# MODES software and its application to the ERA Interim reanalyses

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### Outline

- Introduction
- Development of the normal-mode function (NMF) representation in the terrain-following coordinate
- Solution of NMF problem: vertical structure functions and horizontal structure functions
- Construction of the MODES software
- Some outputs of the MODES applied to the ERA Interim dataset
- Conclusions and Outlook

### Normal mode function (NMF) representation



**Balance:** part of the circulation that is associated with the Rossby type (quasi-geostrophic, ROT) of solutions to the linearized primitive equations.

The unbalanced part projects onto the inertio-gravity solutions that propagate eastward (EIG) or westward (WIG).

### Derivation of NMF following Kasahara&Puri, 1981

Vertical coordinate  $\sigma = p/p_s$ 

A new geopotential variable  $P = \Phi + RT_0 \ln(p_s)$  h' = P / g

System of linearized equations describing osclillations (u',v',h') superimposed on a basic state of rest with  $T_0(\sigma)$ :

$$\begin{aligned} \frac{\partial u'}{\partial t} &- 2\Omega v' \sin(\varphi) = -\frac{g}{a \cos(\varphi)} \frac{\partial h'}{\partial \lambda}, \\ \frac{\partial v'}{\partial t} &+ 2\Omega u' \sin(\varphi) = -\frac{g}{a} \frac{\partial h'}{\partial \varphi}, \\ \frac{\partial}{\partial t} \left[ \frac{\partial}{\partial \sigma} \left( \frac{g\sigma}{R\Gamma_0} \frac{\partial h'}{\partial \sigma} \right) \right] - \nabla \cdot \mathbf{V}' = 0. \end{aligned}$$

The static stability parameter: 
$$\Gamma_0 = \frac{\kappa T_0}{\sigma} - \frac{dT_0}{d\sigma}$$

### Derivation of NMF following Kasahara&Puri, 1981

Solutions in terms of horizontal and vertical dependencies:

$$[u', v', h']^{\mathrm{T}}(\lambda, \varphi, \sigma, t) = [u, v, h]^{\mathrm{T}}(\lambda, \varphi, t) \times G(\sigma)$$

An equation for the vertical structure and a set of equations for 2D horizontal motions identical in form to the global shallow water equations. 3D and 2D motions are connected by particular values of a separation parameter – "equivalent depth" D.

The vertical structure equation:  

$$\frac{d}{d\sigma} \left( \frac{\sigma}{S} \frac{dG}{d\sigma} \right) + \frac{H_*}{D} G = 0$$
and its boundary conditions  

$$\frac{dG}{d\sigma} + rG = 0 \text{ where } r = \frac{\Gamma_o}{T_o} \text{ at the bottom } \sigma = 1$$

$$\sigma \frac{dG}{d\sigma} = 0 \text{ at the model top } \sigma = \sigma_T$$
provide solutions for G( $\sigma$ ) which are orthogonal  $\int_{\sigma_T}^1 G_i(\sigma)G_j(\sigma)d\sigma = \delta_{ij}$ ,

### Vertical expansion of discrete data onto NMF

An input data vector **X** is defined on the horizontal  $\mathbf{X}(\lambda,\varphi,\sigma) = (u,v,h)^{T}$ regular Gaussian grid and vertical sigma levels at time *t*:

Projection of a single data point on j-th sigma level is performed on the precomputed vertical structure functions **G**, the horizontal Hough vector functions in the meridional direction and waves in the longitudinal direction:

$$\mathbf{X}(\lambda,\varphi,\sigma_j) = \sum_{m=1}^{M} \mathbf{S}_m \mathbf{X}_m(\lambda,\varphi) \cdot G_m(j)$$
(1)

 $\mathbf{X}_{m}(\lambda,\varphi) = \left(\tilde{u}_{m},\tilde{v}_{m},\tilde{h}_{m}\right)^{T} = \left(\frac{u_{m}}{\sqrt{gD_{m}}},\frac{v_{m}}{\sqrt{gD_{m}}},\frac{h_{m}}{D_{m}}\right)^{T}$  Normalization factors define diagonal elements of 3x3 matrix normalization matrix **S** 

The vector  $\mathbf{X}_m$  is obtained by the reverse transform of (1):

$$\mathbf{X}_{m}(\lambda,\varphi) = \mathbf{S}_{m}^{-1} \sum_{j=1}^{J} \left(u,v,h\right)_{j}^{T} G_{m}(j)$$
(2)

### Solutions of the vertical structure function

Example of ERA Interim dataset: L60





Example of ERA Interim dataset: L60 versus L21 (standard p levels)



### Derivation of NMF following Kasahara&Puri, 1981

The horizontal structure equation

$$\frac{\partial}{\partial t}\mathbf{X}_m + \mathbf{L}\mathbf{X}_m = 0$$

describe 2D motions for vector  $\mathbf{X}_m(\lambda, \varphi)$  every equivalent depth  $D_m$ :

$$\mathbf{X}_{m}(\lambda,\varphi) = \left(\tilde{u}_{m}, \tilde{v}_{m}, \tilde{h}_{m}\right)^{T} = \left(\frac{u_{m}}{\sqrt{gD_{m}}}, \frac{v_{m}}{\sqrt{gD_{m}}}, \frac{h_{m}}{D_{m}}\right)^{T}$$

 $\begin{array}{c|c} \mathbf{L} \text{ is the linear} \\ \text{differential matrix} \\ \text{operator} \end{array} \begin{array}{c|c} \mathbf{L} = \left| \begin{array}{c} 0 & -\sin(\varphi) & \frac{\gamma}{\cos(\varphi)} \frac{\partial}{\partial \lambda} \\ \sin(\varphi) & 0 & \gamma \frac{\partial}{\partial \varphi} \\ \frac{\gamma}{\cos(\varphi)} \frac{\partial}{\partial \lambda} & \frac{\gamma}{\cos(\varphi)} \frac{\partial}{\partial \varphi} [\cos(\varphi)()] & 0 \end{array} \right|$ 

Solution  $\mathbf{X}_m$  is harmonic in time:

$$\mathbf{X}_{m}(\lambda,\varphi,\tilde{t}) = \mathbf{H}_{n}^{k}(\lambda,\varphi)e^{-i\nu_{n}^{k}}$$

With global inner product defined as  $\langle \mathbf{X}_{\mathbf{l}}, \mathbf{X}_{\mathbf{m}} \rangle = \frac{1}{2\pi} \int_{0}^{2\pi} \int_{-1}^{1} \left( \tilde{u}_{l} \tilde{u}_{m}^{*} + \tilde{v}_{l} \tilde{v}_{m}^{*} + \tilde{h}_{l} \tilde{h}_{m}^{*} \right) d\mu d\lambda$ 

functions **H** are expressed as  $\mathbf{H}_{n}^{k}(\lambda, \varphi) = \Theta_{n}^{k} e^{ik\lambda}$  Hough harmonics

The horizontal coefficient vector  $\mathbf{X}_m$  for a given vertical mode is projected onto the Hough harmonics  $\mathbf{H}^n_k(\lambda, \varphi, m)$  as

$$\mathbf{X}_{m}(\lambda,\varphi) = \sum_{n=1}^{R} \sum_{k=-K}^{K} \chi_{n}^{k}(m) \mathbf{H}_{n}^{k}(\lambda,\varphi,m)$$
(3)

The subscript *n* indicates all meridional modes including rotational (ROT), and eastward and westward propagating inertio-gravity (EIG and WIG, respectively) modes

The scalar complex coefficients  $\boldsymbol{\chi}$  are obtained as

$$\chi_{n}^{k}(m) = \frac{1}{2\pi} \int_{0}^{2\pi} \int_{-1}^{1} \left( \tilde{u}_{m}, \tilde{v}_{m}, \tilde{h}_{m} \right)^{T} \left[ \mathbf{H}_{n'}^{k'} \right]^{*} d\mu d\lambda$$
(4)

Here,  $\mu = sin(\phi)$ .

Equations (3-4) are the horizontal transform pair.

Equations (1-2) represent the vertical transform pair.

### Two kinds of Hough functions



Frequencies of spherical normal modes for different equivalent depths

### NMF expansion: horizontal expansion functions

HSFs are pre-computed for a given number of vertical modes, M For every m=1,...,M, i.e. for every D<sub>m</sub>

Meridional structure for Hough functions is computed for a range of the zonal wavenumbers K, k=-K,...,0,...,Kand a range of meridional modes for the balanced, N<sub>ROT</sub>, a range of EIG, N<sub>EIG</sub>, and a range of WIG, N<sub>WIG</sub>, modes.

 $R=N_{ROT} + N_{EIG} + N_{WIG}$ 



### Expansion of 3D data onto NMF: MODES software

#
TOPPATH=/Users/nedjeljka/NMF/Sigma/NMF_MODES
#
INSTALLDIR=\$(TOPPATH)/bin
SHAREDIR=\$(TOPPATH)/share
HOUGHDIR=\$(TOPPATH)/hough
NORMALDIR=\$(TOPPATH)/normal
VSFDIR=\$(TOPPATH)/VSF
MAINDIR=\$(TOPPATH)/main
LIBDIR=/usr/local/lib
# FOR LINUX
SYSTEM = LINUX
LAPACK_LIB=/usr/local/lib
ALFPACK_LIB=/usr/local/lib
GRIB_API_LIB=/usr/local/grib_api-1.9.18/lib
GRIB_API_INCLUDE=/usr/local/grib_api-1.9.18/include
NETCDF_LIB=/opt/local/lib
NETCDF_INCLUDE=/opt/local/include
LIBGRIB_F90=/usr/local/grib_api-1.9.18/lib/libgrib_api_f90.a
LIBGRIB=/usr/local/grib_api-1.9.18/lib/libgrib_api.a

FC = gfortran FFLAGS= -O3 -I\$(SHAREDIR) -I\$(HOUGHDIR) -I\$(NORMALDIR) -I\$ (VSFDIR) -L\$(LIBDIR) -fopenmp -I\$(GRIB\_API\_INCLUDE) LD = \$(FC) LFLAGS = -O3 -fopenmp Makefile and namelist controlled

Required libraries for the input data in grib and netcdf format

Input data on the Gaussian grid and model levels (sigma or hybrid)

Uses LAPACK or equivalent

Preparation steps requires computation of the stability profile

Five executables which are run in subsequent steps:

- Preparation of the horizontal grid
- Computation of the vertical structure functions
- Computation of the horiozntal structure functions
- Projection
- Filtering of selected modes to physical space

### Expansion of 3D data onto NMF: MODES software

#	& vsfcalc_cnf stab_fname = 'stability_L60.data',	
TOPPATH=/Users/nedjeljka/NMF/Sigma/NMF_MODES	vgrid_fname = 'sigma_leveis_L60.data', vsf fname = 'vsf L60.data'.	
<pre># INSTALLDIR=\$(TOPPATH)/bin SHAREDIR=\$(TOPPATH)/share HOUGHDIR=\$(TOPPATH)/hough NORMALDIR=\$(TOPPATH)/normal VSFDIR=\$(TOPPATH)/VSF MAINDIR=\$(TOPPATH)/main LIBDIR=/usr/local/lib #FOR LINUX</pre>	equiheight_fname = 'vsi_Loo.uata', equiheight_fname = 'equivalent_height_L60 num_vmode = 60, mp = 60, hstd = 8000.0d0, suft = 288.0d0, given_stability = .true., ocheck = .true., / & filter_c	).data', cnf
SYSTEM = LINUX LAPACK_LIB=/usr/local/lib	eig.	_n_s = 1, n.e = 70,
ALFPACK_LIB=/usr/local/lib ALFPACK_LIB=/usr/local/grib_api-1.9.18/lib GRIB_API_LIB=/usr/local/grib_api-1.9.18/include NETCDF_LIB=/opt/local/lib NETCDF_INCLUDE=/opt/local/include LIBGRIB_F90=/usr/local/grib_api-1.9.18/lib/libgrib_api_f90.a LIBGRIB=/usr/local/grib_api-1.9.18/lib/libgrib_api.a	&normal_cnf_inverse wig	_n_s = 1, _n_e = 70,
	$nx = 512,$ $rot_{-}$ $ny = 256,$ $km_{-}$	_n_e = 70, ode_s = 100,
	nz = 56, kmc coef3DNMF_fname = 'Houghcoeff_', vmc	ode_e = 85, ode_s = 410, ode_e = 70.
FC = gfortran	inverse_fname = ' Inv_', /	,
FFLAGS= -O3 -I\$(SHAREDIR) -I\$(HOUGHDIR) -I\$(NORMALDIR) -I\$ (VSFDIR) -L\$(LIBDIR) -fopenmp -I\$(GRIB_API_INCLUDE)	inv2hybrid = .false.,	
LD = \$(FC) LFLAGS = -O3 -fopenmp	ps_fname = ' Ps_',	
	meant_fname = 'Tmean_',	
	saveasci = .true.,	
	afname = 'Inverse_ascii_',	
	aformat = '(512F16 4 1x)'	

### NMF software structure: grid and vsf namelists

```
& gaussian
      N = 256.
     gauss_fname = 'gauss256.data',
 /
& vsfcalc cnf
    stab fname = 'stability L60.data',
    vgrid fname = 'sigma levels L60.data',
    vsf fname = 'vsf L60.data',
    equiheight_fname = 'equivalent_height_L60.data',
    num_vmode = 60,
    mp = 60,
    hstd = 8000.0d0,
    suft = 288.0d0.
    given stability = .true.,
   ocheck = .true.,
1
```

### NMF software structure: hsf namelists

```
&houghcalc_cnf
```

1

```
szw = 0,
ezw = 200,
maxl = 70,
my = 256,
freq_fname = 'freq.data',
ocheck = .true.,
```

```
& output
	output_gmt = .false.,
	ofname_gmt = 'hough_gmt',
	ofname_bin = 'hough',
	bin_combine = 'zonal',
```

```
& meridional_grid
ygrid_fname = 'gauss256.data',
/
```

```
& vsf_cnf
equiheight_fname = 'equivalent_height.data',
num_vmode = 43,
/
```

### NMF software structure: projection namelists

	& time
& normal cnf	datetype = 'yyyymmddhh',
- ny - 512	syear = 2007,
	smon = $07$ ,
ny = 256,	sday = 01,
nz = 56,	shour = $12$ ,
nsten = 1	smins = $00$ ,
	ssec = 00,
coef3DNMF_fname = 'Hough_coeff_',	slen = $00$ ,
output_3DNMF = .true.,	eyear $= 2007$ ,
	emon = 07,
Saveps – .uue.,	eday = 01,
savemeant = .true.,	ehour = $12$ ,
ps fname = 'Ps ',	emins = 00,
- $        -$	esec = 00,
meant_mame – mean_,	eien = 00,
saveasci = .true.,	dt = 86400,
afname = 'Indata_',	/ 8 input data
= '(512E20.6.1x)'	& Inpul_uala detefermet_input = 'grib'
$a_{012} = (012 \pm 20.0, 1X),$	$aria = - ECMN/E^{2}$
/	$\bigcup_{i=1}^{n} (i) = \bigcup_{i=1}^{n} (i) = \bigcup_{i=1}^{n$
	zgna_type = nypna

ifile\_grib\_head(1) = 'erai\_N128\_'

```
1
```

### **Energy product**

Starting from the linearized system (u',v',h') and its solutions by using the variable separation, the partition of total energy into the kinetic and available potential energy for every vertical mode is written as :

$$\frac{\partial}{\partial t}\int_{0}^{2\pi}\int_{-1}^{1}\sum_{m}\frac{1}{2}\left(u_{m}^{2}+v_{m}^{2}+\frac{g}{D_{m}}h_{m}^{2}\right)a^{2}d\mu d\lambda=0$$

Global energy product of the m-th vertical mode defined as

$$I_{m} = \frac{1}{2} g D_{m} \sum_{n=1}^{R} \sum_{k=-K}^{K} \chi_{n}^{k}(m) \left[ \chi_{n}^{k}(m) \right]^{*}$$

is equivalent to

$$I_{m} = \frac{1}{2} g D_{m} \int_{0}^{2\pi} \int_{-1}^{1} \left( \tilde{u}_{m}^{2} + \tilde{v}_{m}^{2} + \tilde{h}_{m}^{2} \right) d\lambda d\mu = \int_{0}^{2\pi} \int_{-1}^{1} \left( K_{m} + P_{m} \right) d\lambda d\mu$$

### NMF representation of ERA Interim dataset



### NMF representation of ERA Interim dataset



### Scale-dependent climatological distribution of atmospheric total energy



Percentage of ROT/IG energy in each scale

Based on 30-year period between 1980 and 2009

12 UTC time

Up to 10% of wave energy in IG modes ⇔ 1/3 of circulation

### Application of the normal mode representation

### Modal-space diagnosis of the vertical energy propagation by the Kelvin waves



Evolution of the zonal wavenumber k=1 Kelvin wave in July 2007, filtered for periods shorter than 36 hours. The best agreement between the datasets exists for the Kelvin wave.

### Application of the normal mode representation

### Modal-space diagnosis of the vertical energy propagation by the Kelvin waves



The zonal wind component in the Kelvin wave at subsequent days (left) shows the downward phase propagation (upward energy propagation).

The difference in the depth of the atmosphere in DART/CAM and NCEP/NCAR on one hand and ECMWF and NCEP on the other appears to be one reason for different propagation properties.

(Zagar et al., 2009)

### NMF software structure: filtering namelists

#### &normal\_cnf\_inverse

1

nv –	512	
11X –	512,	eig_
ny =	256,	eig_
nz =	56,	wig_
	AE from - "Lloughoooff"	wig_
COET3DININ	iF_mame = Houghcoem_,	rot
inverse_fr	name ='Inv_',	rot_
inv2hybrid	= .false.,	kmo
ps fname	= ' Ps ',	KMO VMO
		VIIIO
meant_ma	ame = 1 mean_,	, vino
saveasci	= .true.,	1
afname	= 'Inverse_ascii_',	
aformat	= '(512F16.4.1x)'.	
	$\langle \bullet : = = : \bullet : :, : \times \rangle$	

& filter_cnf	
eig_n_s	= 1,
eig_n_e	= 70,
wig_n_s	= 1,
wig_n_e	= 70,
rot_n_s	= 2,
rot_n_e	= 70,
kmode_s	= 100,
kmode_e	= 85,
vmode_s	= 410,
vmode e	= 70.

### NMF representation of ERA Interim dataset



### NMF representation of ERA Interim dataset





# Real-time modal view of ECMWF forecast http://meteo.fmf.uni-lj.si/MODES



### Modal representation of the MJO

Regression between the MJO index and ERA Interim

$$\mathcal{R}_{n}^{k}(m,\tau) = \frac{1}{N-1} \frac{\sum_{t=1}^{N} \left[ \left( \chi_{n}^{k}(m,t) - \overline{\chi_{n}^{k}(m)} \right) \left( \mathcal{M}_{i}(t,\tau) - \overline{\mathcal{M}_{i}} \right) \right]}{Var\left( \overline{\mathcal{M}_{i}} \right)}$$



### Modal representation of MJO

#### Quantification of the role of various modes in MJO and its teleconnections



- Representation of the global circulation and mass fields in global NWP and climate models in terms of the normal modes of the Navier-Stokes equations offers an alternative and physically attractive approach to the diagnostic of some properties of the models
- A new software for the representation of global atmospheric energy, MODES, has been developed under the ERC funding and it is available to the atmospheric research community
- MODES provides possibility to study circulation changes in the coupled climate models in relation to the (un)balanced dynamic. Unbalanced circulation is climatologically small and difficult to diagnose, but it is critical for understanding atmospheric variability

### Thank you very

## much for your

### attention!

# MODESworkshop

On normal-mode function theory and applications to observations, numerical weather prediction models and climate research

Boulder, CO • 26-28 August 2015 • NCAR, Mesa Lab

#### **Topics of interest:**

 Normal-mode function (NMF) theory
 Lessons learnt from the initialization of NWP models based on NMFs
 Diagnosis of global atmospheric dynamics, NWP and climate models using NMFs
 General session on modes of atmospheric variability





Bennert Machenhauer Ron Errico, John Boyd Robert Dickinson Grant Branstator Jose Castanheira Nils Gustafsson Joan Alexander Rolando Garcia Taroh Matsuno Ronald Madden Pedro Silva Dias Wayne Schubert Matthew Wheeler Christian Franzke Andrew Staniforth Daryl Kleist, Dave Parrish George Kiladis, Erland Kallen

Invited speakers:

Information and registration: https://www2.cgd.ucar.edu /sections/amp/events/20150826



### http://meteo.fmf.uni-lj.si/MODES