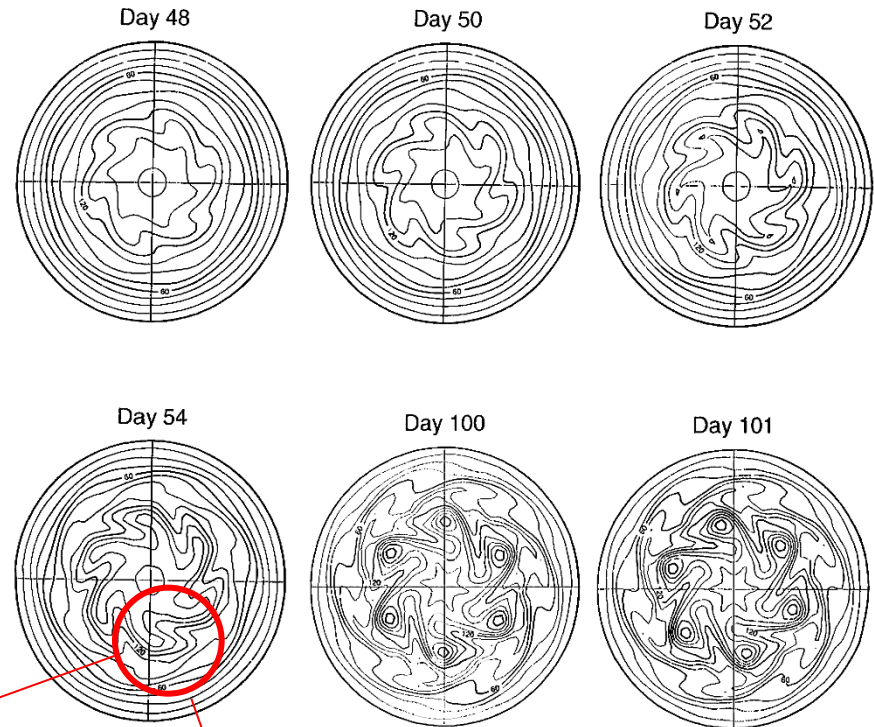
# Rossby wave breaking and saturation

## Time Series

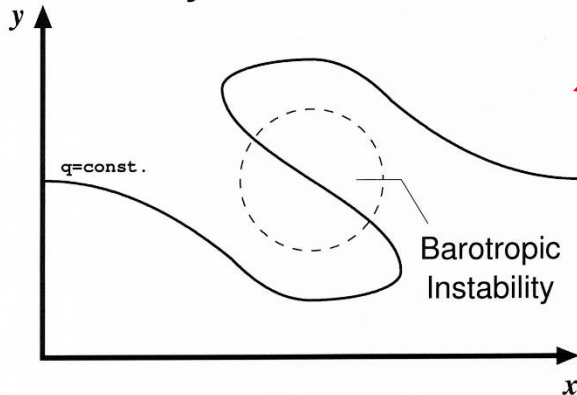


## Potential Vorticity

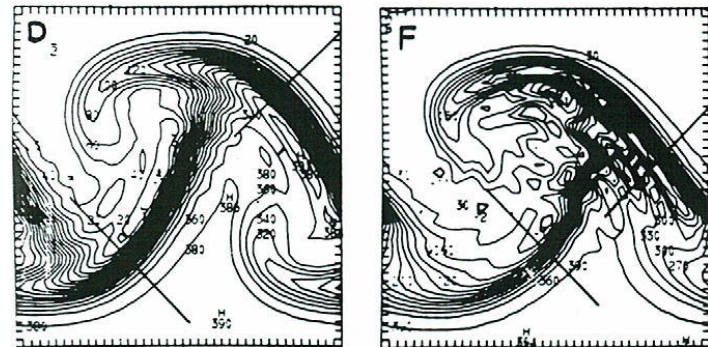
Wave-6 Model



## Rossby Wave



Tanaka and Watarai (1999)



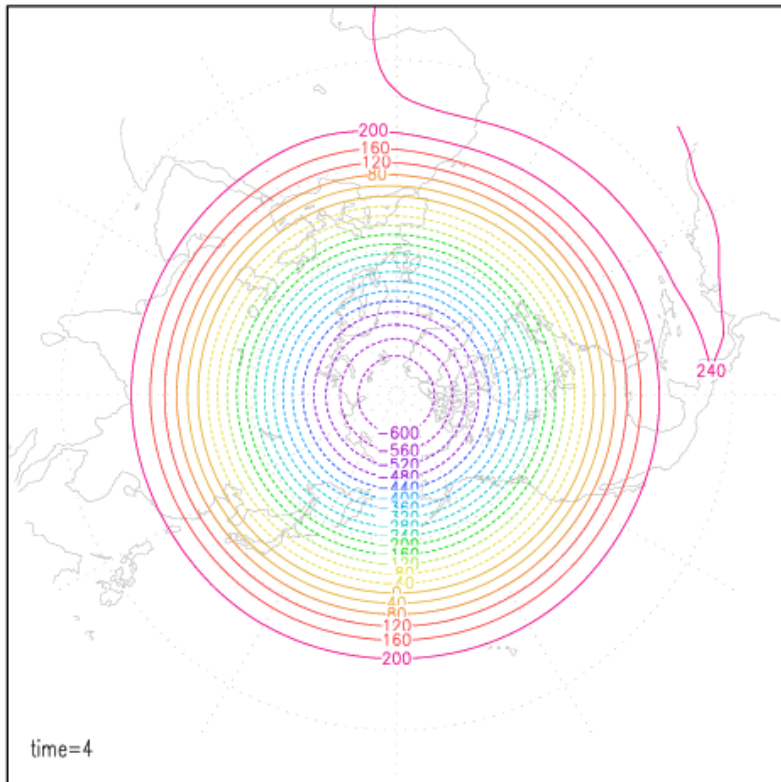
Mudrick (1974)

# Zonalization

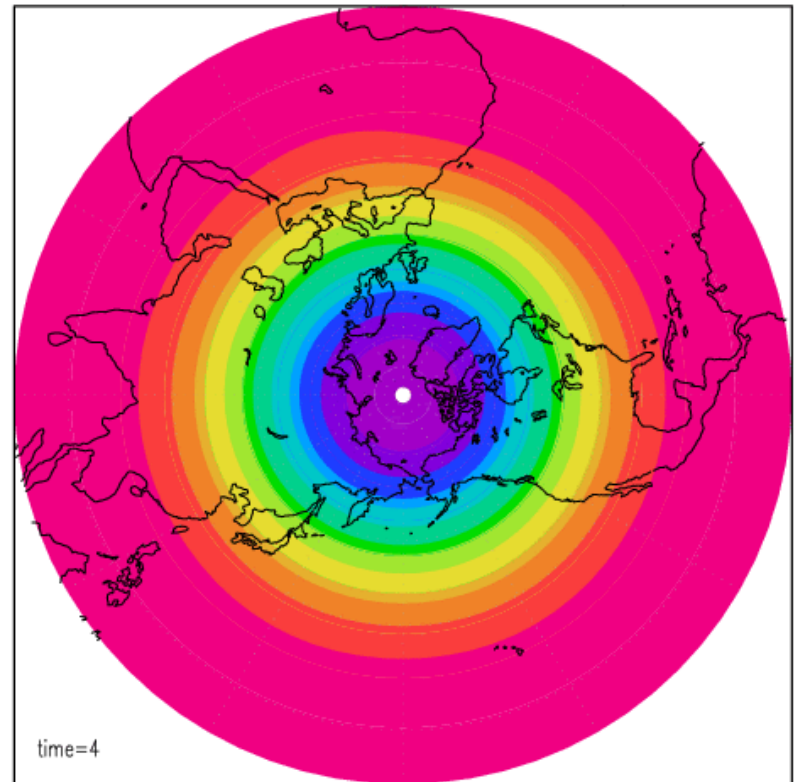
## Rossby wave breaking for $n=6$

Growthrate  $\times 1.7$

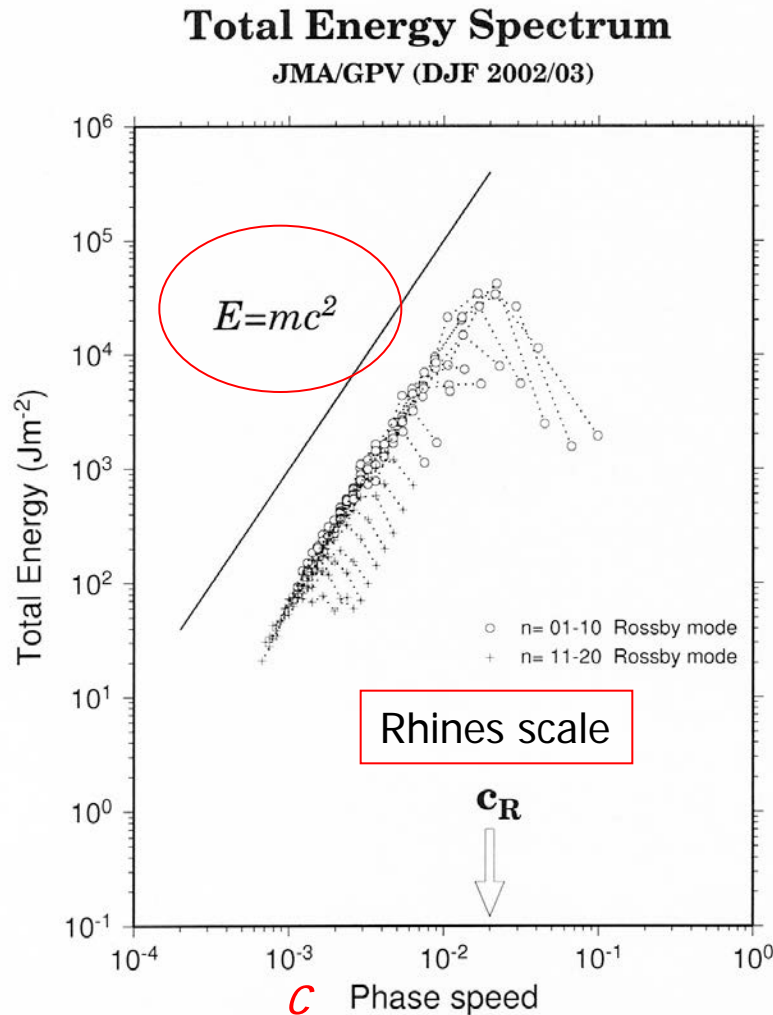
Barotropic Height  
Wavenumber 6



Barotropic Height  
Wavenumber 6



# 3D energy spectrum



By 3D normal mode expansion

$$\frac{dw_i}{d\tau} = -i\sigma_i w_i - i \sum_{jk} r_{ijk} w_j w_k + f_i$$

$$E_i = \frac{1}{2} p_s h_0 |w_i|^2$$

$$c_i = \sigma_i / n \quad \text{Phase speed}$$

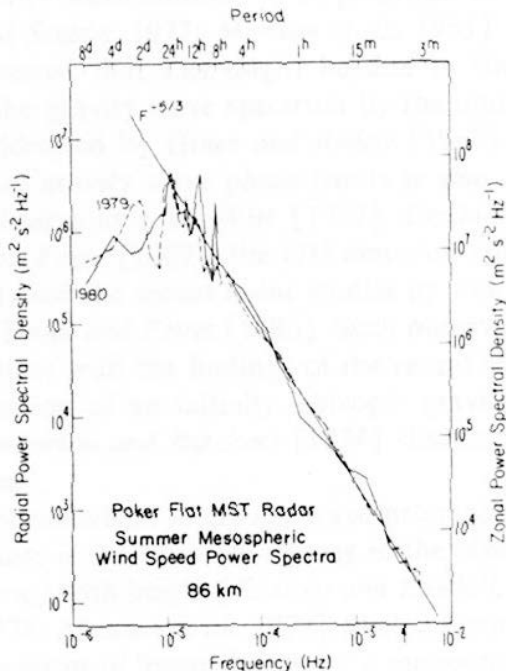
$$E = mc^2 \quad \text{?}$$

Tanaka et al. (2004 GRL)



# Saturation theory in gravity waves

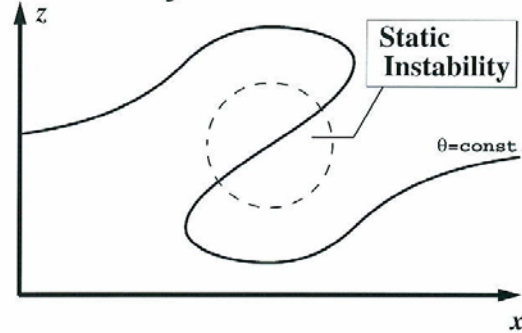
Saturation spectrum



**-3/5 power**

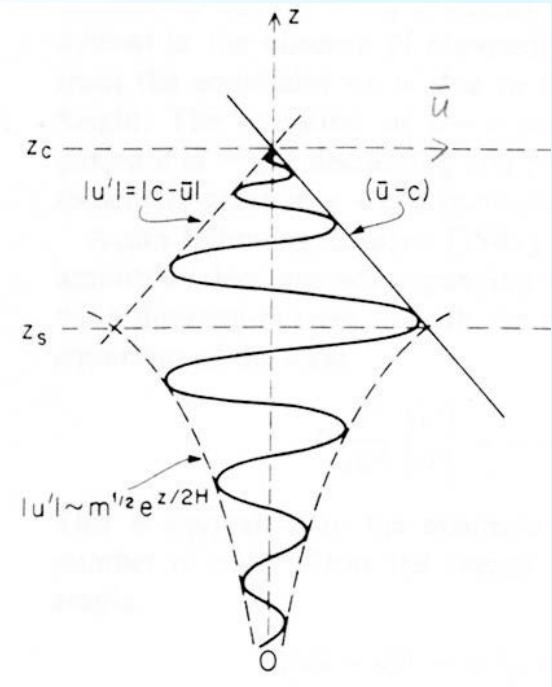
Breaking gravity waves

• Gravity Wave



$$\frac{\partial \theta}{\partial z} < 0$$

Breaking condition



Fritts (1984)

# Saturation theory in Rossby waves

$$\frac{\partial q}{\partial y} < 0, \quad q = \nabla^2 \psi + f$$

Tanaka and Watarai (1999)

PV in barotropic model

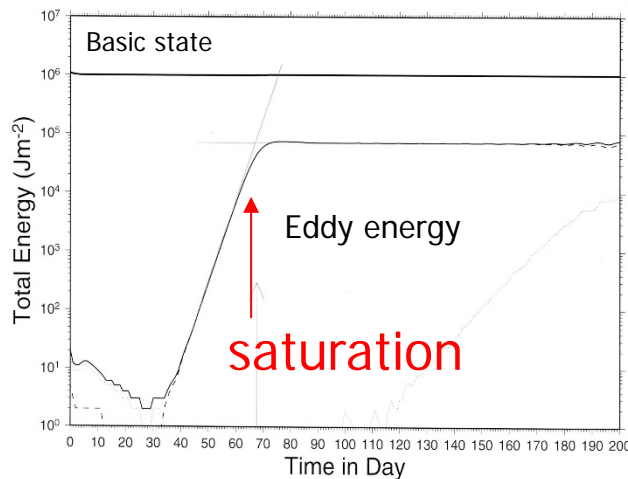
$$\frac{\partial}{\partial y} (\nabla^2 \psi + f) = -\nabla^2 u + \beta < 0$$

$$u < -\frac{\beta}{n^2 + l^2} = c$$

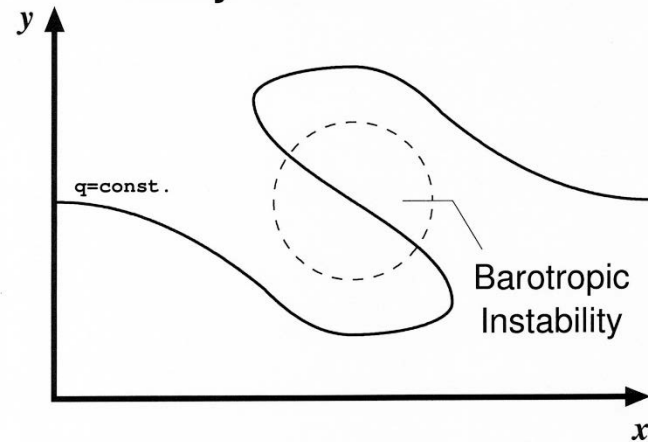
Breaking condition

$$\frac{\partial q}{\partial y} < 0$$

Time Series



Rossby Wave



# Saturation energy spectrum

$$u < -\frac{\beta}{n^2 + l^2} = c$$

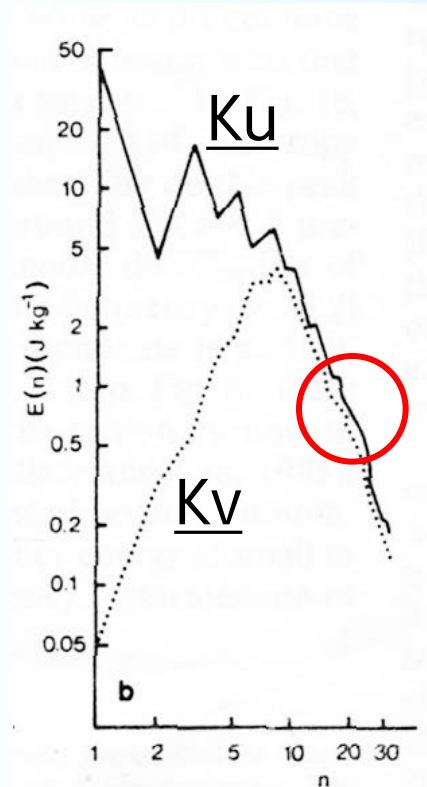
$$|u| \approx |v| \quad (\text{Tanaka and Kasahara 1992})$$

$$E = \frac{1}{g} \int_0^{p_s} \frac{1}{2} (u^2 + v^2) dp$$

$$= \frac{p_s}{g} c^2 = mc^2 \quad m = p_s / g$$

Mass for  
unit area

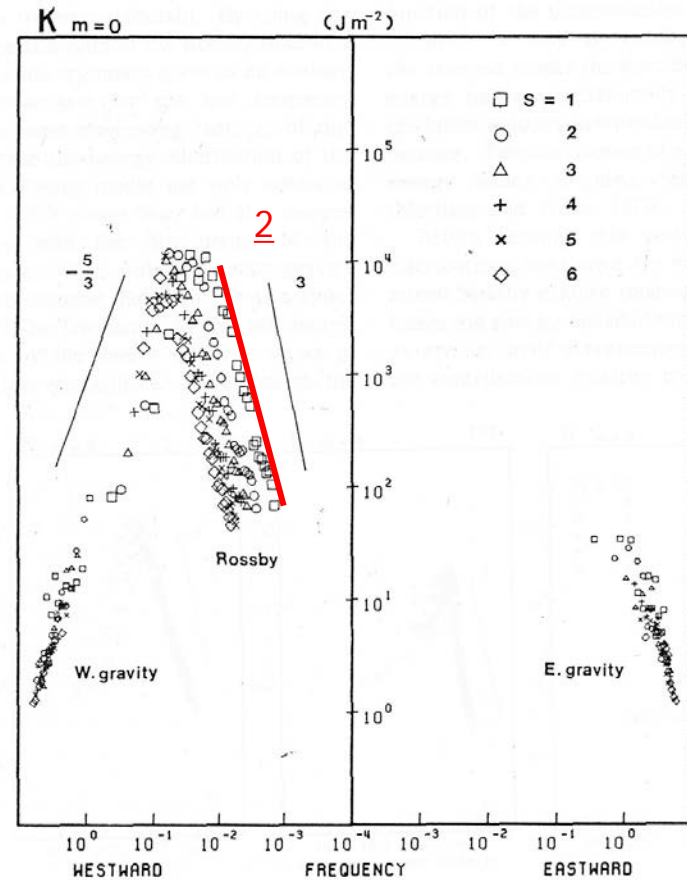
$$\frac{\partial q}{\partial y} < 0 \quad \Rightarrow \quad E = mc^2$$



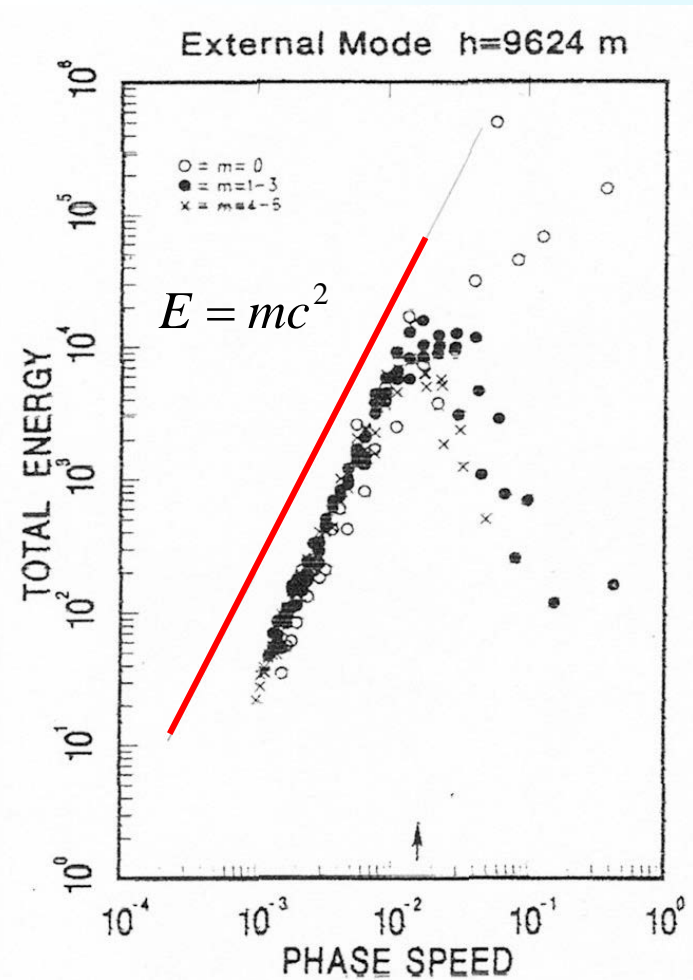
Shepherd (1987)

# Observed energy spectrum in c-domain

FGGE SOP1



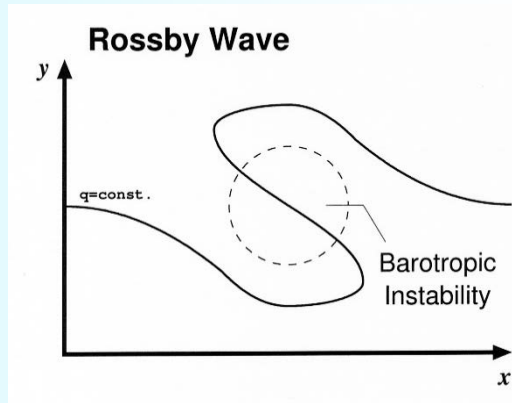
Tanaka (1985)



Tanaka and Kung (1988)

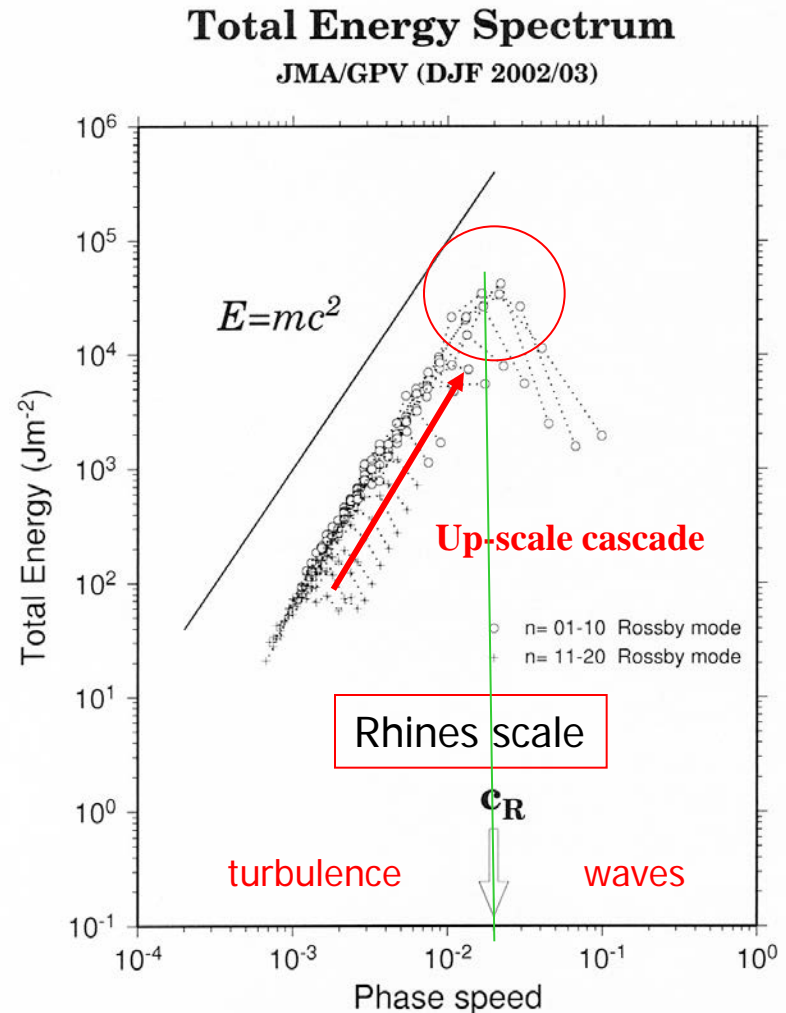
# Global energy spectrum of $E = mc^2$

(Tanaka et al. 2004 GRL)

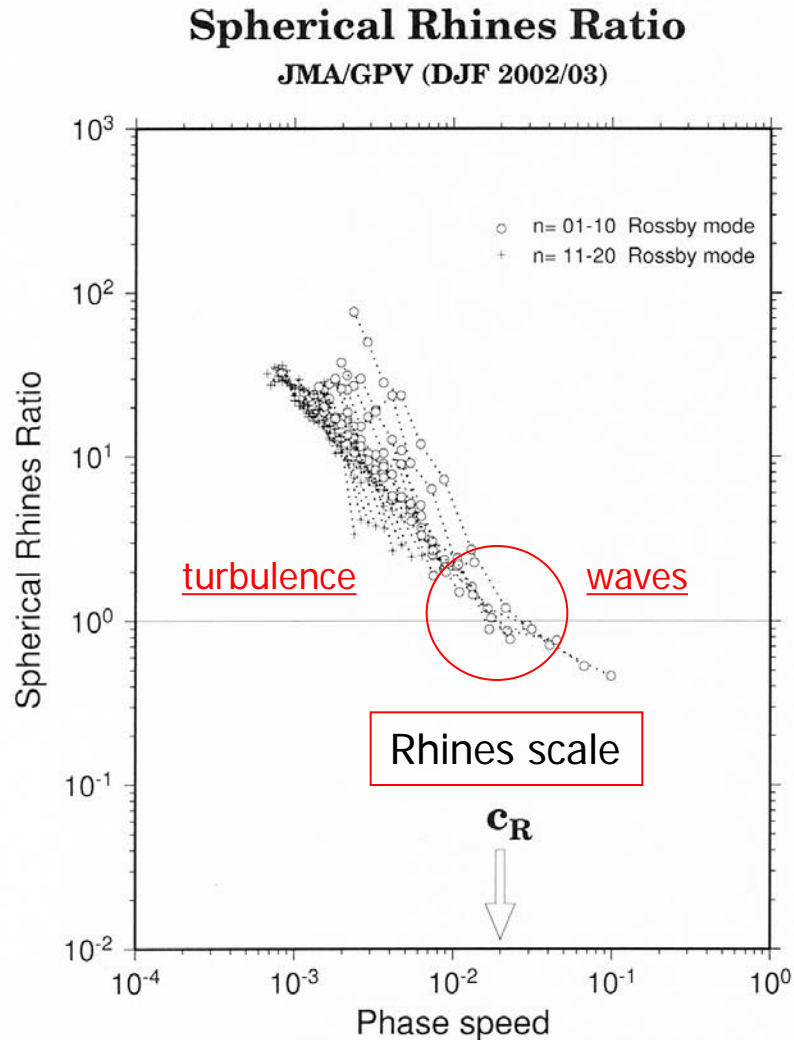


$$\frac{\partial q}{\partial y} < 0 \Rightarrow E = mc^2$$

$c$  Rossby phase speed  
 $m = p_s / g$  Mass of the air



# Rhines scaelae on a sphere



$$\frac{dw_i}{d\tau} = -i\sigma_i w_i - i \sum_{jk} r_{ijk} w_j w_k + f_i$$

$$R_i = \frac{|\sum r_{ijk} w_j w_k|}{|\sigma_i w_i|} \quad \text{Rhines ratio}$$

turbulence       $R_i > 1$

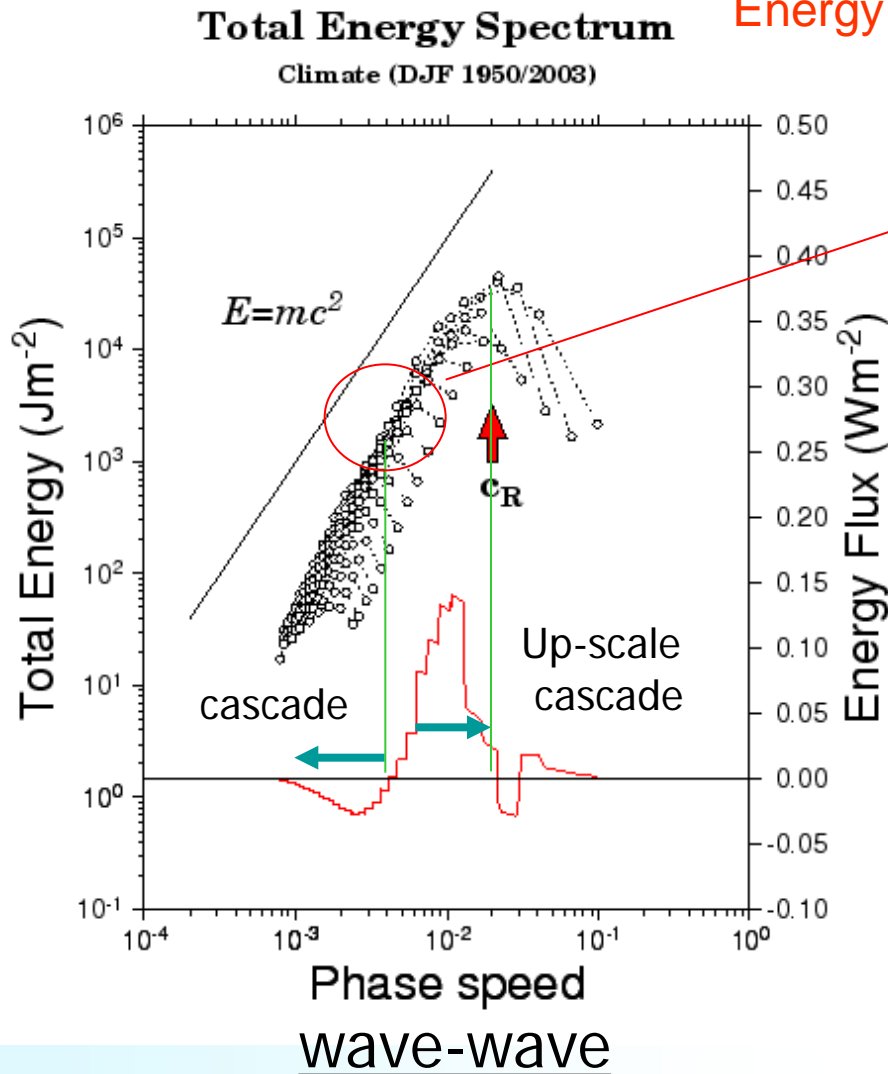
waves       $R_i < 1$

Rhines scale       $R_i = 1$

(Tanaka et al. 2004 GRL)

# Energy flux in $c$ -domain

Energy accumulates at phase speed 0



Source of energy

There is an energy source in the middle of the spectrum: inertial subrange theory fails.

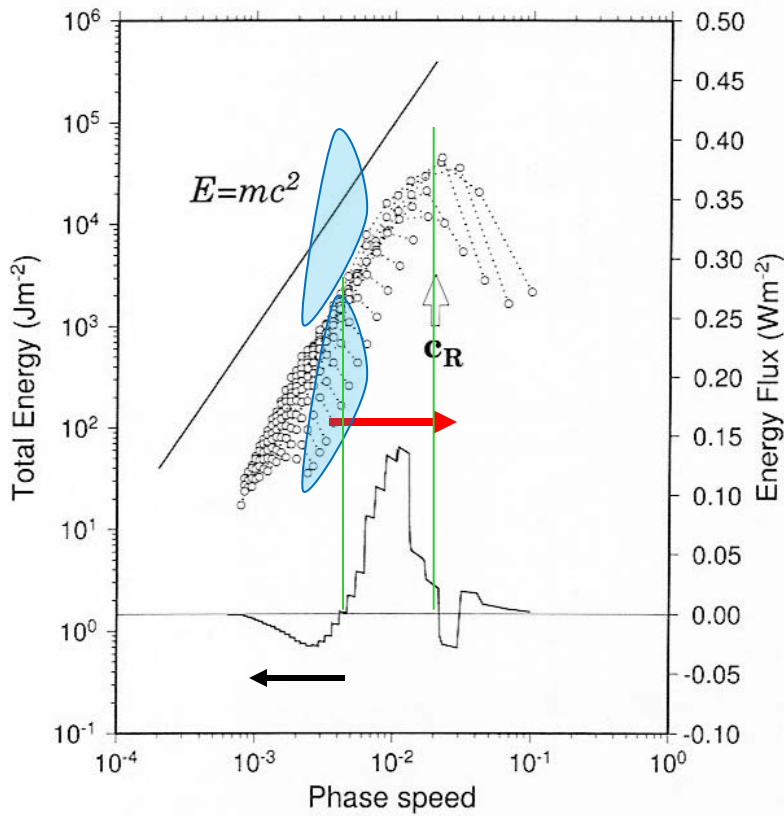


Theory of Rossby wave saturation

# Excitation of blocking and AO by up-scale energy cascade

## Total Energy Spectrum

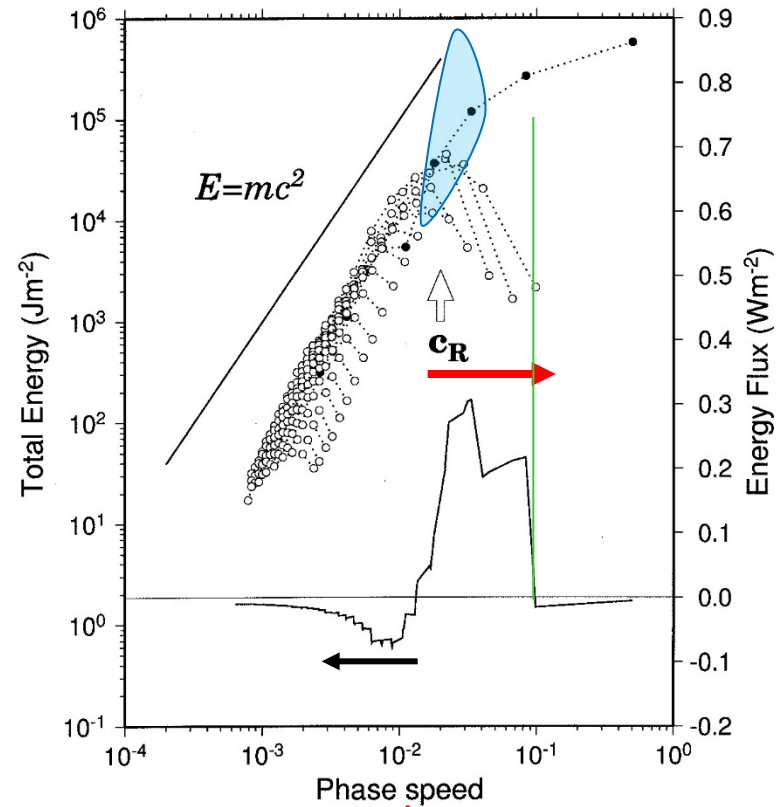
Climate (DJF 1950/2003)



wave-wave

## Zonal-Wave Interactions

Climate (DJF 1950/2003)



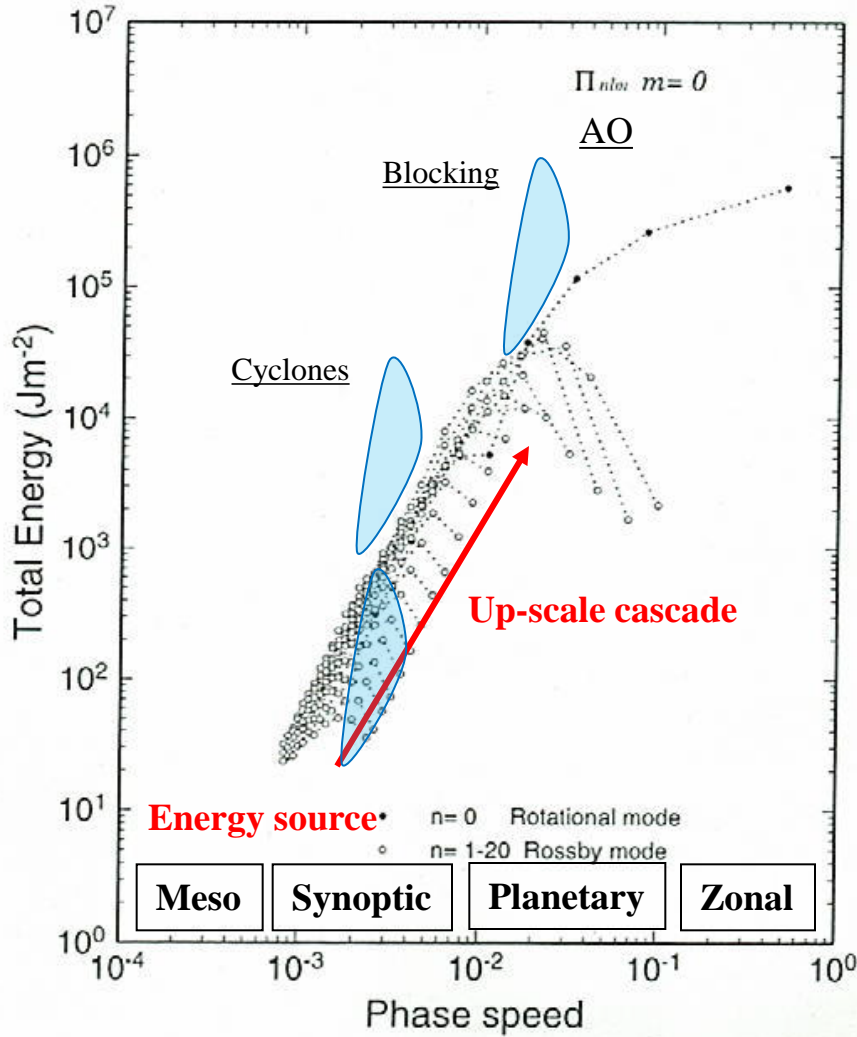
zonal-wave

(Tanaka and Terasaki 2004)

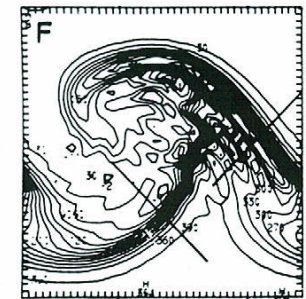
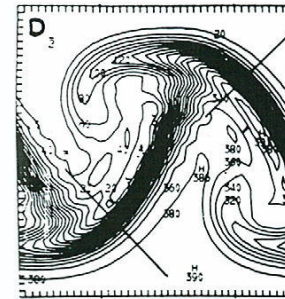
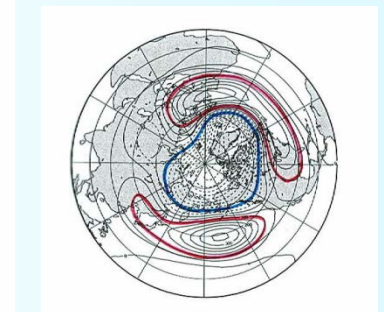
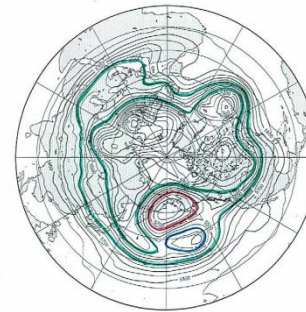


# Total Energy Spectrum

NCEP/NCAR DJF 1950-1999



# Low-frequency variability of the atmosphere





# Summary

- (1) Energy spectrum is examined in the 3D wavenumber domain using **phase speed  $c$** .
- (2) Energy spectrum of  $E=mc^2$  is obtained and explained by **Rossby wave saturation**
- (3) Up-scale energy cascade to Rhine's scale forms **blocking**
- (4) Further up-scale cascade to zonal energy forms the **Arctic Oscillation**



END

Thank you.

## Barotropic S-Model

## Model description

### ○ 3-D Spectral Model

$$\frac{dw_i}{d\tau} + i\sigma_i w_i = -i \sum_{jk} r_{ijk} w_j w_k + f_i, \quad i = 1, 2, 3, \dots, \quad (7)$$

where the symbols denote:  
 $w_i$  : expansion coefficient of  $U$   
 $f_i$  : expansion coefficient of  $F$   
 $\sigma_i$  : Laplace's tidal frequency  
 $r_{ijk}$  : nonlinear interaction coefficient  
 $\tau$  : dimensionless time

### Barotropic S-Model

### ○ Barotropic Spectral Model

$$\frac{dw_i}{d\tau} + i\sigma_i w_i = -i \sum_{jk} r_{ijk} w_j w_k + s_i, \quad i = 1, 2, 3, \dots, \quad (m = 0), \quad (8)$$

where the external forcing  $s_i$  includes barotropic-baroclinic interactions.

## ○ External Forcing

External forcing  $s_i$  is statistically obtained by observed data using the least square method:

$$s_i = \tilde{s}_i + \mathbf{A}_{ij}w_j + \mathbf{B}_{ij}w_j^* + \epsilon'_i. \quad (9)$$

where  $\tilde{s}_i$  is the climate of  $s_i$  and the matrices  $\mathbf{A}_{ij}$   $\mathbf{B}_{ij}$  are evaluated by minimizing  $\epsilon_i$ :

$$\mathbf{A}_{ij} = \overline{s'_i w_j^+}. \quad (10)$$

where  $s'_i = s_i - \tilde{s}_i$  and pseudo-inverse of  $w_j$  is

$$w_j^+ = w_k^H (\overline{w_k w_j^H})^{-1} \quad (11)$$

The matrix  $\mathbf{B}_{ij}$  is similarly obtained stepwise by minimizing the first residual  $\delta_i$

$$\mathbf{B}_{ij} = \overline{\delta_i w_j^{*+}}, \quad (12)$$

Finally, the external forcing is given by

$$s_i = \tilde{s}_i + \mathbf{A}_{ij}w_j + \mathbf{B}_{ij}w_j^* + (BC)_{ij}w_j \quad (13)$$

$$+ (DF)_{ij}w_j + (DZ)_{ij}w_j + (DE)_{ij}w_j. \quad (14)$$

Physical processes considered in the S-Model:

(BC): baroclinic instability

(DF): biharmonic diffusion

(DZ): zonal surface stress and

(DE): Ekman pumping.

# Barotropic forcing