MODES software and its application to the ERA Interim reanalyses

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Based on

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Outline

- Introduction
- Development of the normal-mode function (NMF) representation in the terrain-following coordinate
- Solution of NMF problem: vertical structure functions and horizontal structure functions
- Construction of the MODES software
- Some outputs of the MODES applied to the ERA Interim dataset
- Conclusions and Outlook

Normal mode function (NMF) representation

Balance: part of the circulation that is associated with the Rossby type (quasi-geostrophic, ROT) of solutions to the linearized primitive equations.

The unbalanced part projects onto the inertio-gravity solutions that propagate eastward (EIG) or westward (WIG).

Derivation of NMF following Kasahara&Puri, 1981

²¹⁰ σ level and R the gas constant of air. It is convenient to in-

Vertical coordinate *σ=p/p_s*

A new geopotential variable $P = \Phi + RT_0 \ln(p_s)$ $h' = P/g$ \mathbf{S} postantial variable \mathbf{D} A new geopotential variable $\mathbf{F} = \mathbf{\Psi} + \mathbf{\Lambda} \mathbf{I}_0 \ln(P_s)$, $n' = P / g$

System of linearized equations describing osclillations (u',v',h') superimposed on a basic state of rest with T₀(σ): (uli
List $\frac{1}{2}$) superimposed on a basic state of rest with temperature on a basic state of rest with temper-
In the most with temper-basic state of rest with temperature of rest with temperature of the most with the mos 22 System of linearized equations describing osclillations (μ' ν' h') rast with $T(\sigma)$ is obtained as a continuous of the continuous of the continuous of the continuous set of the c $\frac{1}{\sqrt{2}}$

$$
\frac{\partial u'}{\partial t} - 2\Omega v' \sin(\varphi) = -\frac{g}{a \cos(\varphi)} \frac{\partial h'}{\partial \lambda},
$$

$$
\frac{\partial v'}{\partial t} + 2\Omega u' \sin(\varphi) = -\frac{g}{a} \frac{\partial h'}{\partial \varphi},
$$

$$
\frac{\partial}{\partial t} \left[\frac{\partial}{\partial \sigma} \left(\frac{g \sigma}{R \Gamma_0} \frac{\partial h'}{\partial \sigma} \right) \right] - \nabla \cdot \mathbf{V}' = 0.
$$

The static stability parameter:
$$
\Gamma_0 = \frac{\kappa T_0}{\sigma} - \frac{dT_0}{d\sigma}
$$

Derivation of NMF following Kasahara&Puri, 1981 represented as the product of 2D motions and the vertical product of 2D motions and the vertical product of 2D $\frac{1}{1000}$ in the section. $\frac{1}{2}$ N. $\overline{}$ structure function $\overline{}$ is a solution of the solution of the theorem of the theorem of the theorem of the solution of the

Solutions in terms of horizontal and vertical dependencies: VSE written in the dimensionless form as d <u>" σ</u> dG

$$
[u', v', h']^{\mathrm{T}}(\lambda, \varphi, \sigma, t) = [u, v, h]^{\mathrm{T}}(\lambda, \varphi, t) \times G(\sigma)
$$

An equation for the vertical structure and a set of equations for 2D horizontal motions identical in form to the global shallow water equations. 3D and 2D motions are connected by particular values of a separation parameter – "**equivalent depth" D**. and 2D motions are connected by particular values of a separation
、motor、"**equivalent denth" D** \mathcal{L} is the government of the \mathcal{L} . of dotate and a cotter equations for ED
in form to the alobal shallow water equistions $T_{\rm eff}$ structure function σ structure function σ ucture and a set of equations for 2D dσ S dσ with to the giodal shallow water equations. f'' stratification and a typical profile of S is shown in S is shown in 270 where S(σ) where S(σ) = RF0
(gH∗). Here, H∗ is a scaling constant constant constant constant constant constant constant constant constant form to the global shallow water equations. $\frac{1}{2}$ for stratification and a typical properties of $\frac{1}{2}$ separation Fig. 1 in the next section. The extendion is dependent in the equivalent is denoted by D_i ehri n

The vertical structure equation:
\n
$$
\frac{d}{d\sigma} \left(\frac{\sigma}{S} \frac{dG}{d\sigma} \right) + \frac{H_*}{D} G = 0
$$
\nand its boundary conditions
\n
$$
\frac{dG}{d\sigma} + rG = 0 \text{ where } r = \frac{\Gamma_o}{T_o} \text{ at the bottom } \sigma = 1
$$
\n
$$
\sigma \frac{dG}{d\sigma} = 0 \text{ at the model top } \sigma = \sigma_T
$$
\n\nprovide solutions for $G(\sigma)$ which are orthogonal
\n
$$
\int_{\sigma_T}^{1} G_i(\sigma) G_j(\sigma) d\sigma = \delta_{ij},
$$

Vertical expansion of discrete data onto NMF

 $\mathbf{X}(\lambda, \varphi, \sigma) = (u, v, h)^T$ An input data vector **X** is defined on the horizontal regular Gaussian grid and vertical sigma levels at time *t*:

Projection of a single data point on j-th sigma level is performed on the precomputed vertical structure functions **G**, the horizontal Hough vector functions in the meridional direction and waves in the longitudinal direction:

$$
\mathbf{X}(\lambda, \varphi, \sigma_j) = \sum_{m=1}^{M} \mathbf{S}_m \mathbf{X}_m(\lambda, \varphi) \cdot G_m(j)
$$
 (1)

$$
\mathbf{X}_{m}(\lambda,\varphi) = \left(\tilde{u}_{m},\tilde{v}_{m},\tilde{h}_{m}\right)^{T} = \left(\frac{u_{m}}{\sqrt{gD_{m}}},\frac{v_{m}}{\sqrt{gD_{m}}},\frac{h_{m}}{D_{m}}\right)^{T}
$$

Normalization factors define diagonal elements of 3x3 matrix normalization matrix **S**

The vector \mathbf{X}_m is obtained by the reverse transform of (1):

$$
\mathbf{X}_{m}(\lambda,\varphi)=\mathbf{S}_{m}^{-1}\sum_{j=1}^{J}\left(u,v,h\right)_{j}^{T}G_{m}(j)
$$
 (2)

Solutions of the vertical structure function

Example of ERA Interim dataset: L60

Example of ERA Interim dataset: L60 versus L21 (standard p levels)

Derivation of NMF following Kasahara&Puri, 1981

The horizontal structure equation

$$
\frac{\partial}{\partial t} \mathbf{X}_m + \mathbf{L} \mathbf{X}_m = 0
$$

describe 2D motions for vector **X***m*(λ,φ) every equivalent depth D*m*:

$$
\mathbf{X}_{m}(\lambda,\varphi) = \left(\tilde{u}_{m},\tilde{v}_{m},\tilde{h}_{m}\right)^{T} = \left(\frac{u_{m}}{\sqrt{gD_{m}}},\frac{v_{m}}{\sqrt{gD_{m}}},\frac{h_{m}}{D_{m}}\right)^{T}
$$

L is the linear differential matrix operator ${\bf L} =$! ! | | | | ! 0 $-\sin(\varphi)$ $\frac{\gamma}{\cos(\varphi)}$ ∂ ∂ λ $\sin(\varphi)$ 0 $\gamma \frac{\partial}{\partial \varphi}$ γ $\cos(\varphi)$ <u>∂</u> ∂ λ γ $\cos(\varphi)$ $\frac{\partial}{\partial \varphi} [\cos(\varphi)(\theta)]$ 0

Solution \mathbf{X}_m is harmonic in time: $\left\| \mathbf{X}_m (\lambda, \varphi, t) \!=\! \mathbf{H}^{\kappa}_n (\lambda, \varphi) e^{-i \mathbf{V}_n t} \right\|$

nic in time:
$$
\mathbf{X}_{m}(\lambda, \varphi, \tilde{t}) = \mathbf{H}_{n}^{k}(\lambda, \varphi)e^{-i\mathbf{v}_{n}^{k}\tilde{t}}
$$

Since (18) is a linear system with respect to time, the so-

! ! | | | ! !

 \tilde{t}

speed of Earth: With global inner product defined as $\left\langle \mathbf{X}_{\text{l}},\mathbf{X}_{\text{m}}\right\rangle$ = $\frac{1}{2^{j}}$ $2a_{0-1}$
2a⁰ a_{-1} λ 2π $\left(\tilde{u}_l \tilde{u}_m^* + \tilde{v}_l \tilde{v}_m^* + \tilde{h}_l \tilde{h}_m^* \right)$ -1 1 \int 0 2π $\int\int\left(\tilde{u}_{\iota}\tilde{u}_{\iota}^*+\tilde{\nu}_{\iota}\tilde{v}_{\iota\iota}^*+\tilde{h}_{\iota}\tilde{h}_{\iota\iota}^*\right)d\mu d\lambda$ functions **H** are expressed as $\mathbf{H}_{n}^{k}(\lambda, \varphi)$ = $\Theta_{n}^{k}e^{ik\lambda}$ Hough harmonics

Horizontal expansion of discrete data onto NMF

The horizontal coefficient vector **X***m* for a given vertical mode is projected onto the Hough harmonics **H**ⁿ k(λ,φ,m) as

$$
\mathbf{X}_{m}(\lambda,\varphi)=\sum_{n=1}^{R}\sum_{k=-K}^{K}\chi_{n}^{k}(m)\mathbf{H}_{n}^{k}(\lambda,\varphi,m)
$$
 (3)

The subscript *n* indicates all meridional modes including rotational (ROT), and eastward and westward propagating inertio-gravity (EIG and WIG, respectively) modes

The scalar complex coefficients χ are obtained as

$$
\chi_n^k(m) = \frac{1}{2\pi} \int_{0}^{2\pi} \int_{-1}^{1} \left(\tilde{u}_m, \tilde{v}_m, \tilde{h}_m\right)^T \left[\mathbf{H}_{n'}^{k'}\right]^* d\mu d\lambda \tag{4}
$$

Here, $\mu = \sin(\phi)$.

Equations (3-4) are the horizontal transform pair.

Equations (1-2) represent the vertical transform pair.

Two kinds of Hough functions

Frequencies of spherical normal modes for different equivalent depths

NMF expansion: horizontal expansion functions

HSFs are pre-computed for a given number of vertical modes, M For every m=1,…,M, i.e. for every D_m

Meridional structure for Hough functions is computed for a range of the zonal wavenumbers K, k=-K,..,0,...,K and a range of meridional modes for the balanced, N_{ROT} , a range of EIG, N_{FIG} , and a range of WIG, N_{WIG} , modes.

 $R=N_{ROT} + N_{FIG} + N_{WIG}$

Expansion of 3D data onto NMF: MODES software

 $FC = gfortran$ FFLAGS=-03-I\$(SHAREDIR)-I\$(HOUGHDIR)-I\$(NORMALDIR)-I\$ (VSFDIR)-L\$(LIBDIR)-fopenmp-I\$(GRIB_API_INCLUDE) $LD = \frac{1}{2}$ (FC) $LFLAGS = -O3$ -fopenmp

Makefile and namelist controlled

Required libraries for the input data in grib and netcdf format

Input data on the Gaussian grid and model levels (sigma or hybrid)

Uses LAPACK or equivalent

Preparation steps requires computation of the stability profile

Five executables which are run in subsequent steps:

- Preparation of the horizontal grid
- Computation of the vertical structure functions
- Computation of the horiozntal structure functions
- Projection
- Filtering of selected modes to physical space

Expansion of 3D data onto NMF: MODES software

/

NMF software structure: grid and vsf namelists

```
& gaussian
      N = 256,
       gauss_fname = 'gauss256.data', 
 / 
& vsfcalc_cnf 
    stab fname = 'stability L60.data',
    vgrid fname = 'sigma levels L60.data',
    vsf fname = 'vsf L60.data',
     equiheight_fname = 'equivalent_height_L60.data', 
    num\_vmode = 60,mp = 60,hstd = 8000.0d= 0,
    \text{suffix} = 288.0d0,given stability = .true.,
    ocheck = true./
```
NMF software structure: hsf namelists

 $\overline{1}$

```
&houghcalc_cnf
```
/

```
szw = 0.
ezw = 200,maxl = 70my = 256freq_fname = 'freq.data',
ocheck = true.
```

```
& output
   output gmt = false.
   of name_gmt = 'hough_gmt',
   of name\_bin = 'hough',bin combine ='zonal',
```

```
& meridional_grid 
  ygrid_fname = 'gauss256.data', 
/
```

```
& vsf_cnf
   equiheight_fname = 'equivalent_height.data', 
  num vmode = 43,
/
```
NMF software structure: projection namelists

ifile_grib_head (1) = 'erai_N128_'

/

Energy product

Starting from the linearized system (u',v',h') and its solutions by using the variable separation, the partition of total energy into the kinetic and available potential energy for every vertical mode is written as :

$$
\frac{\partial}{\partial t} \int_{0}^{2\pi} \int_{-1}^{1} \sum_{m} \frac{1}{2} \left(u_m^2 + v_m^2 + \frac{g}{D_m} h_m^2 \right) a^2 d\mu d\lambda = 0
$$

Global energy product of the m-th vertical mode defined as

$$
I_m = \frac{1}{2} g D_m \sum_{n=1}^{R} \sum_{k=-K}^{K} \chi_n^k(m) \left[\chi_n^k(m) \right]^*
$$

is equivalent to

$$
I_m = \frac{1}{2} g D_m \int_0^{2\pi} \int_{-1}^1 \left(\tilde{u}_m^2 + \tilde{v}_m^2 + \tilde{h}_m^2 \right) d\lambda d\mu = \int_0^{2\pi} \int_{-1}^1 (K_m + P_m) d\lambda d\mu
$$

NMF representation of ERA Interim dataset

NMF representation of ERA Interim dataset

Scale-dependent climatological distribution of atmospheric total energy

Percentage of ROT/IG energy in each scale

Based on 30-year period between 1980 and 2009

12 UTC time

Up to 10% of wave energy in IG modes \Leftrightarrow 1/3 of circulation

Application of the normal mode representation

Modal-space diagnosis of the vertical energy propagation by the Kelvin waves

Evolution of the zonal wavenumber *k*=1 Kelvin wave in July 2007, filtered for periods shorter than 36 hours. The best agreement between the datasets exists for the Kelvin wave.

Application of the normal mode representation

Modal-space diagnosis of the vertical energy propagation by the Kelvin waves

component in the Kelvin wave at subsequent days (left) shows the downward phase propagation (upward energy propagation).

The difference in the depth of the atmosphere in DART/CAM and NCEP/NCAR on one hand and ECMWF and NCEP on the other appears to be one reason for different propagation properties.

(Zagar et al., 2009)

NMF software structure: filtering namelists

&normal_cnf_inverse

/

 ℓ , filter cnf s = 1, $e = 70$, $s = 1$, e = 70, $= 2$, $= 70,$ k mode_s = 100, k mode $_e$ = 85, $vmode_s = 410,$ $e = 70$,

NMF representation of ERA Interim dataset

NMF representation of ERA Interim dataset

Jan8

Jul8

Real-time modal view of ECMWF forecast http://meteo.fmf.uni-lj.si/MODES

Modal representation of the MJO patterns, we perform a linear regression between the MJO index Mi(t) and modal ex- $\frac{1}{2}$ as follows. First we compute the mean values of the time $\frac{1}{2}$

Regression between the MJO index and ERA Interim series of the expansion coefficients and the MJO index and denote the M

$$
\mathcal{R}_n^k(m,\tau) = \frac{1}{N-1} \frac{\sum_{t=1}^N \left[\left(\chi_n^k(m,t) - \overline{\chi_n^k(m)} \right) \left(\mathcal{M}_i(t,\tau) - \overline{\mathcal{M}_i} \right) \right]}{Var\left(\overline{\mathcal{M}_i} \right)}
$$

Modal representation of MJO

Quantification of the role of various modes in MJO and its teleconnections

- Representation of the global circulation and mass fields in global NWP and climate models in terms of the normal modes of the Navier-Stokes equations offers an alternative and physically attractive approach to the diagnostic of some properties of the models
- A new software for the representation of global atmospheric energy, MODES, has been developed under the ERC funding and it is available to the atmospheric research community
- MODES provides possibility to study circulation changes in the coupled climate models in relation to the (un)balanced dynamic. Unbalanced circulation is climatologically small and difficult to diagnose, but it is critical for understanding atmospheric variability

Thank you very

much for your

attention!

MODES workshop

On normal-mode function theory and applications to observations, numerical weather prediction models and climate research

Boulder, CO • 26-28 August 2015 **Q** NCAR, Mesa Lab

Topics of interest:

- Normal-mode function (NMF) theory Lessons learnt from the initialization of NWP models based on NMFs · Diagnosis of global atmospheric dynamics, NWP and climate models using NMFs · General session on modes of atmospheric variability

Bennert Machenhauer Ron Errico, John Boyd Robert Dickinson **Grant Branstator** Jose Castanheira Nils Gustafsson Joan Alexander Rolando Garcia Taroh Matsuno Ronald Madden Pedro Silva Dias Wayne Schubert Matthew Wheeler Christian Franzke Andrew Staniforth Daryl Kleist, Dave Parrish George Kiladis, Erland Kallen

Invited speakers:

Information and registration: https://www2.cgd.ucar.edu /sections/amp/events/20150826

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