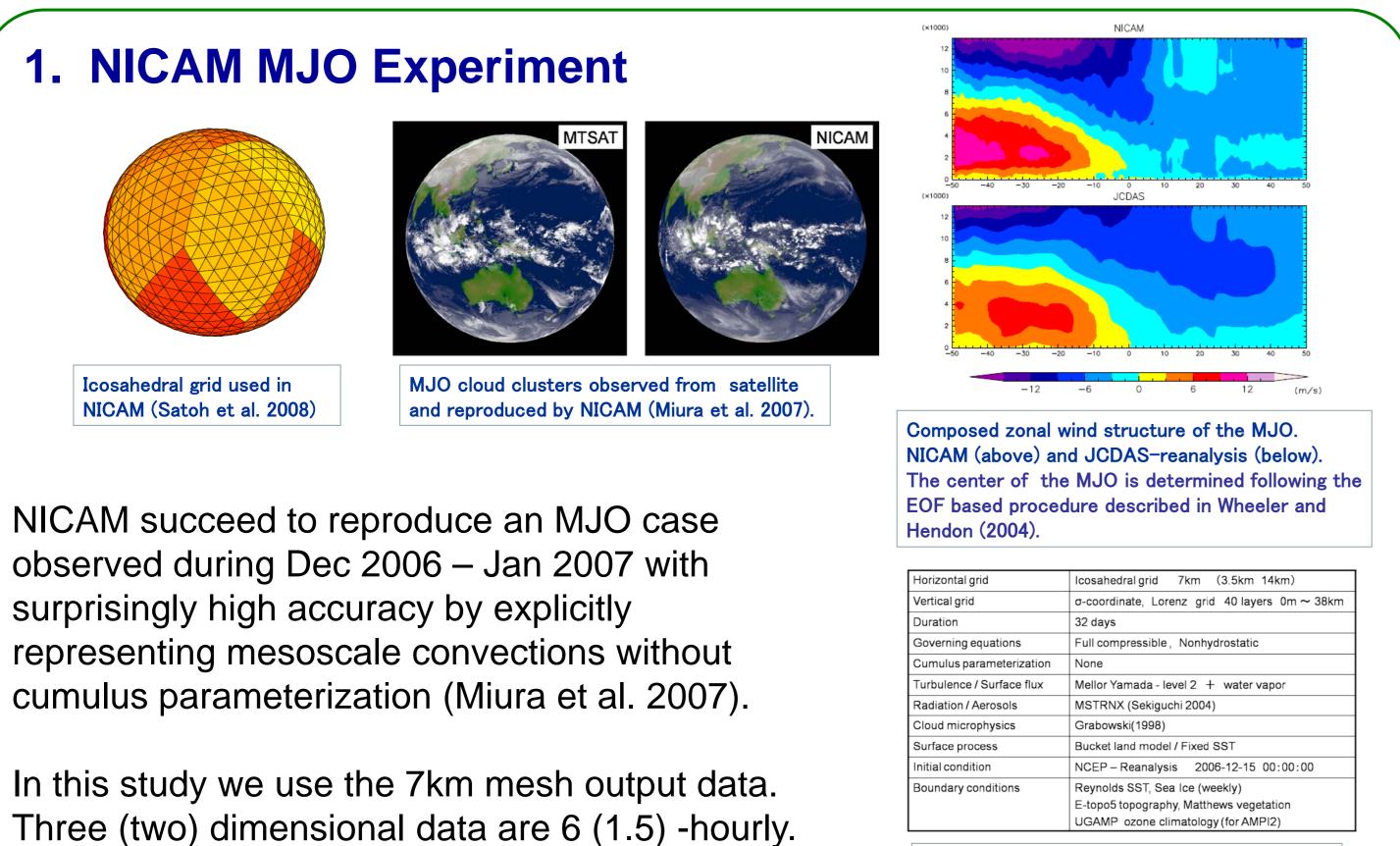
# A study on the Effects of Convective Momentum Transport Associated with Rain Bands within the Madden-Julian Oscillation

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Upscale effects from unresolved Mesoscale Convective Systems (MCSs) are known sources of uncertainty in General Circulation Models (GCMs), which show general difficulty in reproduction of MJOs. The Nonhydrologiccal ICosahedral Atmospheric Model (NICAM) successfully reproduced an MJO case using extremely fine mesh, directly resolving MCSs without cumulus parameterization.

We analyzed the upscale effect of Convective Momentum Transport (CMT) associated with rainbands of MCSs embedded within the

convectively active region of the reproduced MJO case. The upscale zonal acceleration ensemble of CMT formed a three-storied structure: positive near the surface (below 1.6km); negative at low to mid levels (2km - 6.5km); positive at upper levels (above 11km). CMT accounted for -160% of the 2km - 6.5km averaged wind difference that occur associated with the MJO, suggesting that exclusion of CMT effect may result in larger propagation speed of the convective region, possibly up to 10 -15 m/s according to a simplified estimation.



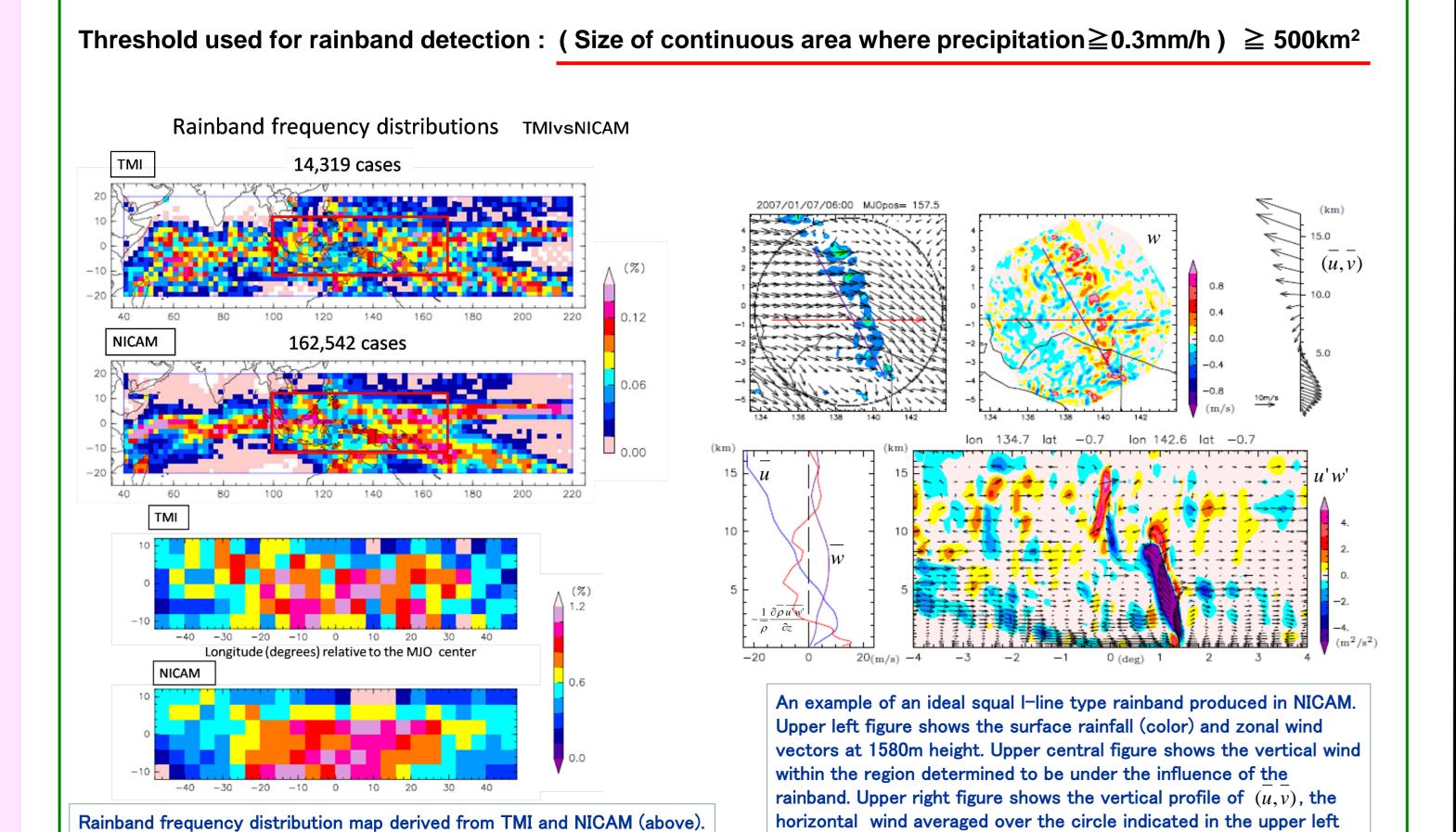
Model configuration for the MJO experiment

### 4. Distribution of CMT Acceleration Vectors Area assumed to be affected from the rainband Examples of various vertical profile purple: (u'w') Contribution from the orange rectangular area to this term (upscale acceleration on u) is considered as the CMT effect of this rainband MJO relative distribution of upscale CMT acceleration vectors defined by the deviation momentum flux divergence shows a three storied structure: positive zonal component near the surface (below 1.6km); negative zonal component at low to mid troposphere (2km-6.5km); positive zonal component at Plan view (above) and zonal-height cross section (below, the upper troposphere (above 11km). colors show u) of CMT vector distribution relative to the

#### 2. Upscale Convective Momentum Transport within the MJO **MJO** convective region propagation speed ≒5m/s Heats the troposphere Rainband Schematic showing the relation between mesoscale convection **Cloud cluster** and grids of typical GCM (left) and NICAM (right). Kelvin response to Hemisphere Schematic of a typical squall-line type MCS. Momentum flux divergence Schematic structure of MJO and embedded convection. Arrows show the (acceleration due to CMT), affected by the horizontal pressure gradient Matsuno-Gill type response due to the cumulus heating ensemble. This study stands in the point of view that eastward propagation of the MJO is across the vertical flow, can have counter gradient effects on the vertical controlled by the eastward shift of the region favorable for convection. wind shear when the convection are organized in to a linear structure.

CMT associated with rainbands embedded in MJOs cause upscale acceleration. This effect is usually parameterized as sub-grid mixing components in GCMs. However, it is known from observation and numerical experiments that organized rainbands can have counter gradient effects on the vertical wind shear (e.g. LeMone and Moncrieff, 1994; Tung and Yanai 2002).

3. Rainbands Detected in the NICAM MJO case



Total number and frequency distribution of rainbands produced in NICAM are consistent with satellite data (TRMM/TMI) considering data retrieval differences. The area within 100°E - 170°E / 12°N – 12°S (red box), where the cases are abundant and the distribution agree well, is chosen as the main analysis region.

The expected ratio of rainband case numbers estimated from the data

retrieval difference is about 10 - 15. Red rectangle indicates the main

spatial distribution of rainbands relative to the MJO center.

analysis region used hereafter in this study. The lower figure shows the

figure. The lower left figure show vertical profiles of :  $\overline{u}$  (blue);  $\overline{w}$ 

(purple); and upscale acceleration by  $-\frac{1}{2}\frac{\partial \rho u'w'}{\partial z}$  (red). The lower right

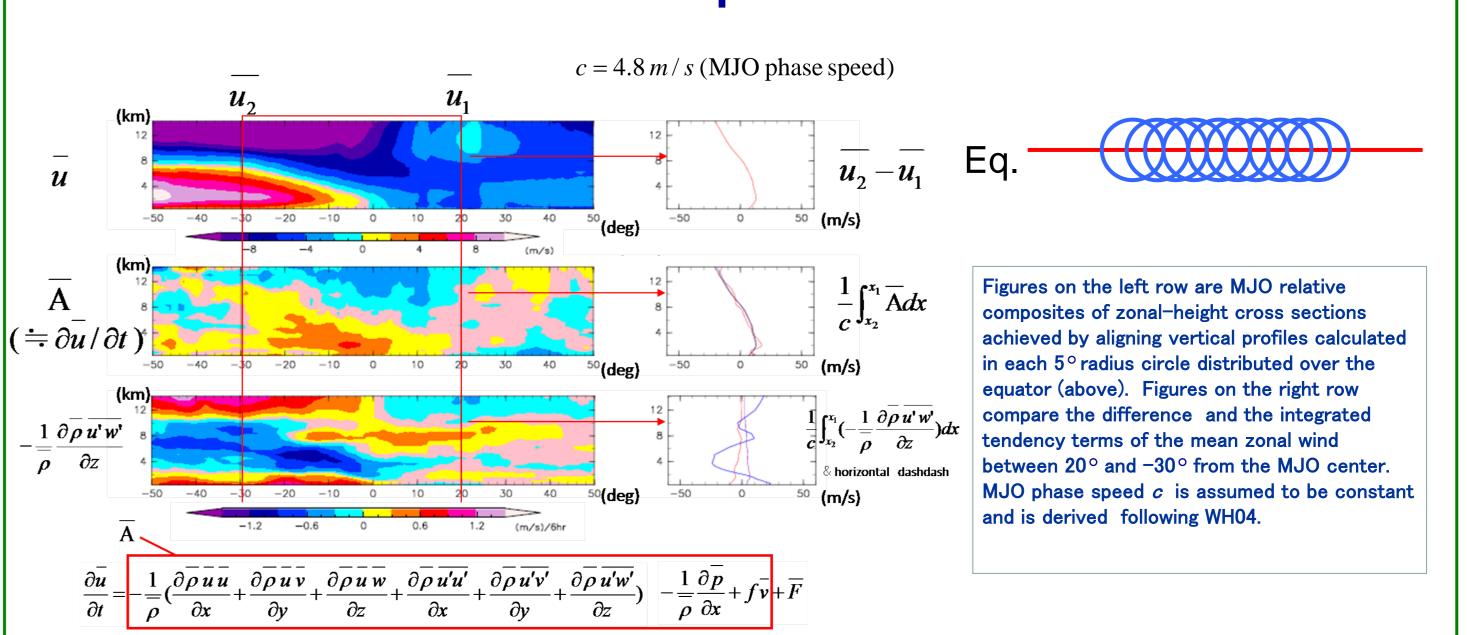
figure is a zonal-height cross section of u'w'. The red arrow in the

upper left figure indicates the zonal axis of the lower right figure.

While some ideal squall line type rainbands are found (upper right figure), there were many cases with complicated structures (not shown).

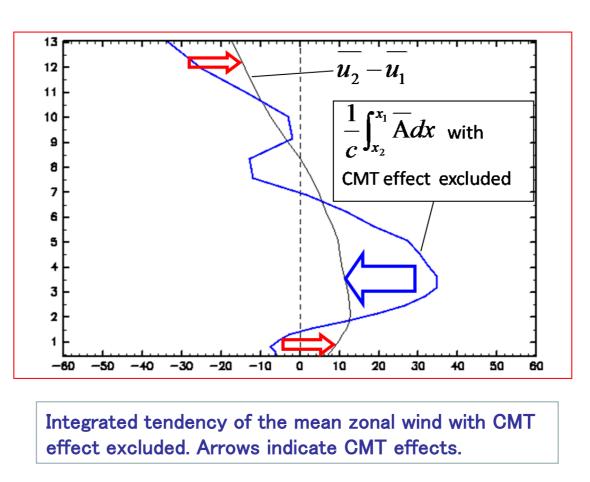
## 5. Quantitative Evaluation of Upscale Acceleration due to CMT

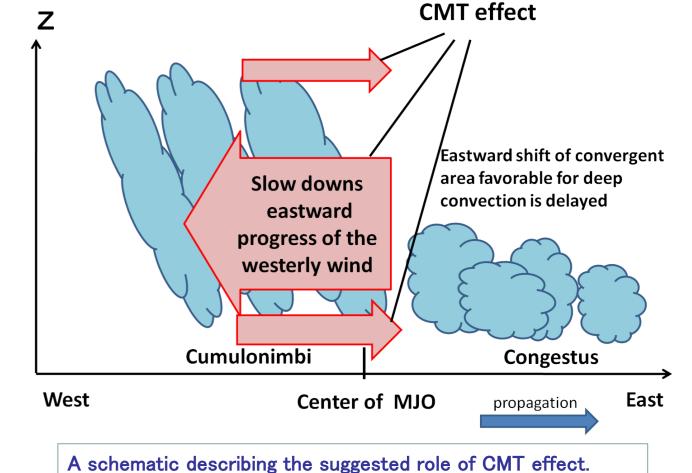
**MJO** center



By distributing the 5° radius circles evenly over the equator, quantitative evaluation of upscale acceleration due to  $-\frac{1}{\rho}\frac{\partial \overline{\rho}\,\overline{u'w'}}{\partial z}$  is performed.  $-\frac{1}{\rho}\frac{\partial \overline{\rho}\,\overline{u'w'}}{\partial z}$  has significant contribution to the wind difference between 20° and -30° from the MJO center. In the low to mid troposphere (2km-6.5km), its contribution on the zonal wind difference was as much as -160% in average.







 $\mathcal{U}$  (m/s)

From the point of view that eastward propagation is controlled by the eastward shift of favorable region for convection, we state further assumptions as follows: (1) development of new convection following the eastward shift of the favorable region and the Matsuno-Gill response occur instantaneously; (2) the MJO wind structure does not change when CMT effects are excluded.

Then, when CMT effects are excluded, the eastward progress of the low to mid level westerly wind speeds up by 260%, thereby speeding up the eastward shift of low to mid level convergent area favorable for deep convection. New convection develops and the MJO wind structure follows instantaneously, and the phase speed of the entire MJO also speeds up by 260%.

Although this estimation is possibly over-simplified, it demonstrates that the impact can be large, and suggests that sufficient representation of mesoscale convections are required for accurate MJO reproduction (In order to do a more realistic estimation, a numerical experiment is required).